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**NUCLEAR CHARGE, CONVECTION CURRENT
AND MAGNETIZATION CURRENT DENSITIES**

by

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Abstract

Nuclear electromagnetic charge and current densities and related topics are discussed. Formulas expressing the multipole-decomposed density distributions in terms of the transition density matrix, geometrical factors and single-particle radial wavefunctions are given in detail. These formulas are used for coding a program, MICRØDENS, written in Fortran-IV, to compute the densities. MICRØDENS also computes the form factors, or the Bessel transforms of the densities. Instructions for the use of the program are provided. Three appendices discuss the densities in connection with several physical processes (gamma-emission, electron scattering and photo nucleon-emission), present the formulas for the computation of the transition density matrix in the most commonly used nuclear models, and give a complete listing of MICRØDENS.

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Densités des charges nucléaires, des courants de convection et des courants de magnétisation

par

H.C. Lee

Résumé

L'auteur de ce rapport fait des commentaires au sujet des densités de courant et de charge électromagnétique nucléaire ainsi qu'au sujet de questions connexes. Il décrit en détail les formules exprimant la répartition des densités à décomposition multipolaire, en fonction de la matrice des densités de transition, de facteurs géométriques et de fonctions ondulatoires radiales à particule unique. Ces formules sont employées pour le codage d'un programme, MICRØDENS, écrit en FORTRAN-IV et destiné à calculer les densités. Par ailleurs, le programme MICRØDENS permet de calculer les facteurs de forme ou les fonctions de Bessel pour les densités. Des instructions sont données pour l'emploi du programme. La première annexe concerne les densités par rapport à plusieurs procédés physiques (émission gamma, diffusion des électrons et émission photo-nucléon), la seconde donne des formules pour calculer la matrice des densités de transition dans les modèles nucléaires les plus couramment utilisés et la troisième est la liste complète des programmes MICRØDENS.

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TABLE OF CONTENTS

1.	INTRODUCTION	1
2.	CHARGE AND CURRENT DENSITY DISTRIBUTIONS	7
2.1	Notation and Phase Conventions	7
2.2	Multipole Decomposition	9
2.3	The Charge Density	10
2.4	The Convection Current Density	11
2.5	The Magnetization Current Density	15
2.6	Parity Selection Rules	18
2.7	Hermitian Property of the Density Operators	19
2.8	Time-Reversal Property of the Density Operators	24
2.9	Isospin Densities	28
3.	THE COMPUTER CODE "MICRØDENS"	29
3.1	Brief Description, Flow Chart	29
3.2	Input Data	31
3.3	Physical Constants	32
3.4	Computation	33
3.5	Output	34
3.6	Subroutines and their Functions	35
	CREDITS	36
	ACKNOWLEDGEMENTS	36
	REFERENCES	37
	APPENDIX A	39
A.1	Gauge Invariance, Charge Conservation and Siegert's Theorem	39
A.2	Gamma-Emission	47
A.3	Electron Scattering	54
A.4	(γ ,N) Reactions	72

APPENDIX B	79
B.1 Transition Density Matrix in the Particle-Hole Model	79
B.2 Transition Density Matrix in the Shell Model	85
B.3 Transition Density Matrix in the Projected Hartree-Fock Approximation	87
APPENDIX C	90
C.1 Listing of MICRØDENS	92
C.2 Sample Output of MICRØDENS	103
C.3 Special Relations for Checking the Results	110

1. INTRODUCTION

The electromagnetic interaction is one of the best tools that can be used to probe the structure of the nucleus. The reasons for this are manifold. Firstly, it is the best understood interaction among those (strong, electromagnetic and weak) that are used to investigate nuclear structure. Secondly, viewed as an operator, the leading and in many cases the only important portion of the interaction is a one-body operator. From a practical point of view the advancement of experimental technologies in the areas of γ -ray detectors, electron accelerators and photon beam generators in the last two decades or so has produced a great wealth of data from which much information in nuclear structure can be extracted.

The nucleus interacts with the electromagnetic field via its charge and current. From a study of the charge and current densities one can infer certain aspects of nuclear structure. In practice often the charge and current densities are computed in a particular model with specific assumptions. These densities are then used in (electromagnetic) reaction calculations and the results are compared with experimental data. Such a comparison may then lead to the rejection, acceptance or modification of the model used, or some of the assumptions made, or both. This report is mainly concerned with the calculation of nuclear charge and current densities.

Suppose that, under the influence of the electromagnetic interaction, the nucleus goes through a transition from state Ψ_i to state Ψ_f (f may be identical to i), then the transition charge, $\rho_{(\vec{r})}^{i \rightarrow f}$, and current, $\vec{J}_{(\vec{r})}^{i \rightarrow f}$, densities are

$$\rho(\vec{r}) = e \int_k \delta(\vec{r} - \vec{r}_k) g_L^k \{\Psi_f^* \Psi_i\} \quad (1)$$

$$\vec{J}(\vec{r}) = \vec{J}^c(\vec{r}) + \vec{J}^m(\vec{r}) + \vec{J}^{exch}(\vec{r}) \quad (2)$$

$$\vec{J}^c(\vec{r}) = \frac{e\hbar}{i2M} \int_k \delta(\vec{r} - \vec{r}_k) g_L^k \{\Psi_f^* \vec{\nabla}_k \Psi_i - \Psi_i \vec{\nabla}_k \Psi_f^*\}, \quad (3)$$

$$\vec{J}^m(\vec{r}) = \frac{e\hbar}{2M} \int_k \delta(\vec{r} - \vec{r}_k) \mu_s^k \vec{\nabla}_k \times \{\Psi_f^* \vec{\sigma}_k \Psi_i\}, \quad (4)$$

where σ_k are the Pauli matrices and we have suppressed the superscripts $i \rightarrow f$. In (1), (3)

and (4), integration over all internal coordinates other than $\vec{r}_k$ is implied for the expression in the curly

brackets; g_L is the nucleon orbital g-factor ($g_L^{\text{proton}} = 1$,

$g_L^{\text{neutron}} = 0$) and μ_s is the magnetic moment of the

free nucleon ($\mu_s^{\text{proton}} = 2.79$, $\mu_s^{\text{neutron}} = -1.91$) in

nuclear magnetons. $\vec{J}^c$ is the convection current due to

the motion of individual nucleons within the nucleus and

$\vec{J}^m$ is the magnetization current generated by the magnetic

moments of the nucleons. $\vec{j}^{\text{exch}}$ is the current that will arise from an exchange or a momentum-dependent nuclear interaction term or both. Although the exchange current has been a subject of interest since the early days of nuclear physics^{1,2)} the nuclear interaction is not sufficiently well known to enable us to treat $\vec{j}^{\text{exch}}$ in a general way. In fact $\vec{j}^{\text{exch}}$ may require a separate treatment for each individual transition, since it is very model dependent. For example a truncation of the shell model basis may be viewed as being effected by a strongly momentum-dependent (in fact, singular) interaction. In this report $\vec{j}^{\text{exch}}$ will be completely ignored. Recently studies of the exchange current can be found in the literature³⁻⁵⁾.

Having decided to consider only the three densities ρ , $\vec{j}^c$ and $\vec{j}^m$ of which the operative forms are well known, we still must decide on a method to treat the nuclear wavefunctions. Currently in the literature there are many models that are used to describe the many-body nuclear wavefunction. Most models have their own distinctive usefulness and together they form a complementary set. It is therefore easy to conclude that for the purpose of this report one should not commit oneself to any one model but rather adopt a more global approach. A natural

linkage between any many-body wavefunction and a one-body density function is the one-body transition density matrix. In this work we shall assume that the density matrix is known, and begin our evaluation of the charge and current densities from this knowledge.

The plan of this report is as follows. In section 2 we first express the density functions in terms of single-particle density functions and the one-body density matrix. We then expand ρ and $\vec{J}$ in terms of scalar and vector spherical multipoles respectively. This expansion is suitable for finite systems such as the nucleus. The multipole single-particle density functions are expressed in terms of geometric coefficients and single-particle radial functions and their derivatives. The hermitian conjugate and the time-reversal properties of the density operators are discussed in two following subsections. A knowledge of these properties may be used to reduce the complexity of the computation. As well, it establishes the realness of the densities (if the radial wavefunctions are real) and enables one to identify a special selection rule, namely that the electric current densities (both longitudinal and transverse) in an elastic transition (i.e., $\Psi \rightarrow \Psi$) vanish. In the last subsection of section 2 we show how the isoscalar and isovector components of the densities are calculated.

In section 3 we describe a package of computer programs, MICRØDENS, written in FORTRAN-IV for the CDC-6600 installation at Chalk River Nuclear Laboratories, for the computation of the nuclear charge and current densities $\rho(\vec{r})$, $\vec{J}^c(\vec{r})$ and $\vec{J}^m(\vec{r})$. The formulas derived in section 2 are used for the coding. The nuclear form factors, which are essentially the spherical Bessel transforms of the densities, are also computed in MICRØDENS. Spherical harmonic oscillator functions (Hermite polynomials) are used as the radial wavefunctions. The advantages of using these functions are their wide use in structure calculations, their well-known analytical properties, and the fact that a complete set of such functions is determined by a single parameter, the oscillator frequency. A flow chart of the program, adequate instructions for the preparation and the assembly of input cards, and brief descriptions of the functions of all subroutines are provided.

A knowledge of the nuclear charge and current densities as isolated entities is academic, since their properties are manifested only through interacting with external fields. In Appendix A, formulas relating the densities to the external field in several physical processes (γ -decay, electron scattering and photo nucleon-emission) are presented. These may be in the form of transition probabilities or strengths or scattering amplitudes or cross sections. The (e, e')

scattering amplitude in the distorted-wave Born approximation is derived in detail. A new formulation for the amplitude, better than those presently available in the literature, is given. The important concepts of gauge invariance and charge conservation and the related Siegert's theorem²⁾ are also briefly discussed at the beginning of the Appendix.

In Appendix B we show how the one-body transition density matrix can be extracted from the result of structural calculations in the three most commonly used nuclear models, namely the particle-hole model and its variants, the shell model, and the angular-momentum-projected Hartree-Fock model.

A complete listing of the computer code MICRØDENS and sample outputs are contained in Appendix C.

2. CHARGE AND CURRENT DENSITY DISTRIBUTIONS

2.1 NOTATION AND PHASE CONVENTIONS

We use the Greek alphabet, except λ , which is reserved for the tensor rank, to represent the complete set of quantum numbers defining a single-particle state; we use the corresponding Roman alphabet for the same set excluding the magnetic quantum number. Thus

$$\alpha = (\tau_a n_a l_a j_a m_a \dots) = (a, m_a); \quad \bar{\alpha} = (a, -m_a).$$

The single-particle wavefunction is l -s coupled

$$\psi_{\alpha}(\vec{r}) = \langle \vec{r} | \alpha \rangle = \sum_{m, \sigma} \langle l_a m \frac{1}{2} \sigma | j_a m_a \rangle i^{l_a} u_a(r) Y_{l_a}^{m_a}(\vec{r}) \chi_{1/2}^{\sigma}. \quad (5)$$

Here $u_a(r)$ is the radial wavefunction. The phase i^{l_a} assures that $|\alpha\rangle$ will have the desired property under time-reversal,

$$T|\alpha\rangle = (-)^{j_a - m_a} |\bar{\alpha}\rangle, \quad (6)$$

when $u_a(r)$ is real. $\chi_{1/2}^{\sigma}$ is the spin wavefunction.

Let O_{λ} be a spherical tensorial operator of rank λ . Using the Wigner-Eckart theorem⁶⁾ we define the reduced matrix element,

$$\langle f \| O_{\lambda} \| i \rangle = \sum_{M_i, \mu} \langle J_i M_i \lambda \mu | J_f M_f \rangle \langle f M_f | O_{\lambda \mu} | i M_i \rangle. \quad (7)$$

We also define a partially reduced matrix element, where only the polar coordinates are integrated over

$$\langle J_f \| O_{\lambda}(r) \| J_i \rangle \equiv \frac{1}{r^2} \langle J_f \| \delta(\vec{r}-\vec{r}') O_{\lambda}(\vec{r}') \| J_i \rangle. \quad (8)$$

In general, if O_{λ} is a one-body operator, the reduced matrix element can be expressed in terms of reduced matrix elements between single-particle states and the one-body transition density matrix, ρ^{if} ,

$$\langle f \| O_{\lambda}(r) \| i \rangle = \sum_{a,b} \frac{\hat{j}_a}{\hat{\lambda}} \rho_{ba,\lambda}^{if} (a \| O_{\lambda}(r) \| b) \quad (9)$$

where $\hat{j}_a = (2j_a + 1)^{1/2}$ etc. and

$$\rho_{ba,\lambda}^{if} \equiv \langle f \| [C_a^{\dagger} \otimes C_b]_{\lambda} \| i \rangle; [C_a^{\dagger} \otimes C_b]_{\lambda \mu} = \sum_{m_a, m_b} (-)^{j_b - m_b} \langle j_a m_a j_b -m_b | \lambda \mu \rangle C_{\alpha}^{\dagger} C_{\beta}. \quad (10)$$

$C^{\dagger}$ and C are respectively the single-particle creation and annihilation operators. We see from (9) that to compute the transition charge or current density it is not necessary to know the wavefunctions of $|i\rangle$ and $|f\rangle$ in their entirety. Only the transition density matrix is needed.

2.2 MULTIPOLE DECOMPOSITION

From now on it will be understood that the transition is from the state $|i\rangle$ to the state $|f\rangle$, and all super- or sub-scripts indicating this fact will be suppressed. Since only certain portions, or multipoles of the density will affect a transition between states of definite angular momentum, it is convenient to decompose the density into various multipoles. We define the multipole charge density ρ_λ and current density $\rho_{\lambda\ell}$ as follows:

$$\rho(\vec{r}) = e \sum_{\lambda\mu} (-i)^\lambda \langle J_i M_i \lambda \mu | J_f M_f \rangle \rho_\lambda(r) Y_\lambda^\mu(\hat{r}) \quad (11)$$

$$\vec{J}^{c,m}(\vec{r})/c = e \sum_{\lambda\ell\mu} (-i)^\ell \langle J_i M_i \lambda \mu | J_f M_f \rangle \rho_{\lambda\ell}^{c,m}(r) \vec{Y}_{\lambda\ell}^{\mu*}(\hat{r}) \quad (12)$$

where

$$\vec{Y}_{\lambda\ell}^{\mu}(\hat{r}) = \sum_{mv} \langle \ell m \ell v | \lambda \mu \rangle Y_\ell^m(\hat{r}) \hat{e}_v \quad (13)$$

is the vector spherical harmonic. $\hat{e}_v$, $v=\pm 1, 0$ are the spherical unit vectors. The quantities ρ_λ , $\rho_{\lambda\ell}^{c,m}$ to be calculated are now functions of the radial variable only.

2.3 THE CHARGE DENSITY

Recall that the total charge density is given as

$$\rho(\vec{r}) = e \sum_k \delta(\vec{r}-\vec{r}_k) g_L^k (\psi_f^* \psi_i). \quad (1)$$

Multiplying the right-hand side of (1) and (11) by $i^\lambda Y_\lambda^\mu(\hat{r})$ and integrating over the polar variables and then using (7) and (8) we get

$$\rho_\lambda(r) = \sum_{a,b} \frac{\hat{j}_a}{\hat{\lambda}} \rho_{ba\lambda} (a \| i^\lambda y_\lambda(\hat{r}) \| b) \quad (14a)$$

with

$$\begin{aligned} (a \| i^\lambda y_\lambda(\hat{r}) \| b) &= -(-)^{j_b + \frac{1}{2} + \lambda} i^{\ell_b + \lambda - \ell_a} \frac{\hat{\lambda} \hat{j}_b}{\sqrt{4\pi}} \begin{pmatrix} j_a & j_b & \lambda \\ \frac{1}{2} & -\frac{1}{2} & 0 \end{pmatrix} u_a^*(r) u_b(n), \\ &\quad \text{if } \ell_a + \ell_b + \lambda \text{ is even;} \\ &= 0, \quad \text{if otherwise.} \end{aligned} \quad (14b)$$

The above equation is general. For the special case when the radial function $u(r)$ is taken to be the oscillator function,

$$u_a(r) = d^{-3/2} R_{n_a \ell_a}(r/d), \quad (15)$$

where $R_{n\ell}(x)$ is the (dimensionless) Hermite polynomial⁷⁾, d is the oscillator length parameter ($d^{-2} = \frac{M\omega}{\hbar} = 0.02412 \hbar\omega$ (F^{-2}), and $\hbar\omega$ is the oscillator frequency in MeV).

2.4 THE CONVECTION CURRENT DENSITY

We repeat the expression for $\vec{J}^c$ given in (3),

$$\vec{J}^c(\vec{r}) = \frac{e\hbar}{2iM} \sum_k \delta(\vec{r}-\vec{r}_k) g_L^k(\psi_f^* \vec{\nabla}_k \psi_i - \psi_i \vec{\nabla}_k \psi_f^*). \quad (3)$$

One must always be especially careful whenever evaluating the matrix element of an operator involving the derivative.

Consider the integral

$$I = \int d\vec{r} \psi_i f(r) \vec{Y}_{\lambda\ell 1}^{\mu}(\hat{r}) \cdot \vec{\nabla} \psi_f^* \quad (16a)$$

where $f(r)$ is a function of r only and $\psi_i \psi_f^* f(r) r^2$ vanishes at $r \rightarrow \infty$. Integrating by parts we get

$$I = - \int d\vec{r} \psi_f^* \vec{\nabla} \cdot f(r) \vec{Y}_{\lambda\ell 1}^{\mu}(\hat{r}) \psi_i. \quad (16b)$$

We left multiply both sides of (3) by $f \vec{Y}_{\lambda\ell 1}^{\mu}$, integrate over all space, and use (9) and (12) to obtain

$$\int r^2 dr f(r) \rho_{\lambda\ell}^c(r) = - \frac{\hbar}{2Mc} \sum_{ab} g_L^a \frac{j_a}{\hat{\lambda}} \rho_{ba\lambda} \langle a \| i^{1+\ell} (f \vec{Y}_{\lambda\ell 1}^{\mu} \cdot \vec{\nabla} + \vec{\nabla} \cdot f \vec{Y}_{\lambda\ell 1}^{\mu}) \| b \rangle. \quad (17)$$

The RHS does not have the desired form since f is being operated upon.

We now note that the Hermitian adjoint of $i^{1+\ell} \vec{\nabla} \cdot f \vec{Y}_{\lambda\ell 1}^{\mu}$ is

$$(i^{1+\ell} \vec{\nabla} \cdot \vec{r} \vec{Y}_{\lambda\ell 1}^{\mu})^{\dagger} = (-)^{1+\lambda+\mu} i^{1+\ell} \vec{r} \vec{Y}_{\lambda\ell 1}^{*\mu} \cdot \vec{\nabla}. \quad (18)$$

Using this and the identity

$$\langle \alpha | O | \beta \rangle \equiv \langle \beta | O^{\dagger} | \alpha \rangle^{*}$$

where O is any operator, we get

$$\begin{aligned} \langle a | i^{1+\ell} \vec{\nabla} \cdot \vec{r} \vec{Y}_{\lambda\ell 1}^{\mu} | b \rangle &= (-)^{1+\lambda+j} b^{-j} a^j \frac{j_b}{j_a} \langle b | i^{1+\ell} \vec{r} \vec{Y}_{\lambda\ell 1}^{*\mu} \cdot \vec{\nabla} | a \rangle^{*} \\ &= (-)^{1+\lambda+j} b^{-j} a^j \frac{j_b}{j_a} \int r^2 dr f(b) i^{1+\ell} \vec{Y}_{\lambda\ell 1} \cdot \vec{\nabla} | a \rangle^{*}. \end{aligned} \quad (19)$$

From (17) and (19), we finally get, independent of r

$$\rho_{\lambda\ell}^c(r) = - \frac{\hbar}{2Mc} \hat{\lambda}^{-1} \sum_{ab} \rho_{ba\lambda} g_L^a \{ j_a (a | i^{1+\ell} \vec{Y}_{\lambda\ell 1} \cdot \vec{\nabla} | b) - (-)^{\lambda+j} b^{-j} a^j j_b (b | i^{1+\ell} \vec{Y}_{\lambda\ell 1} \cdot \vec{\nabla} | a) \}. \quad (20)$$

From (5) the RHS of (19) has a phase factor $i^{\ell_a+1+\ell-\ell_b}$.

Due to the parity of the operator $\vec{Y}_{\lambda\ell 1}^{\mu} \cdot \vec{\nabla}$, which is $(-)^{1+\ell}$,

the non-vanishing matrix elements must have an even

$\ell_a + \ell_b + 1 + \ell$. Therefore the phase factor mentioned above is

real. In other words, $\rho_{\lambda\ell}^c(r)$ is real, if the radial

wavefunctions and $\rho_{ba\lambda}$ are real. We now introduce the

ket $|\ell_a m_a\rangle$ for the spatial part of the single-particle

wavefunction,

$$\langle \vec{r} | \ell_a m_a \rangle \equiv Y_{\ell_a}^{m_a}(\hat{r}) u_a(r); \quad \langle \vec{r} | \ell_a^* m_a^* \rangle = Y_{\ell_a}^{m_a}(\hat{r}) u_a^*(r). \quad (21)$$

The spin wavefunctions can now be contracted out⁶⁾

$$(a \| i^{1+\ell_a} \vec{Y}_{\lambda \ell 1} \cdot \vec{\nabla} \| b) = i^{\ell_b+1+\ell_a-\ell_a} (-)^{j_a+\ell_b+\lambda-\frac{1}{2}} \hat{j}_b \hat{\ell}_a W(\ell_a \ell_b j_a j_b; \lambda \frac{1}{2}) (\ell_a \| \vec{Y}_{\lambda \ell 1} \cdot \vec{\nabla} \| \ell_b). \quad (22)$$

Substituting (22) into (20), we finally get

$$\begin{aligned} \rho_{\lambda \ell}^c(r) = & \frac{\hbar}{Mc} \hat{\lambda}^{-1} \sum_{a,b} i^{\ell_b+1+\ell_a-\ell_a} (-)^{j_a+\frac{1}{2}+\ell_b+\lambda} o_{ba\lambda} g_L^a \hat{j}_a \hat{j}_b W(\ell_a \ell_b j_a j_b; \lambda \frac{1}{2}) \\ & \times \frac{1}{2} \{ (\ell_a \| \vec{Y}_{\lambda \ell 1} \cdot \vec{\nabla} \| \ell_b) - (-)^{\lambda} \ell_b (\ell_b^* \| \vec{Y}_{\lambda \ell 1} \cdot \vec{\nabla} \| \ell_a^*) \}. \end{aligned} \quad (23)$$

The evaluation of the reduced matrix element in (23) is now straightforward; we find

$$\begin{aligned} (\ell_a \| \vec{Y}_{\lambda \ell 1} \cdot \vec{\nabla} \| \ell_b) = & (-)^{\lambda+\ell_b} \frac{\hat{\lambda} \hat{\ell}}{\sqrt{4\pi}} \\ & \sum_{\ell'=\ell_b-1}^{\ell_b+1} (\ell'-\ell_b) \hat{\ell}' \begin{pmatrix} \ell_a & \ell & \ell' \\ 0 & 0 & 0 \end{pmatrix} W(\ell_b 1 \ell_a \ell; \ell' \lambda) u_a^*(r) D_{\ell_b, \ell'-\ell_b}(r) u_b(r) \end{aligned} \quad (24)$$

$$D_{\ell_b, \pm 1}(r) = (\ell_b + \frac{1}{2}(1 \pm 1))^{1/2} \left[\frac{d}{dr} \mp \frac{\ell_b + \frac{1}{2}(1 \mp 1)}{r} \right] \quad (25)$$

For $(\ell_b^* \| \vec{Y}_{\lambda \ell 1} \cdot \vec{\nabla} \| \ell_a^*)$, we interchange ℓ_a with ℓ_b , and $u_a^*(r)$ with $u_b(r)$, in (24).

It is not always necessary to compute the second term on the RHS* of (23). We note that $\vec{Y}_{\lambda\lambda 1}^{\mu}$ is proportional to $\vec{L} Y_{\lambda}^{\mu}$; it therefore commutes with the gradient operator, $\vec{\nabla}$. Consequently for the current that affects magnetic transitions, $\rho_{\lambda\lambda}^c(r)$, the second term is equal to the first term.

Equations (23) and (24) are applicable for any radial wavefunction. When oscillator functions $R_{n\ell}(x)$ are employed, the relations given below are useful. For $\ell > 0$, we have

$$\left(\frac{d}{dx} + \frac{\ell+1}{x}\right)R_{n\ell}(x) = (\ell+n+\frac{1}{2})^{\frac{1}{2}} R_{n\ell-1}(x) + (n+1)^{\frac{1}{2}} R_{n+1,\ell-1}, \quad (26a)$$

$$\left(\frac{d}{dx} - \frac{\ell}{x}\right)R_{n\ell}(x) = \left(\frac{d}{dx} + \frac{\ell+1}{x}\right)R_{n\ell}(x) - \frac{2\ell+1}{x} R_{n\ell}. \quad (26b)$$

When $\ell = 0$, we have

$$\frac{d}{dx} R_{n0}(x) = -(n+\frac{3}{2})^{\frac{1}{2}} R_{n1}(x) - n^{\frac{1}{2}} R_{n-1,1}(x). \quad (26c)$$

In (26), n is the number of nodes of the oscillator function in the range $0 < r < \infty$.

*RHS - right-hand side

2.5 THE MAGNETIZATION CURRENT DENSITY

We repeat equation (4)

$$\vec{j}^m(\vec{r}) = \frac{e\hbar}{2M} \sum_j \delta(\vec{r}-\vec{r}_j) \mu_s^j \vec{\nabla}_j \times (\psi_f^* \vec{\sigma}_j \psi_i), \quad (4)$$

and consider the integral

$$\begin{aligned} I &= i^\ell \int d\vec{r} f(r) \vec{Y}_{\lambda\ell 1}^\mu(r) \cdot [\vec{\nabla} \times (\psi_f^* \vec{\sigma} \psi_i)] \\ &= i^\ell \int d\vec{r} \psi_f^* \vec{\sigma} \psi_i \cdot \vec{\nabla} \times (f \vec{Y}_{\lambda\ell 1}^\mu). \end{aligned} \quad (27)$$

Using the well-known relations

$$(\vec{\nabla} \times f \vec{Y}_{\lambda\ell 1}^\mu)_q = \frac{\sqrt{2}}{i} \sum_{mv} (-)^k \langle \ell m 1 - k | \lambda \mu \rangle \langle 1 v 1 k | 1 q \rangle \nabla_v (f Y_\ell^m) \quad (28)$$

and

$$\begin{aligned} \nabla_v f Y_\ell^m &= \left(\frac{\ell+1}{2\ell+3}\right)^{1/2} \langle \ell m | V | \ell+1, m+v \rangle Y_{\ell+1}^{m+v} \left(\frac{d}{dr} - \frac{\ell}{r}\right) f \\ &\quad - \left(\frac{\ell}{2\ell-1}\right)^{1/2} \langle \ell m | V | \ell-1, m+v \rangle Y_{\ell-1}^{m+v} \left(\frac{d}{dr} + \frac{\ell+1}{r}\right) f \end{aligned} \quad (29)$$

and calling the left-hand side of (28) $D_q(r, \frac{d}{dr})f$, we rewrite (27) as

$$\begin{aligned}
 I &= i^{\ell+1} \sum_q (-)^q \int d\vec{r} (\Psi_f^* \sigma_{-q} \Psi_i) D_q(r, \frac{d}{dr}) f \\
 &= -i^{\ell+1} \sum_q (-)^q \int d\vec{r} f D_q(r, -\frac{d}{dr}) r^2 \Psi_f^* \sigma_{-q} \Psi_i,
 \end{aligned} \tag{27'}$$

upon integration by parts. Equation (27') has the desired form, since f is not operated upon. From (27'), (28), (7), (9) and (12), and using standard angular momentum recoupling techniques, we have

$$\begin{aligned}
 c_{\lambda\ell}^m(r) &= -\frac{\hbar}{2Mc} \hat{\lambda}^{-1} \sum_{a,b} \hat{j}_a \rho_{ba\lambda} \mu_s^a i^{1+\ell} \sqrt{6} \\
 &\times \{ (\ell+1) W(11\ell+1\ell; 1\lambda) (\frac{d}{dr} + \frac{\ell+2}{r}) (a \| T_\lambda(Y_{\ell+1} \otimes \sigma) \| b) \\
 &- \ell^{\frac{1}{2}} W(11\ell-1, \ell; 1\lambda) (\frac{d}{dr} - \frac{\ell-1}{r}) (a \| T_\lambda(Y_{\ell-1} \otimes \sigma) \| b) \}.
 \end{aligned} \tag{30}$$

$T_\lambda(Y_\ell \otimes \sigma)$ is a rank- λ tensor product of Y_ℓ and $\vec{\sigma}$:

$$T_{\lambda\mu}(Y_\ell \otimes \sigma) = \sum_{mq} \langle \ell m 1 q | \lambda \mu \rangle Y_\ell^m \sigma_q.$$

With (5), we have

$$\begin{aligned}
 &(a \| T_\lambda(Y_\ell \otimes \sigma) \| b) \\
 &= (-)^{\ell_a} i^{\ell_b - \ell_a} \left(\frac{3}{2\pi} \right)^{\frac{1}{2}} \hat{\lambda} \hat{\ell} \hat{j}_b \hat{\ell}_a \hat{\ell}_b \begin{pmatrix} \ell_a & \ell & \ell_b \\ 0 & 0 & 0 \end{pmatrix} \begin{pmatrix} j_a & j_b & \lambda \\ \ell_a & \ell_b & \ell \\ \frac{1}{2} & \frac{1}{2} & 1 \end{pmatrix} u_a^*(r) u_b(r).
 \end{aligned} \tag{31}$$

From (30) and (31), $\ell_a + \ell_b + \ell + 1$ must be even, i.e., the phase factor $i^{\ell_a + \ell_b - \ell}$ is real, and $\rho_{\lambda\ell}^m$ is real, if u_a and u_b and $\rho_{ba\lambda}$ are real. Since $\ell = \lambda, \lambda \pm 1$, the two equations above may be further simplified by evaluating the Racah and 9-j coefficients. We finally get

$$\rho_{\lambda\ell}^m(r) = -\frac{\hbar}{2Mc} \sum_{ab} i^{\ell_b - \ell_a + 1 + \ell} (-)^{\lambda + j_b + \frac{1}{2}} \rho_{ba\lambda} \mu_s \frac{\hat{j}_a \hat{j}_b}{\hat{\lambda} \sqrt{4\pi}} \begin{pmatrix} j_a & j_b & \lambda \\ \frac{1}{2} & -\frac{1}{2} & 0 \end{pmatrix} F_{\lambda\ell}^{ab}(r) u_a^*(r) u_b(r) \quad (32)$$

$$F_{\lambda\lambda}^{ab}(r) = -\left(\frac{2\lambda+1}{\lambda(\lambda+1)}\right)^{\frac{1}{2}} \left[\frac{\lambda(\lambda+1)}{r} + (x_a + x_b) \left(\frac{d}{dr} + \frac{1}{r} \right) \right] \quad (33a)$$

$$F_{\lambda, \lambda \pm 1}^{ab}(r) = \frac{1}{\lambda + \frac{1}{2} \pm \frac{1}{2}} (x_a - x_b) D_{\lambda, \pm 1}(r) \quad (33b)$$

with $D_{\lambda, \lambda \pm 1}(r)$ given by (25), and $x_a = (\ell_a - j_a)(2j_a + 1)$.

2.6 PARITY SELECTION RULES

We recall that for transitions affected by the electric or Coulomb field, the parity change from $|i\rangle$ to $|f\rangle$ is "natural" $((-)^{L_i-L_f} = (-)^\lambda)$, and for magnetic transitions it is "unnatural" $((-)^{L_i-L_f} = (-)^{\lambda+1})$. These selection rules must be reflected in the transition densities. In table 1 these rules are summarized.

TABLE 1. PARITY SELECTION RULES

density	type	$L_i+L_f+\lambda$	$\ell_a+\ell_b+\lambda$	$\ell_a+\ell_b+\ell+1$
$\rho_\lambda(r)$	Coulomb	even	even	—
$\rho_{\lambda,\ell=\lambda}^{C,m}(r)$	Magnetic	odd	odd	even
$\rho_{\lambda,\ell=\lambda\pm 1}^{C,m}(r)$	Electric	even	even	even

2.7 HERMITIAN PROPERTY OF THE DENSITY OPERATORS

When the Hermitian adjoint of a tensor operator O_{λ}^{μ} is

$$(O_{\lambda}^{\mu})^{\dagger} = (-)^{\lambda+\mu} O_{\lambda}^{-\mu},$$

we shall call O_{λ}^{μ} Hermitian. Conversely, if

$$(O_{\lambda}^{\mu}) = -(-)^{\lambda+\mu} O_{\lambda}^{-\mu}$$

is true, we call O_{λ}^{μ} anti-Hermitian.

The Hermitian property of an operator can be exploited to shorten the calculation of its matrix elements. We define a multipole density operator to be an operator such that its reduced matrix elements between $|\Psi_i\rangle$ and $|\Psi_f\rangle$ give the correct density. From (1), the charge density operator is

$$\rho_{op, \lambda \mu}(\vec{r}) = \frac{1}{r^2} \sum_j g_L^j \delta(\vec{r} - \vec{r}_j) i^{\lambda} Y_{\lambda}^{\mu}(\hat{r}_j). \quad (34)$$

From (3) and (17), the convection current density operator is

$$\rho_{op, \lambda \mu, \ell}^c(\vec{r}) = - \frac{\hbar}{2Mc r^2} i^{1+\ell} \sum_j g_L^j \{ \delta(\vec{r} - \vec{r}_j) \vec{Y}_{\lambda \ell 1}^{\mu}(\hat{r}_j) \cdot \vec{\nabla}_j + \vec{\nabla}_j \cdot \delta(\vec{r} - \vec{r}_j) \vec{Y}_{\lambda \ell 1}^{\mu}(\hat{r}_j) \}. \quad (35)$$

We cannot immediately write down the magnetization density operator from (4). However, by partial integration, we find

$$\rho_{op, \lambda \mu, \ell}^m(\vec{r}) = \frac{\hbar}{2Mc} i^\ell \sum_j \mu_s^j \{ \delta(\vec{r} - \vec{r}_j) \vec{Y}_{\lambda \ell 1}^{\mu}(\hat{r}_j) \cdot (\vec{\nabla}_j \times \vec{\sigma}_j) + (\vec{\sigma}_j \times \vec{\nabla}_j) \cdot \delta(\vec{r} - \vec{r}_j) \vec{Y}_{\lambda \ell 1}^{\mu}(\hat{r}_j) \} \quad (36)$$

Recalling that

$$\vec{Y}_{\lambda \ell 1}^{\mu} \cdot \vec{A} = \sum_{m, q} \langle \ell m 1 q | \lambda \mu \rangle Y_{\ell}^m A_q$$

and

$$\vec{Y}_{\lambda \ell 1}^{\mu} \cdot (\vec{A} \times \vec{B}) = \frac{\sqrt{2}}{i} \sum_{mrs} \langle \ell m 1 q | \lambda \mu \rangle \langle 1 r 1 s | 1 q \rangle Y_{\ell}^m A_r B_s,$$

where $\vec{A}$ and $\vec{B}$ are any rank-1 tensors, and $(i^\ell Y_{\ell}^m)^{\dagger} = (-)^{\ell-m} i^\ell Y_{\ell}^{-m}$; $(\nabla_q)^{\dagger} = (-)^{1-q} \nabla_q$ and $(\sigma_q)^{\dagger} = (-)^q \sigma_{-q}$, it is easily shown that

$$(\rho_{op, \lambda \mu})^{\dagger} = (-)^{\lambda+\mu} \rho_{op, \lambda -\mu}, \quad (37a)$$

and

$$(\rho_{op, \lambda \mu, \ell}^{c, m})^{\dagger} = -(-)^{\lambda+\mu} \rho_{op, \lambda -\mu, \ell}^{c, m}. \quad (37b)$$

That is, the charge density operator is Hermitian and the current density operators are anti-Hermitian. Note that in (35) as well as in (36), either one of the two terms on the RHS by itself does not have definite Hermiticity.

Now consider a tensor O_λ^μ which has definite Hermiticity, $(O_\lambda^\mu)^\dagger = (-)^{p+\mu} O_\lambda^{-\mu}$; we find

$$\begin{aligned} \langle JM | O_\lambda^\mu | J'M' \rangle &= \langle J'M' | (O_\lambda^\mu)^\dagger | JM \rangle^* \\ &= (-)^{p-\mu} \langle J'M' | O_\lambda^{-\mu} | JM \rangle^* \end{aligned}$$

leading to

$$\langle J | O_\lambda | J \rangle = (-)^{p+J-J'} \hat{J}' \langle J' | O_\lambda | J \rangle^*. \quad (38)$$

In the previous sections, we have shown that, when $\rho_{ba\lambda}$ and the radial functions are real[†], reduced single-particle matrix elements of $\rho_{op,\lambda\mu}$ and $\rho_{op,\lambda\mu,\ell}^{c,m}$ are all real. In this case, from (37,38) we can make the replacement on the LHS of (15), for the charge density

$$\sum_{ab} \rho_{ba,\lambda} \rightarrow \sum_{a \geq b} \frac{1}{1+\delta_{ab}} [\rho_{ba,\lambda} + (-)^{\lambda+j_b-j_a} \rho_{ab,\lambda}]; \text{ for } \rho_\lambda(r). \quad (39)$$

Similarly, for the current densities, on the LHS of (23) and (32),

$$\sum_{a,b} \rho_{ba,\lambda} \rightarrow \sum_{a \geq b} \frac{1}{1+\delta_{ab}} [\rho_{ba,\lambda} - (-)^{\lambda+j_b-j_a} \rho_{ab,\lambda}]; \text{ for } \rho_{\lambda\ell}^{c,m}(r). \quad (40)$$

[†] In the rest of this section, we shall assume this to be the case, unless otherwise stated.

Equation (40) leads to an interesting result for electric transitions. Consider the pair $a=b$. Due to the parity selection rule for electric transitions, λ must be even. Therefore we see that the pair $a=b$ cannot generate a convection current which contributes to $E\lambda$ transitions. A special case of this property is when $|\Psi_i\rangle = |\Psi_f\rangle$. For this case it follows immediately from (3) that $\vec{j}^c(\vec{r}) = 0$.

Since the gradient operator is spin independent, we can even make a stronger statement; no convection current can be generated if the spatial wavefunctions of $|i\rangle$ and $|f\rangle$ are identical. Microscopically, this is manifestly true from (23). In the spherical harmonic oscillator model, the spatial wavefunctions of the two states with $j = 2 \pm \frac{1}{2}$ are identical: therefore there is no convection current between such two states. In more realistic (single-particle) models their wavefunctions will in general be different. The difference will be small, however, because it is generated only by the spin-orbit potential, which is a small part of the total potential. This observation, coupled to the fact that only currents can affect γ -transitions (due to gauge invariance, or equivalently, the conservation of charge), seems to lead to the following statement: $E\lambda$ transitions between spin-orbit partner states, such as $d_{3/2} \rightarrow d_{5/2}$,

are always weak. On the other hand, we know this conjecture is not supported by experiments. The answer to this apparent paradox lies in the fact that only convection current has been considered so far. Currents generated by exchange and momentum-dependent forces or due to basis truncations (i.e. J^{exch}) have been ignored. The fact that there are strong E2 transitions in nuclei across the whole periodic table is proof that J^{exch} is not negligible. Yet for reasons stated in the Introduction, a global treatment of J^{exch} is not feasible. This difficulty is resolved, at least in part, for low energy processes (wave length of photon $\geq$ nucleon size), by the so-called Siegert's theorem²⁾. This theorem enables us to side-step the question of currents, and relate the electric transition directly to the nuclear charge distribution. A comprehensive discussion of Siegert's theorem is somewhat outside the scope of this report; however, it is discussed briefly at the beginning of Appendix A.

2.8 TIME-REVERSAL PROPERTY OF THE DENSITY OPERATORS

We shall use the time-reversal properties of the density operators to establish the realness of the densities. Under time-reversal (TR), all momenta and spins change sign, and all c-numbers become their respective complex-conjugates. Let us call the TR transformation T, and for any state $|\psi\rangle$

$$|\psi_T\rangle \equiv T|\psi\rangle. \quad (41)$$

We shall call a state of good angular momentum TR-invariant if under the transformation T, other than gaining a phase $(-)^{J-M}$ only its magnetic quantum number changes sign,

$$T|\gamma JM\rangle = (-)^{J-M}|\gamma J, -M\rangle, \quad (42)$$

where γ is the set of all other quantum numbers. Thus as stated in (6) the single-particle wavefunction given in (5) is TR-invariant, if $u(r)$ is real.

Let O be any operator, and

$$O|\psi\rangle = \sum_n |\phi_n\rangle \langle \phi_n | O | \psi \rangle. \quad (43)$$

We want to find the property of O under TR, which is defined as

$$TO|\psi\rangle = TOT^{-1}T|\psi\rangle \equiv O_T|\psi_T\rangle$$

(44)

$$= \sum_n |\phi_{nT}\rangle \langle \phi_n | O | \psi \rangle^*,$$

the second line comes from (43), taking into account that $\langle \phi_n | O | \psi \rangle$ is c-number. It follows that

$$\langle \phi_T | O_T | \psi_T \rangle = \langle \phi | O | \psi \rangle^*. \quad (45)$$

Suppose

$$(O_\lambda^\mu)_T = (-)^{p-\mu} O_\lambda^{-\mu},$$

and let $|JM\rangle$ and $|J'M'\rangle$ be TR-invariant, then from (45)

$$(-)^{J-M-J'+M'+p-\mu} \langle J-M | O_\lambda^{-\mu} | J'-M' \rangle = \langle JM | O_\lambda^\mu | J'M' \rangle^*.$$

Taking the reduced matrix elements on both sides, we have

$$\langle J || O_\lambda || J' \rangle^* = (-)^{p-\lambda} \langle J || O_\lambda || J' \rangle. \quad (46)$$

Therefore $\langle J || O_\lambda || J' \rangle$ is real, if $p=\lambda$, and is purely imaginary if $p = \lambda+1$.

From (34), (35) and (36) we find that the charge and current density operators transform similarly under T

$$(\rho_{op,\lambda\mu})_T = (-)^{\lambda-\mu} \rho_{op,\lambda-\mu}, \quad (47a)$$

$$(\rho_{op,\lambda\mu,\ell}^{c,m})_T = (-)^{\lambda-\mu} \rho_{op,\lambda-\mu,\ell}^{c,m}. \quad (47b)$$

Since we have chosen a phase convention such that the single-particle wavefunctions are TR-invariant (Equation (5)) it follows that all single-particle reduced matrix elements of the density operators are real. This has already been pointed out in Section 2. To establish the realness of the density itself, it remains to show that the many-body states $|\Psi_i\rangle$ and $|\Psi_f\rangle$ are TR-invariant. This can easily be shown by induction, if (5) is true, and if all the coupling coefficients are real. Here we only show it to be true for a two-particle state. Let

$$|ab; JM\rangle = |\alpha\beta; JM\rangle = \sum_{m_a, m_b} \langle j_a m_a j_b m_b | JM \rangle |\alpha\rangle |\beta\rangle.$$

Then from (5),

$$\begin{aligned} T|ab; JM\rangle &= \sum_{m_a, m_b} \langle j_a m_a j_b m_b | JM \rangle (-)^{j_a - m_a} (-)^{j_b - m_b} |\bar{\alpha}\rangle |\bar{\beta}\rangle \\ &= (-)^{J-M} \sum_{m_a, m_b} \langle j_a -m_a j_b -m_b | J-M \rangle |\bar{\alpha}\rangle |\bar{\beta}\rangle \\ &= (-)^{J-M} |ab; J-M\rangle. \end{aligned}$$

Therefore $|JM\rangle$ is TR-invariant. We have thus shown that for the phase convention used here, the charge and current densities are real. A corollary is that the one-body density matrix element $\rho_{ba\lambda}$ is real.

It should be emphasized that Equations (5), (37) and (47) follow strictly from the phase convention adopted in Section 2.1, which is essentially replacing the spherical harmonic Y_{ℓ}^m by $i^{\ell}Y_{\ell}^m$ whenever the former appears. Any other phase-convention, if consistently used, will be equally acceptable but may result in imaginary or complex densities.

The realness of the current, together with the Hermitian property of the current operator, lead to a special selection rule when $|\Psi_f\rangle$ is equal to $|\Psi_i\rangle$. From the identity

$$\langle \Psi_f | O_{\lambda\mu} | \Psi_i \rangle^* = \langle \Psi_i | (O_{\lambda\mu})^{\dagger} | \Psi_f \rangle$$

and (37b), we have

$$(\rho_{\lambda\ell}^{i\rightarrow i}(r))^* = \rho_{\lambda\ell}^{i\rightarrow i}(r) = -(-)^{\lambda} \rho_{\lambda\ell}^{i\rightarrow i}(r).$$

For magnetic transitions ($\ell=\lambda$), $(-)^{\lambda}$ must be odd, from the parity selection rule, so $-(-)^{\lambda} = 1$. Therefore a magnetic transition is allowed. A similar argument shows that the diagonal electric transition is not allowed.

2.9 ISOSPIN DENSITIES

So far we have discussed the charge and current distributions in terms of proton and neutron densities. If the proton and neutron wavefunctions are close approximates to each other, a condition which is certainly fulfilled in all light and intermediate nuclei, then it is convenient to discuss the distributions in terms of isoscalar and isovector densities. Let us call the isoscalar density matrix $\rho_{ba\lambda}^{(0)}$ and the isovector density matrix $\rho_{ba\lambda}^{(1)}$. It is easy to see that

$$\rho_{ba}^{(0)} = \frac{1}{\sqrt{2}} (\rho_{ba\lambda}^{\text{proton}} + \rho_{ba\lambda}^{\text{neutron}}), \quad (48a)$$

$$\rho_{ba}^{(1)} = \frac{1}{\sqrt{2}} (\rho_{ba\lambda}^{\text{proton}} - \rho_{ba\lambda}^{\text{neutron}}). \quad (48b)$$

We can therefore pretend that there is only one kind of particle with orbital g-factor and magnetic moment which is isospin dependent. Thus

$$g_L^{(T)} = \frac{1}{\sqrt{2}} (g_L^{\text{proton}} + (-)^T g_L^{\text{neutron}}), \quad (49)$$

$$\mu_S^{(T)} = \frac{1}{\sqrt{2}} (\mu_S^{\text{proton}} + (-)^T \mu_S^{\text{neutron}}); \quad T = 0, 1. \quad (50)$$

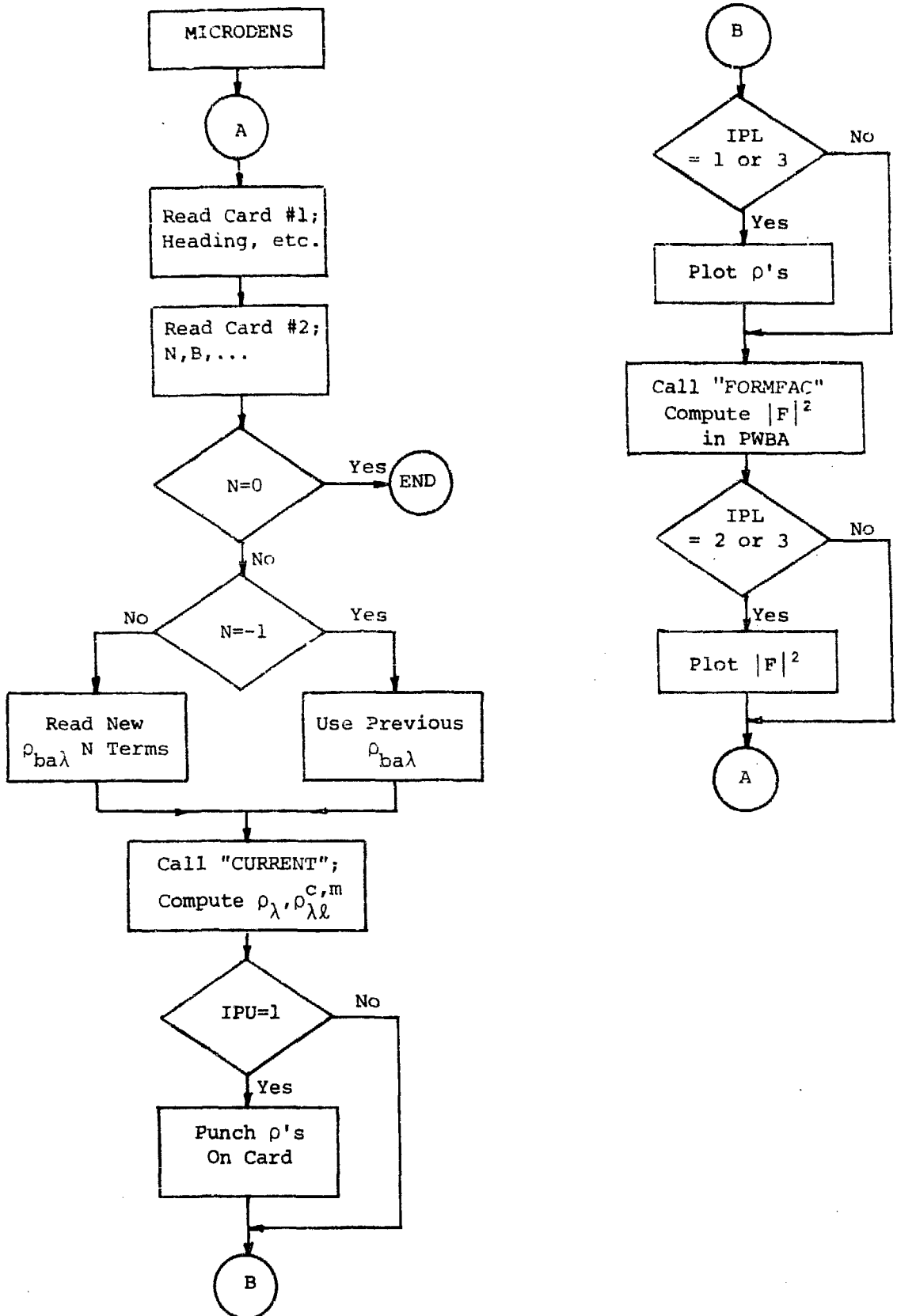
3. THE COMPUTER CODE "MICRØDENS"

3.1 BRIEF DESCRIPTION, FLOW CHART

MICRØDENS is a self-contained package of computer programs, written in the Fortran-IV language, for calculating the charge and current densities, given the one-body transition density matrix. The execution of the program is controlled by input data cards, and the results are printed and/or punched on cards and plotted. The flow chart of the code is given in Fig. 1. The execution begins at point A. Upon the completion of a full cycle of computation, the command returns to point A. The job is terminated upon reading of #2 data card with $N=0$.

A second part of the program, shown as part B in the flow chart, calculates the Coulomb, electric and magnetic form factors (see Appendix A, Equations (27) and (27c')) corresponding to the one-body transition density matrix.

FIGURE 1. FLOW CHART OF MICRODENS



3.2 INPUT DATA

For each transition for which the densities are to be calculated, there are three sets of data cards. Each of the first two sets contains only one card. The number of cards in the third set varies.

The contents and format of each card is as follows:

Content/Format

card - NAME, KIND
no. 1 7A10, A10

card - N, B, DX, NX, DQ, QMAX, VJI, VJF, Z, EFC, EFM, JS, JL, IPL, IPU
no. 2 I4, F6.3, F5.3, I5, (7F5.1), (4I2)

card - NLJA(1), NLJB(1), IRO(1), RØ(1), NLJA(2), ...
no. 3 I3, I3, I1, F8.5, I3,

NAME is the identifier of the particular calculation. It may be up to seven words (10 characters per word) long, and is used as the heading for all output. KIND may be either one of the three left-adjusted words COULOMB, (ρ_{λ} computed), ELECTRIC ($\rho_{\lambda, \lambda \pm 1}^{c, m}$ computed) and MAGNETIC ($\rho_{\lambda \lambda}^{c, m}$ computed). N is the number of $\rho_{ba\lambda}$ to be read. It therefore determines the number $(N/5+1)$ of cards in set 3. B is the oscillator length parameter (Section 2.3) in Fermies. The meanings of other entries on card No. 2

are clearly explained in the listing of MICRØDENS (see Appendix C), and will not be repeated here. For each pair (a,b), NLJA, NLJB, IRØ and RØ have respectively the following meanings:

$$NLJA = 32*n_a + 2*\ell_a + j_a + \frac{1}{2} - \ell_a,$$

with NLJB similarly defined; IRØ=1(0) if (a,b) is a proton (neutron) pair; and RØ = $\rho_{ba\lambda}$. $n_a \geq 0$ is the principal quantum number minus one.

3.3 PHYSICAL CONSTANTS

We give the names and values of the physical constants used in MICRØDENS:

HBC = $\hbar c = 197.33$ (MeV.F);

BOHR = $\mu_n/e = \hbar/2Mc = 0.105(F)$, μ_n is the nuclear magneton;

GLP = $g_L^p = 1$, orbital g-factor for proton;

GLN = $g_L^n = 0$, orbital g-factor for neutron;

MUSP = $\mu_s^p = 2.79$, magnetic moment for protons in μ_n ;

MUSN = $\mu_s^n = -1.91$, magnetic moment for neutron in μ_n .

The effective charge for proton (neutron), in units of e, is GLP+EFCH (GLN+EFCH). EFCH is read as an input parameter on card No. 2.

3.4 COMPUTATION

The charge and current densities are calculated in the subroutine CURRENT. When KIND = COULOMB, the dimensionless density $\tilde{\rho}_\lambda(x) = d^3 \rho_\lambda(r) \big|_{r=xd}$ is computed using (14). When KIND = ELECTRIC (or MAGNETIC), $\rho_{\lambda, \lambda \pm 1}^{c,m}(x) = d^3 \rho_{\lambda, \lambda \pm 1}(r) \big|_{r=xd}$ (or $\tilde{\rho}_{\lambda\lambda}^{c,m}(x)$) are calculated using (23-26) and (30-33); d is the oscillator length parameter (see Section 2.3).

In the subroutine FORMFAC, the Coulomb electric and magnetic form factors are computed. These form factors are defined in Appendix A, section 3. They are essentially spherical Bessel transforms of the appropriate densities (see (A.27)). However, in FORMFAC the form factors are not calculated as such, but are directly calculated from the one-body density matrix. For example, if we seek the Bessel transform of the density $(f \| \theta_\lambda(r) \| i)$ with kernel $j_\ell(qr)$, then in (9) we replace the partially reduced matrix element $(a \| \theta_\lambda(r) \| b)$ by the reduced matrix element $\langle a \| j_\ell(qr) \theta_\lambda(r) \| b \rangle$. The point is that for harmonic oscillators such reduced matrix elements can be evaluated analytically (by the routine RADQ) and the form factor thus obtained provides at least a consistency check when compared against the (numerically evaluated) Bessel transform of the appropriate density.

In FORMFAC the form factors are calculated for momentum transfer $q = 1$ to $Q_{MAX}(\text{MeV}/c)$, in steps of DQ .

3.5 OUTPUT

The calculated densities and form factors are printed in a format that is self-explanatory (see sample output on p.92).

When IPU=1, the density will also be punched on the card. The first card will contain the first three words of NAME and other relevant information (see card MICRØD.95). Then $\tilde{\rho}^c(x_1)$, $\tilde{\rho}^m(x_1)$, $\tilde{\rho}^c(x_2)$, $\tilde{\rho}^m(x_2)$... are punched, in format (6E12.5). NX points are punched, with $x_1=0$, $x_{n+1}-x_n = DX$. When KIND = COULOMB, $\tilde{\rho}^c(x) = \tilde{\rho}(x)$, and $\tilde{\rho}^m(x) = 0$.

When IPL = 1 or 3, the output includes plots of the density vs. x . When IPL = 2 or 3, the natural-logarithmic of the form factor squared is plotted against q .

3.6 SUBROUTINES AND THEIR FUNCTIONS

Subroutines that are used in MICRØDENS are described in Table 2.

TABLE 2. ROUTINES IN MICRØDENS

<i>Name</i>	<i>Function</i>	<i>Called by</i>
MICRØD	Main program	
CURRENT	Computes densities	MICRØD
FORMFAC	Computes (e, e') form factors	MICRØD
TYDEL or TYDELQ	Computes $(\ell_a \ T_\lambda(Y_\ell \Theta V) \ \ell_b)$ or $\langle \ell_a \ j_\ell(qr) T_\lambda(Y_\ell \Theta V) \ \ell_b \rangle$	FORMFAC, CURRENT
TJYLL	Computes $\langle \ell_a \ j_\ell(qr) T_\lambda(Y_\ell \Theta L) \ \ell_b \rangle$	FORMFAC
RDFUNC	Computes oscillator function $R_a(x)$	CURRENT, TYDEL
RADQ	Computes the integral $\int_0^\infty x^2 dx R_a(x) j_\ell(Qx) R_b(x)$	FORMFAC, TJYLL, TYDEL
CG000	Computes the 3-j symbol $\begin{pmatrix} \ell_a & \ell_b & \ell \\ 0 & 0 & 0 \end{pmatrix}$	TJYLL, TYDEL
CLEBS	Computes the 3-j symbol $\begin{pmatrix} a & b & \lambda \\ \frac{1}{2} & -\frac{1}{2} & 0 \end{pmatrix}$	FORMFAC, CURRENT
WCOEF	Computes Racah or 6-j Coefficients	FORMFAC, CURRENT, TJYLL, TYDEL
LINPLØT	Linear plot	MICRØD
SEMILOG	Semi-log plot	MICRØD

The subroutines LINPLOT and SEMILOG call a system routine PLOT, for plotting purposes. This routine is available at the computing center at Chalk River Nuclear Laboratories only. Users at other computer installations may substitute local versions of plotting routines for PLOT. If this proves impossible, the calls, in MICROD, to SEMILOG and LINPLOT must be bypassed.

A complete listing of MICRODENS is given in Appendix C.

CREDITS

The routine RADQ, of which a revised version is used here, was originally written by F.C. Khanna. WCOEF, a high speed program which employs the Regge symmetries to calculate $C-j$ coefficients, was written by R.Y. Cusson. The two plotting routines LINPLOT and SEMILOG were kindly supplied by M.A. Lone.

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APPENDIX A

In this appendix we present a collection of formulas describing some physical processes most often exploited to probe the nuclear charge and current densities, namely γ -emission, photonuclear-disintegration leading to particle emission, and elastic and inelastic electron scattering. These formulas will all be given explicitly in terms of the density functions ρ_λ , $\rho_{\lambda\ell}^{c,m}$ described in Section 2. Detailed attention will be given only to the development of the formalism for the (e,e') in the distorted wave Born approximation, in view of the fact that this topic has received a less than comprehensive treatment in the literature.

A special property of electromagnetism is its gauge invariance which is closely related to the conservation of charge. A very useful relation which arises from this property is the so-called Siegert's theorem. We shall start the appendix on this topic.

A.1 GAUGE INVARIANCE, CHARGE CONSERVATION AND SIEGERT'S THEOREM

We shall use the compact 4-vector notation $x_\nu = (\vec{x}, ct)$, $x^\nu = (\vec{x}, -ct)$. The 4-vector electromagnetic potential is $\vec{A}_\nu = (\vec{A}, \phi)$, the current is $J_\nu = (\vec{J}/c, \rho)$, and the

gradient is $\partial_\nu = (\vec{\nabla}, -\frac{1}{c} \frac{\partial}{\partial t})$. The gauge invariance of electromagnetism states that Maxwell's equations are invariant under a gauge transformation of the second kind⁸⁾,

$$A'_\nu = A_\nu + \partial_\nu \Lambda \quad (A.1)$$

where Λ is any function of x_ν . The gauge invariance is closely related to the conservation of charge.

Consider the interaction of A'_ν with current J_ν . We have

$$\begin{aligned} \int A'_\nu J^\nu dx^4 &= \int A_\nu J^\nu d^4x + \int (\partial_\nu \Lambda) J^\nu d^4x \\ &= \int A_\nu J^\nu d^4x - \int \Lambda (\partial_\nu J^\nu) d^4x, \end{aligned}$$

integrating by parts. The second term vanishes, however, because charge is conserved:

$$\partial_\nu J^\nu = \frac{1}{c} (\vec{\nabla} \cdot \vec{J} + \frac{\partial \rho}{\partial t}) = 0. \quad (A.2)$$

The relation above is also called the equation of continuity.

The gauge invariance allows us, for example, to work in the Lorentz gauge, expressed by the equation

$$\partial_\nu A^\nu = \vec{\nabla} \cdot \vec{A} + \frac{1}{c} \frac{\partial \phi}{\partial t} = 0$$

We shall see that this is consistent with describing A as being generated by a conserved current, J_ν , via a retarded Green's function⁹⁾ G_{ret} ,

$$A_\nu(r) = -i \int J_\nu(r') G_{\text{ret}}(r, r') d^4 r' \quad (\text{A.3})$$

where

$$G_{\text{ret}}(r, r') = \frac{1}{|\vec{r} - \vec{r}'|} \delta(t - t' - \frac{1}{c} |\vec{r} - \vec{r}'|).$$

The Lorentz condition is automatically satisfied,

$$\begin{aligned} \partial_\nu A^\nu(r) &= -i \int J^\nu(r') \partial_\nu G_{\text{ret}}(r, r') d^4 r' \\ &= +i \int J^\nu(r') \partial'_\nu G_{\text{ret}}(r, r') d^4 r' \\ &= -i \int (\partial'_\nu J^\nu(r')) G_{\text{ret}}(r, r') d^4 r' = 0. \end{aligned}$$

Since J_ν is conserved, $\partial_\nu J^\nu = 0$.

From now on we shall restrict our discussion only to fields with definite frequencies

$$A_\nu(r) = A_\nu(\vec{r}) e^{-i\omega t}. \quad (\text{A.4})$$

Naturally, fields of any time-dependence can be constructed as integrals of such fields. Let us consider a field

created by a source at infinity. In this case Maxwell's equations, in the Lorentz gauge, become

$$\nabla^2 A_\nu + k^2 A_\nu = 0,$$

where $k = \omega/c$. The three independent solutions can be multipole decomposed⁶⁾

$$\phi(\vec{r}) = a \sum_{\lambda} \phi_{\lambda 0}(\vec{r}) \quad (\text{A.5a})$$

$$\vec{A}(\vec{r}) = a \sum_{\lambda} \vec{\mathcal{E}}_{\lambda 0} + \sum_{\lambda, \lambda=\pm 1} b_{\mu} (\vec{\mathcal{E}}_{\lambda \mu} + \mu \vec{\mathcal{M}}_{\lambda \mu}(\vec{r})) \quad (\text{A.5b})$$

with

$$\phi_{\lambda \mu} = [4\pi(2\lambda+1)]^{1/2} i^{\lambda} j_{\lambda}(kr) Y_{\lambda}^{\mu}(\hat{r}) \quad (\text{A.5c})$$

$$\vec{\mathcal{E}}_{\lambda \mu} = (ik)^{-1} \vec{\nabla} \phi_{\lambda \mu} = -i^{1+\lambda} \sqrt{4\pi} (\sqrt{\lambda} j_{\lambda-1} \vec{Y}_{\lambda, \lambda-1, 1}^{\mu} + \sqrt{\lambda+1} j_{\lambda+1} \vec{Y}_{\lambda, \lambda+1, 1}^{\mu}), \quad (\text{A.5d})$$

$$\vec{\mathcal{E}}_{\lambda \mu} = [k^2 \lambda(\lambda+1)]^{-1/2} \vec{\nabla} \times \vec{L} \phi_{\lambda \mu} = i^{1+\lambda} \sqrt{4\pi} (\sqrt{\lambda+1} j_{\lambda-1} \vec{Y}_{\lambda, \lambda-1, 1}^{\mu} - \sqrt{\lambda} j_{\lambda+1} \vec{Y}_{\lambda, \lambda+1, 1}^{\mu}), \quad (\text{A.5e})$$

$$\vec{\mathcal{M}}_{\lambda \mu} = (\lambda(\lambda+1))^{-1/2} \vec{L} \phi_{\lambda \mu} = i^{\lambda} \sqrt{4\pi(2\lambda+1)} j_{\lambda} \vec{Y}_{\lambda \lambda 1}^{\mu} \quad (\text{A.5f})$$

where the direction of the z-axis is along $\vec{k}$, and a and $b_{\pm 1}$ are arbitrary c-numbers. $\vec{\mathcal{E}}_{\lambda \mu}$, $\vec{\mathcal{E}}_{\lambda \mu}$, $\vec{\mathcal{M}}_{\lambda \mu}$ are respectively the longitudinal and transverse electric, and magnetic multipoles. Note that by setting the gauge field Λ in (A.1) to be $\frac{ia}{k} \phi e^{+i\omega t}$, we reduce the gauge transformed ϕ' and $\vec{\mathcal{E}}'$ to zero. Only $\vec{\mathcal{E}}'$ and $\vec{\mathcal{M}}'$ are left, resulting in a two-component field

perpendicular to k . To see this differently we let the field interact with current $J_v(r) = J_v(\vec{r})e^{i\omega t}$.

We get

$$\begin{aligned} A_v J^v d^4r &= \frac{1}{c} \int \vec{E} \cdot \vec{J} d^4r - \int \phi \rho d^4r + (\vec{E} \text{ and } \vec{M} \text{ terms}) \\ &= -\frac{1}{i\omega} \int \phi \vec{\nabla} \cdot \vec{J} d^4r + \int \phi \rho d^4r + \dots \end{aligned} \quad (A.6)$$

From the equations of continuity, and (A.4), we have

$$\vec{\nabla} \cdot \vec{J} = -\frac{\partial \rho}{\partial t} = -i\omega \rho. \quad (A.2')$$

Therefore the first two terms in (A.6) cancel. In other words, due to charge conservation, only the $\vec{E}$ and $\vec{M}$ fields are felt by the current.

To proceed further, we expand the fields in powers of kr . Note that since the nucleus is finite $\langle (kr)^n \rangle \leq k^n R^n$, where R is the nuclear radius. So for long wave-length photons, $\frac{1}{k} \ll R$, this expansion is meaningful. We keep only the lowest order term. Then

$$\vec{\mathcal{E}}_{\lambda\mu} = \frac{1}{2} \sqrt{\frac{\lambda}{2\lambda+1}} \frac{(kr)^{\lambda-1}}{(2\lambda-1)!!} \vec{Y}_{\lambda,\lambda-1,1}^{\mu} (1 + O(k^2 r^2)), \quad (\text{A.7a})$$

$$\vec{\mathcal{E}}_{\lambda\mu} = - \frac{1}{2} \sqrt{\frac{\lambda+1}{2\lambda+1}} \frac{(kr)^{\lambda-1}}{(2\lambda-1)!!} \vec{Y}_{\lambda,\lambda-1,1}^{\mu} (1 + O(k^2 r^2)). \quad (\text{A.7b})$$

Thus, to lowest order

$$\vec{\mathcal{E}}_{\lambda\mu} \approx - \sqrt{\frac{\lambda+1}{\lambda}} \vec{\mathcal{E}}_{\lambda\mu}. \quad (\text{A.8})$$

Using the equations of continuity again, we have

$$\frac{1}{c} \int \vec{\mathcal{E}}_{\lambda\mu} \cdot \vec{j} \, d^3r \approx \sqrt{\frac{\lambda+1}{\lambda}} \int \phi_{\lambda\mu} \rho \, d^3r. \quad (\text{A.9})$$

The above relation which states that in the long wave length limit a knowledge of the charge density of matter alone is sufficient to describe the interaction between the matter and the electric potential, is generally referred to as "Siegert's theorem"²⁾. The RHS of (A.9) is obviously easier to compute than the LHS. Conceptually it is also easier to grasp, as it is related to static properties of the nucleus.

It should be stressed that in actual calculations, (A.9) will be realized only if (A.2) is explicitly satisfied. For example (A.9) will not be

realized when $\vec{J}^{\text{exch}}$ is non-vanishing and not included in $\vec{J}$. Therefore it is always advantageous to use the RHS of (A.9) whenever possible, since unlike (the $\vec{J}^{\text{exch}}$ term in) $\vec{J}$ the charge density is known. One of the prescriptions that will guarantee gauge invariance and charge conservation is to replace the operator $\frac{\hbar}{i} \vec{\nabla}$ in (3) by the canonical momentum operator $[\vec{H}, \vec{r}]$. This will include in $\vec{J}$ the exchange current due to the explicit dependence of the interaction on momentum. There are other more subtle reasons for the breakdown of (A.2). For example, truncating the basis in a shell model calculation is equivalent to making the two-body interaction effectively momentum-dependent, even if it is not explicitly so. This part of the $\vec{J}^{\text{exch}}$ can be either eliminated (or reduced) by enlarging the basis¹⁰⁾ or it can be calculated by perturbation theory. In the literature the latter procedure is referred to as "renormalization" of the density operators^{11,4,12)}.

There is no equivalent of Siegert's theorem for the magnetic interaction $\int \vec{m}_{\lambda\mu} \cdot \vec{J} d^4r$. Although the contributions from the magnetization current of (4) almost always dominates this interaction, in some cases³⁾ the contributions of the exchange current are known to be important.

Siegert's theorem can be generalized so that for the electric interactions explicit reference to the current is deleted for all photon energies. Sachs and Austern¹³⁾ used gauge invariance and Foldy¹⁴⁾ used the conservation of charge to achieve this goal. The end results of these two approaches are equivalent but not identical. Essentially they involve different series expansions of the interactions in powers of the photon energy $\hbar\omega$. In both cases the expansions can be identified with the usual multipole expansion only in the long wave-length limit.

The equation of continuity, (A.2), can be expressed in terms of multipole densities. For the Fourier component $J_v(r) = J_v(\vec{r})e^{i\omega t}$, we have, from (A.2) and (11,12),

$$k\rho_\lambda(r) + \sqrt{\frac{\lambda+1}{2\lambda+1}} \left(\frac{d}{dr} + \frac{\lambda+2}{r} \right) \rho_{\lambda,\lambda+1}(r) + \sqrt{\frac{\lambda}{2\lambda+1}} \left(\frac{d}{dr} - \frac{\lambda-1}{r} \right) \rho_{\lambda,\lambda-1}(r) = 0$$

(A.2')

where $k = \omega/c$.

A.2 GAMMA-EMISSION

The rates of transition from $|\Psi_i\rangle$ to $|\Psi_f\rangle$, or the numbers of photons of energy $\hbar\omega$ emitted per second, are¹⁵⁾

$$T_{\lambda\mu}^E = \frac{2k}{\hbar(2\lambda+1)} \left| \int \vec{\mathcal{E}}_{\lambda\mu} \cdot \frac{\vec{J}}{c} d^3r \right|^2 \quad (\text{A.10a})$$

and

$$T_{\lambda\mu}^M = \frac{2k}{\hbar(2\lambda+1)} \left| \int \vec{\mathcal{M}}_{\lambda\mu} \cdot \frac{\vec{J}}{c} d^3r \right|^2 \quad (\text{A.10b})$$

respectively due to electric and magnetic λ -pole γ -transitions. Using Siegert's theorem and (11)

$$T_{\lambda\mu}^E = 8\pi e^2 (C_{i,\lambda\mu}^f)^2 \frac{(\lambda+1)}{\lambda[(2\lambda+1)!!]^2} \frac{k^{2\lambda+1}}{\hbar} \left(\int_0^\infty r^{\lambda+2} \rho_\lambda(r) dr \right)^2 \quad (\text{A.11a})$$

in the long wave-length limit. Here $C_{i,\lambda\mu}^f = \langle J_i M_i \lambda\mu | J_f M_f \rangle$.

We may also use the current given by (12), in which case

$$T_{\lambda\mu}^E = \frac{8\pi k e^2}{\hbar} \frac{1}{2\lambda+1} (C_{i,\lambda\mu}^f)^2 \left\{ \int_0^\infty r^2 dr \sqrt{\lambda+1} j_{\lambda-1}(kr) \rho_{\lambda,\lambda-1}(r) + \sqrt{\lambda} j_{\lambda+1}(kr) \rho_{\lambda,\lambda+1}(r) \right\}^2. \quad (\text{A.11b})$$

Similarly

$$T_{\lambda\mu}^M = \frac{8\pi k e^2}{\hbar} (C_{i,\lambda\mu}^f)^2 \left(\int_0^\infty r^2 dr j_\lambda(kr) \rho_{\lambda\lambda}(r) \right)^2 \quad (\text{A.11c})$$

In the long wave-length limit, a condition invariably satisfied in nuclear γ -transitions, we may again expand the spherical Bessel functions and keep the terms lowest order in k . Summing over M_f and μ and averaging over M_i , we define

$$\begin{aligned} T_{M\lambda}^{(E)}(i \rightarrow f) &= \frac{1}{2J_i+1} \sum_{m_i, m_f, \mu} T_{\lambda\mu}^{E,M}(i \rightarrow f) \\ &= \frac{8\pi(\lambda+1)}{\lambda[(2\lambda+1)!!]^2} \frac{k^{2\lambda+1}}{n} B_M^{(E)}(\lambda; i \rightarrow f). \end{aligned} \quad (A.12)$$

B is called the transition strength. From (A.11)

$$B(E\lambda; i \rightarrow f) = e^2 \frac{2J_f+1}{2J_i+1} \left(\int r^{\lambda+2} \rho_{\lambda}^{i \rightarrow f}(r) dr \right)^2, \quad (A.13a)$$

$$= e^2 \frac{2J_f+1}{2J_i+1} \frac{(2\lambda+1)\lambda}{k^2} \left(\int r^{\lambda+1} \rho_{\lambda, \lambda-1}^{i \rightarrow f}(r) dr \right)^2, \quad (A.13b)$$

$$B(M\lambda; i \rightarrow f) = e^2 \frac{2J_f+1}{2J_i+1} \frac{\lambda}{\lambda+1} \left(\int r^{\lambda+2} \rho_{\lambda\lambda}^{i \rightarrow f}(r) dr \right)^2. \quad (A.13c)$$

The strengths are simply moments of the appropriate transition densities. The dimensionality of $B(\lambda)$ is $e^2 L^{2\lambda}$. Note that the RHS of (A.13b) depends explicitly on the photon energy, whereas the expression in (A.13a) does not. This is a reflection of the continuity

equation, which must be satisfied if the equality of (A.13a) and (A.13b) is to be realized. In the literature (A.13a) is invariably used, as its evaluation requires no explicit statement being made of the current.

As for the evaluation of all operators involving the gradient operator, different expressions can be obtained by integrating by parts. We shall use this technique to derive a more familiar expression for $B(M\lambda)$. From (20) and (27) (second line) we have

$$\begin{aligned} & \sqrt{\frac{\lambda}{\lambda+1}} \int r^{\lambda+2} \rho_{\lambda\lambda}^{i \rightarrow f}(r) dr \\ &= \frac{\hbar}{2Mc} \sqrt{\frac{\lambda}{\lambda+1}} i^\lambda \langle \Psi_f \| -2ig_L r^\lambda \vec{Y}_{\lambda\lambda} \cdot \vec{\nabla} + \mu_S (\vec{\nabla} \times r^\lambda \vec{Y}_{\lambda\lambda}) \cdot \vec{\sigma} \| \Psi_i \rangle. \end{aligned} \quad (A.14)$$

Using the identities[†]

$$r^{\lambda+\mu} \vec{Y}_{\lambda\lambda}^\mu \cdot \vec{\nabla} = -\sqrt{\frac{1}{\lambda(\lambda+1)}} \vec{\nabla} (r^\lambda Y_\lambda^\mu) \cdot \vec{L},$$

and

$$\vec{\nabla} \times (r^\lambda \vec{Y}_{\lambda\lambda}^\mu) = i \sqrt{\frac{\lambda+1}{\lambda}} \vec{\nabla} (r^\lambda Y_\lambda^\mu)$$

we have

$$\sqrt{\frac{\lambda}{\lambda+1}} \int r^{\lambda+2} \rho_{\lambda\lambda}^{i \rightarrow f}(r) dr = \langle \Psi_f \| \frac{\hbar}{2Mc} i^{1+\lambda} \left\{ \frac{2g_L}{\lambda+1} \vec{\nabla} (r^\lambda Y_\lambda) \cdot \vec{L} + \mu_S \vec{\nabla} (r^\lambda Y_\lambda) \cdot \vec{\sigma} \right\} \| \Psi_i \rangle. \quad (A.15)$$

[†] The following identities are more general:

$$\vec{M}_{\lambda\mu} \cdot \vec{\nabla} / i = -\frac{k}{\sqrt{\lambda(\lambda+1)}} \vec{L}_{\lambda\mu} \cdot \vec{L}; \quad \langle \vec{M}_{\lambda\mu} \cdot (\vec{\nabla} \times \vec{\sigma}) \rangle = k \langle \vec{L}_{\lambda\mu} \cdot \vec{\sigma} \rangle$$

The operators on the RHS are the usual *effective* magnetic λ -pole operators due to charge and spin respectively.

The evaluation of the complete matrix element can be found in the literature⁶⁾. However, we may also separate the angular and radial integrations and thereby define the appropriate densities. Using the identity

$$\vec{\nabla}(r^\lambda Y_\lambda^\mu) = \sqrt{\lambda(2\lambda+1)} r^{\lambda-1} \vec{Y}_{\lambda,\lambda-1,1}^\mu$$

and defining the orbital and intrinsic spin densities

$$\rho_{\lambda\ell}^{L,i\rightarrow f} = \frac{\hbar}{Mc} \frac{1}{\sqrt{\lambda(\lambda+1)}} (\psi_f \| i^\ell g_L \vec{Y}_{\lambda\ell 1} \cdot \vec{L} \| \psi_i) \quad (A.16)$$

$$\rho_{\lambda\ell}^{S,i\rightarrow f} = \frac{\hbar}{2Mc} (\psi_f \| i^\ell \mu_S \vec{Y}_{\lambda\ell 1} \cdot \vec{\sigma} \| \psi_i), \quad (A.17)$$

We have

$$B(M\lambda; i\rightarrow f) = e^2 \frac{2J_f+1}{2J_i+1} \frac{\lambda}{\lambda+1} (2\lambda+1) \left\{ \int_0^\infty r^{\lambda+1} (\sqrt{\lambda} \rho_{\lambda,\lambda-1}^{L,i\rightarrow f} + \sqrt{\lambda+1} \rho_{\lambda,\lambda-1}^{S,i\rightarrow f}) dr \right\}^2. \quad (A.13d)$$

We use (9) to evaluate the partially reduced matrix elements.

We give the relevant single-particle matrix elements

$$(a \| i^{1+\lambda} \vec{Y}_{\lambda\ell 1} \cdot \vec{L} \| b) = (-)^{j_a + \ell_b - \lambda - \frac{1}{2}} j_b \hat{\ell}_a W(\ell_a \ell_b j_a j_b; \lambda \frac{1}{2}) (\ell_a \| i^{1+\lambda} \vec{Y}_{\lambda\ell 1} \cdot \vec{L} \| \ell_b) \quad (A.18a)$$

$$\begin{aligned} (\ell_a \| i^{1+\lambda} \vec{Y}_{\lambda\ell 1} \cdot \vec{L} \| \ell_b) &= (-)^{\ell_a + \ell_b} i^{\ell_b + 1 + \lambda - \ell_a} \left(\frac{\ell_b(\ell_b+1)}{4\pi} \right)^{\frac{1}{2}} \hat{\ell} \hat{\lambda} (2\ell_b+1) \\ &\times W(\ell_b 1 \ell_a \ell; \ell_b \lambda) \begin{pmatrix} \ell_a & \ell & \ell_b \\ 0 & 0 & 0 \end{pmatrix} u_a^*(r) u_b(r) \end{aligned} \quad (A.18b)$$

$$(a \| i^{1+\lambda} \vec{Y}_{\lambda, \lambda+1, 1} \cdot \vec{\sigma} \| b) = (-)^{\ell_a + \ell_b + j_b + \frac{1}{2}} i^{\ell_b + 1 + \lambda - \ell_a} j_b \begin{pmatrix} j_a & j_b & \lambda \\ \frac{1}{2} & -\frac{1}{2} & 0 \end{pmatrix}$$

$$\times (x_a + x_b)^{\lambda+1} (4\pi)^{\frac{(\lambda+1)}{\lambda}} u_a^*(r) u_b(r); \quad \underline{\ell_a + \ell_b + \lambda \text{ odd}}, \quad (\text{A.19a})$$

and

$$(a \| i^{\lambda} \vec{Y}_{\lambda \lambda 1} \cdot \vec{\sigma} \| b) = -(-)^{j_b + \lambda - \frac{1}{2}} i^{\ell_b + \lambda - \ell_a} j_b \hat{\lambda} \begin{pmatrix} j_a & j_b & \lambda \\ \frac{1}{2} & -\frac{1}{2} & 0 \end{pmatrix}$$

$$\times (x_a - x_b) (4\pi \lambda (\lambda+1))^{-\frac{1}{2}} u_a^*(r) u_b(r); \quad \underline{\ell_a + \ell_b + \lambda \text{ even}}. \quad (\text{A.19b})$$

In the literature, the transition strength is often expressed in terms of the so-called Weisskopf units¹⁵⁾ (WU). This unit represents a rough estimate of the single-particle strength of a proton in transition and is given as follows:

$$B_{WU}(E\lambda) = \frac{(1.2)^{2\lambda}}{4\pi} \left(\frac{3}{\lambda+3}\right)^2 A^{2\lambda/3} (e^2 F^{2\lambda}),$$

$$B_{WU}(M\lambda) = \frac{0.11(1.2)^{2\lambda-2}}{\pi} \left(\frac{3}{\lambda+3}\right)^2 A^{(2\lambda-2)/3} (e^2 F^{2\lambda}),$$

$$= \frac{10(1.2)^{2\lambda-2}}{\pi} \left(\frac{3}{\lambda+3}\right) A^{(2\lambda-2)/3} \mu_n^2 F^{2\lambda-2},$$

where $\mu_n = \frac{\hbar e}{2Mc} = 0.105 \text{ eF}$ is the nuclear magneton. A is the atomic mass number of the nucleus. The Weisskopf

unit strengths, in units of $e^2 F^{2\lambda}$, for $E\lambda$ and $M\lambda$ transitions, are respectively plotted in Fig. A.1 and A.2, as functions of A .

Experimental data on transitions are also expressed in terms of the (partial) width, $\Gamma(\lambda) = \hbar T(\lambda)$ and the associated lifetime $\tau(\lambda) = T(\lambda)^{-1}$. $T(\lambda)$ is the rate of transition given in (A.12). We may define the kinematic coefficients $t(\lambda)$ and $c(\lambda)$,

$$\tau(\lambda; i \rightarrow f)^{-1} = t(\lambda) \cdot E_Y^{2\lambda+1} \cdot B(\lambda; i \rightarrow f),$$

$$\Gamma(\lambda; i \rightarrow f)^{-1} = c(\lambda) E_Y^{2\lambda+1} \cdot B(\lambda; i \rightarrow f),$$

where $E_Y = \hbar c k$ is the transition energy. From (A.12)

$$\begin{aligned} c(\lambda)\hbar = t(\lambda) &= \frac{8\pi(\lambda+1)}{\lambda[(2\lambda+1)!!]^2} \frac{c}{(\hbar c)^{2\lambda+1}} \left(\frac{e^2}{\hbar c}\right), \text{ if } B(\lambda) \text{ in } e^2 F^{2\lambda} \\ &= \frac{8\pi(\lambda+1)}{\lambda[(2\lambda+1)!!]^2} \frac{c}{(\hbar c)^{2\lambda+1}} \left(\frac{e^2}{\hbar c}\right) \left(\frac{\hbar}{2Mc}\right)^2 \text{ if } B(\lambda) \text{ in } \mu^2 F^{2\lambda-2}. \end{aligned}$$

In table A.1, the values for $t(\lambda)$ and $c(\lambda)$ for $\lambda=1$ to 5 are listed, with τ in units of psec (10^{-12} sec), Γ in units of eV, and E_Y in units of MeV.

TABLE A.1. VALUES FOR COEFFICIENTS $t(\lambda)$ AND $c(\lambda)$

λ	$B(\lambda) \text{ in } e^2 F^{2\lambda}$		$B(\lambda) \text{ in } \mu_n^2 F^{2\lambda-2}$	
	$t(\lambda)$	$c(\lambda)$	$t(\lambda)$	$c(\lambda)$
1	1.60(3)*	1.05	1.76(1)	1.16(-2)
2	1.23(-3)	8.08(-7)	1.36(-5)	8.94(-9)
3	5.72(-10)	3.76(-13)	6.33(-12)	4.16(-15)
4	1.70(-16)	1.12(-19)	1.88(-18)	1.24(-21)
5	3.46(-23)	2.28(-26)	3.83(-25)	2.52(-28)

* The number in the bracket is the exponent of 10, i.e.

$$1.60(3) = 1.60 \times 10^3.$$

A.3 ELECTRON SCATTERING

We shall express all momenta and energies in terms of wave numbers, in units of inverse length. The following notations will be used,

$\vec{p}_i, \vec{p}_f$, momenta of the incident and scattered electrons, respectively;

ϵ_i, ϵ_f , initial and final electron energy;

$\vec{q} \equiv \vec{p}_i - \vec{p}_f$, momentum transferred to the nucleus;

$\kappa \equiv \epsilon_i - \epsilon_f$, energy transferred to the nucleus;

$\vec{P} \equiv \frac{1}{2}(\vec{p}_i + \vec{p}_f)$,

and

$q_\mu^2 \equiv q^2 - \kappa^2$.

The EM potential which interacts with the nucleus is generated by the electron in motion. We use (A.3) and write $J_v^{(e)}$ as

$$J_v^{(e)} = e \bar{\psi}_f \gamma_v \psi_i,$$

where ψ is the electron wavefunction, and γ_v is the Dirac matrix. From (A.3), (A.4), and integrating the time out, we get the time-independent 4-vector potential

$$\delta(\epsilon_i - \epsilon_f - \kappa) A_v(\vec{r}) = -e \delta(\epsilon_i - \epsilon_f - \kappa) \int d^3 r' \bar{\psi}_f(\vec{r}') \gamma_v \psi_i(\vec{r}') \frac{e^{ikR}}{R}. \quad (A.20)$$

$R = |\vec{r} - \vec{r}'|$. The interaction matrix element, to lowest order, is

$$\langle H_{int} \rangle = \int d^3r A_v(\vec{r}) J^v(\vec{r}), \quad (A.21)$$

where $J_v(r)$ is the nuclear current. The differential cross section is

$$\frac{d\sigma}{d\Omega} = \frac{2\pi}{\hbar} |\langle H_{int} \rangle|^2 \frac{\rho_f}{j_e} \quad (A.22)$$

where ρ_f is the final density of state. Including the nuclear recoil correction

$$\rho_f = \frac{1}{1 + \frac{2\epsilon_f}{M_T} \sin^2 \frac{\theta}{2}} \frac{p_f \epsilon_f}{(2\pi)^3 \hbar c}$$

where θ is the scattering angle in the laboratory frame and M_T is the nuclear mass. j_e is the incident electron flux

$$j_e = c p_i / \epsilon_i.$$

(The expression for ρ_f and j_e implies that the electron wavefunctions are normalized in a sphere of unit volume.)

Putting everything together, we have

$$\frac{d\sigma}{d\Omega} = \frac{\epsilon_i \epsilon_f p_f}{(2\pi)^2 (\hbar c)^2 p_i} \frac{1}{1 + (2\epsilon_f \sin^2 \frac{\theta}{2}) / M_T} |\langle H_{int} \rangle|^2. \quad (A.23)$$

Only $\langle H_{int} \rangle$ remains to be calculated.

In the plane wave Born approximation, $\bar{\psi}_f(\vec{r}')\psi_i(\vec{r}')$ is proportional to $e^{-i\vec{q}\cdot\vec{r}'}$. Using this property and the relation

$$(\nabla_{\vec{r}'}^2 + k^2) \frac{e^{ikR}}{R} = -\delta(\vec{r}' - \vec{r}),$$

and by integrating by parts, we have

$$\langle H_{int} \rangle_{PWBA} = \frac{e}{q_\mu^2} \int \bar{\psi}_f(\vec{r}) \gamma_\nu J^\nu(\vec{r}) \psi_i(\vec{r}) d^3r. \quad (A.24)$$

The evaluation of the matrix element is straightforward but somewhat tedious; we only give the result for the cross section averaged over initial and summed over final spin projection¹⁶⁾,

$$\begin{aligned} \frac{d\sigma}{d\Omega} &= \frac{\epsilon_i \epsilon_f p_f}{(2\pi)^2 (\hbar c)^2 p_i} \frac{1}{1+2\epsilon_f/M_T \sin^2 \frac{\theta}{2}} \frac{1}{2(2J_i+1)} \sum_{M' S} |\langle H_{int} \rangle_{PWBA}|^2 \\ &= \frac{Z^2 \alpha^2}{q_\mu^4} \frac{p_f}{p_i} \frac{2}{1+2\epsilon_f/M_T \sin^2 \frac{\theta}{2}} \cdot \{V_L(\theta) \sum_{\lambda=0} |F_\lambda^C(q^2)|^2 \\ &\quad + V_T(\theta) \sum_{\lambda=1} (|F_\lambda^M(q^2)|^2 + |F_\lambda^E(q^2)|^2)\} \end{aligned} \quad (A.25)$$

where

$$V_L(\theta) = \frac{1}{2} \frac{q_\mu^4}{q} [(\epsilon_i + \epsilon_f)^2 - q^2] \quad (A.26a)$$

$$V_T(\theta) = p^2 - \left(\frac{\vec{p} \cdot \vec{q}}{q}\right)^2 + \frac{1}{2} q_\mu^2. \quad (A.26b)$$

The Coulomb, electric and magnetic form factors, F^C , F^E and F^M , are given by

$$Z \frac{\hat{J}_i \hat{\lambda}}{\hat{J}_f} F_\lambda^C(q^2) = \frac{1}{e} \langle \Psi_f \| \phi_\lambda \rho_{op} \| \Psi_i \rangle = \sqrt{4\pi} \hat{\lambda} \int \rho_{\lambda, \lambda}(r) j_\lambda(qr) r^2 dr \quad (A.27a)$$

$$\begin{aligned} Z \frac{\hat{J}_i \hat{\lambda}}{\hat{J}_f} F_\lambda^E(q^2) &= \frac{1}{ec} \langle \Psi_f \| \vec{\epsilon}_\lambda \cdot \vec{J}_{op} \| \Psi_i \rangle \\ &= \sqrt{4\pi} \int \{ \sqrt{\lambda+1} \rho_{\lambda, \lambda-1}(r) j_{\lambda-1}(qr) + \sqrt{\lambda} \rho_{\lambda, \lambda+1}(r) j_{\lambda+1}(qr) \} r^2 dr \end{aligned} \quad (A.27b)$$

$$Z \frac{\hat{J}_i \hat{\lambda}}{\hat{J}_f} F_\lambda^M(q^2) = \frac{1}{ec} \langle \Psi_i \| \vec{m}_\lambda \cdot \vec{J}_{op} \| \Psi_i \rangle = \sqrt{4\pi} \hat{\lambda} \int \rho_{\lambda\lambda}(r) j_\lambda(qr) r^2 dr. \quad (A.27c)$$

The multipole fields $\phi_{\lambda\mu}$, $\vec{\epsilon}_{\lambda\mu}$ and $\vec{m}_{\lambda\mu}$ are those defined in (A.5) (with k replaced by q) and ρ_λ and $\rho_{\lambda\ell}$ are the nuclear transition densities defined in Section 2.

When the momentum transfer is small, $qR \lesssim 1$, where R is the nuclear size, we may use Siegert's theorem, resulting in

$$F_\lambda^E(q^2) \approx \frac{k}{q} \sqrt{\frac{\lambda+1}{\lambda}} F_\lambda^C(q^2).$$

We have the extra factor k/q which is not equal to unity because in general the exchanged photon will be off the mass-shell.

As in our discussion of the $B(M\lambda)$ strength, the orbital and spin densities, defined in (A.16) and (A.17), can be used to compute the magnetic form factor. Thus (see footnote on p.49)

$$\begin{aligned} \frac{1}{ce} \langle \Psi_f \| \vec{M}_\lambda \cdot \vec{J}_{op} \| \Psi_i \rangle &= - \frac{\hbar q}{2Mc} \langle \Psi_f \| 2q_L \vec{\mathcal{P}}_{\lambda\mu} \cdot \frac{\vec{L}}{\sqrt{\lambda(\lambda+1)}} - \mu_s \vec{\mathcal{C}}_{\lambda\mu} \cdot \vec{\sigma} \| \Psi_i \rangle \\ &= -\sqrt{4\pi} q \int [(\sqrt{\lambda} \rho_{\lambda,\lambda-1}^L + \sqrt{\lambda+1} \rho_{\lambda,\lambda-1}^S) j_{\lambda-1}(qr) - (\sqrt{\lambda+1} \rho_{\lambda,\lambda+1}^L - \sqrt{\lambda} \rho_{\lambda,\lambda+1}^S) j_{\lambda+1}(qr)] r^2 dr. \end{aligned}$$

(A.27c')

Clearly the LHS of (A.27c) has a simpler appearance. It should also be pointed out that (A.27c) is formally correct for the complete current (including $\vec{J}^{exch}$) density whereas (A.27c') is specialized for $\vec{J}^c$ and $\vec{J}^m$ only. In general the advantage ρ^L and ρ^S have over ρ^c and ρ^m is that the former do not involve the derivative of the nuclear wavefunction. When q is small compared to the nuclear size, only the first two terms on the RHS of (A.27c') need be retained, since $|j_{\lambda+1}(qr)/j_{\lambda-1}(qr)| \ll 1$. In this case, just as in the case of static $M\lambda$ transitions in the long wave-length limit, it is more advantageous to use the orbital and intrinsic spin densities. This is especially so for $M1$, since $\rho_{1,0}^{L,S}$ are quite simple to calculate.

When the charge of the target is large, or more specifically, when the Coulomb potential energy of the scattered electron is not negligible compared to its kinetic energy, distortions of the incoming and outgoing electron waves must be considered. In this case, the condition leading to the simplified expression of (A.24) is no longer satisfied, and partial wave expansions for both the initial and final electron waves must be carried out.

Although the expression for the (e, e') cross section in the distorted wave Born approximation (DWBA) exist in the literature¹⁷⁾, the derivation is often less than transparent and the end result incomplete. We shall therefore present a complete derivation of the formulas here. Readers who are only interested in the end result shall find them in (A.41) and (A.45).

We shall use the abbreviated notations $f = \int d^3r' d^3r$,
 $C_{j_a j_b}^{j_c} = C_{j_a j_b}^{j_c m_c} = \langle j_a m_a j_b m_b | j_c m_c \rangle$, and $\tilde{Y}_{\lambda \pm}^{\mu} \equiv \tilde{Y}_{\lambda, \lambda \pm 1, 1}^{\mu}$ and $\tilde{Y}_{\lambda}^{\mu} \equiv \tilde{Y}_{\lambda \lambda 1}^{\mu}$.

We start from (A.20) and (A.21). Using an explicit representation of the Dirac matrices, and expanding⁹⁾ out the Green's function, we have

$$\begin{aligned}
 H_{int} &= -e \int \psi_f^\dagger(\vec{r}') (\vec{p}(\vec{r}) - \vec{a} \cdot \vec{J}(\vec{r})/c) \psi_i(\vec{r}) e^{ikR/R} \\
 &= -ie k \cdot \vec{a} \sum_{\lambda=0, \mu} \int \psi_f^\dagger(\vec{r}) (\vec{p} - \vec{a} \cdot \frac{\vec{J}}{c}) \psi_i j_\lambda(kr_<) h_\lambda^{(1)}(kr_>) Y_{\lambda}^{*\mu}(\hat{r}') Y_{\lambda}^{\mu}(\hat{r}), \quad (A.28)
 \end{aligned}$$

where $r_<(r_>)$ is the smaller (larger) of r' and r , and $\vec{J} = \begin{pmatrix} J_x \\ J_y \\ J_z \end{pmatrix}$ is an operator in the electron Hilbert space only. In the following we shall use the abbreviation

$$h_\lambda(r') h_\lambda(r) = j_\lambda(kr_<) h_\lambda^{(1)}(kr_>).$$

For any vectors A and B , we have the identities

$$\begin{aligned}
 &\sum_{\ell=0}^{\infty} \int d\Omega A(\vec{r}') \cdot B(\vec{r}) Y_{\ell}^{m*}(\hat{r}') Y_{\ell}^m(\hat{r}) h_{\ell}(r') h_{\ell}(r) \\
 &= \sum_{\ell=0}^{\infty} \sum_{m, \nu} (-1)^{\nu} (\vec{A} \cdot \hat{e}_{-\nu} Y_{\ell}^{m*} h_{\ell}) (\vec{B} \cdot \hat{e}_{\nu} Y_{\ell}^m h_{\ell}) \\
 &= \sum_{\lambda, \mu} \sum_{\ell=\lambda, \lambda \pm 1} \vec{A} \cdot \vec{Y}_{\lambda \ell 1}^{\mu*}(\hat{r}') h_{\ell}(r') h_{\ell}(r) \vec{Y}_{\lambda \ell 1}^{\mu}(\hat{r}) \cdot \vec{B} \quad (A.29a)
 \end{aligned}$$

$$= \sum_{\lambda \mu} \frac{1}{4\pi(2\lambda+1)} \vec{A} \cdot \{ \vec{\mathcal{L}}_{\lambda \mu}^*(\hat{r}') \vec{\mathcal{L}}_{\lambda \mu}(\hat{r}) + \vec{\mathcal{E}}_{\lambda \mu}^*(\hat{r}') \vec{\mathcal{E}}_{\lambda \mu}(\hat{r}) + \vec{\mathcal{M}}_{\lambda \mu}^*(\hat{r}') \vec{\mathcal{M}}_{\lambda \mu}(\hat{r}) \} \cdot \vec{B} \quad (A.29b)$$

where $\hat{e}_{\pm 1}, \hat{e}_0$ are the spherical unit vectors and $\vec{\mathcal{L}}, \vec{\mathcal{E}}, \vec{\mathcal{M}}$ are the spherical vector tensors defined in (A.5), with j_ℓ replaced by h_ℓ (i.e., $j_\ell(kr_<)$ or $h_\ell^{(1)}(kr_>)$). In the summation over λ , λ starts from 0 for $\vec{\mathcal{L}}$, and from 1 for $\vec{\mathcal{E}}$ and $\vec{\mathcal{M}}$. The reason

for having the two expansions (A.29a) and (A.29b) will be made clear later. Briefly the former will lead to a simpler expression whereas only the latter will allow us to make explicit and easy use of the equation of continuity. Putting (A.29) into (A.28), we have

$$\begin{aligned} \langle H_{\text{int}} \rangle = & -4 \text{ iek} \left\{ \int \sum_{\lambda\mu} h_{\lambda} h_{\lambda} (\psi_f^{\dagger} (-i)^{\lambda} Y_{\lambda}^{*\mu} \psi_i) (\rho \text{ i}^{\lambda} Y_{\lambda}^{\mu}) \right. \\ & \left. - \sum_{\lambda\ell\mu} h_{\ell} h_{\ell} (\psi_f^{\dagger} (-i)^{\ell} \vec{\alpha} \cdot \vec{Y}_{\lambda\ell 1}^{*\mu} \psi_i) (\text{i}^{\ell} \vec{Y}_{\lambda\ell 1}^{\mu} \cdot \vec{J}/c) \right\} . \end{aligned} \quad (\text{A.30})$$

We shall do the angular integration

for the RHS of (A.30) first. From the definition of the nuclear charge and current densities, (11) and (12), we immediately have

$$\begin{aligned} \int \rho \text{ i}^{\lambda} Y_{\lambda}^{\mu} d\Omega_r &= e C_{1,\lambda\mu}^f \rho_{\lambda}(r), \\ \int \text{i}^{\ell} \vec{Y}_{\lambda\ell 1}^{\mu} \cdot \vec{J}/c d\Omega_r &= e C_{1,\lambda\mu}^f \rho_{\lambda\ell}(r). \end{aligned} \quad (\text{A.31})$$

To do the same for the electron part we need explicit expressions for the electron wavefunctions. We follow the notation of Rose¹⁸⁾, and write for the incoming and outgoing functions

$$\psi_{(\rho, \vec{r})}^{(\pm)\sigma} = \frac{4\pi}{\sqrt{2}} \sum_{\kappa, m, j} e^{\pm i\delta_{\kappa}} i^{\ell} C_{\ell m \frac{1}{2}\sigma}^{j\nu} Y_{\ell}^{m*}(\hat{p}) \psi_{\kappa}^{\nu}(\vec{r}) \quad (\text{A.32a})$$

where δ_{κ} is the phase-shift, σ is the spin projection, and ψ_{κ}^{ν} is a two-component wavefunction,

$$\psi_{\kappa}^{\nu}(\vec{r}) = \begin{pmatrix} g_{\kappa}(r) \phi_{\kappa}^{\nu}(\hat{r}) \\ i f_{\kappa}(r) \phi_{-\kappa}^{\nu}(r) \end{pmatrix} \quad (\text{A.32b})$$

$$\phi_{\kappa}^{\nu}(\hat{r}) = \sum_{m, \sigma} C_{\ell_{\kappa} m, \frac{1}{2}\sigma}^{j_{\kappa} \nu} Y_{\ell_{\kappa}}^m(\hat{r}) \chi_{\frac{1}{2}}^{\sigma} \quad (\text{A.32c})$$

with $j_{\kappa} = |\kappa| - \frac{1}{2}$, and $\ell_{\kappa} = \kappa$, if $\kappa > 0$ and $\ell_{\kappa} = -\kappa - 1$, if $\kappa < 0$. The radial wavefunctions g and f are solutions of the coupled radial Dirac equations¹⁸⁾ and are functions of the electron energy and the nuclear charge distribution:

$$\frac{dg_{\kappa}}{dr} = (E + m_e - V) f_{\kappa} - \frac{\kappa + 1}{r} g_{\kappa} \quad (\text{A.32d})$$

$$\frac{df_{\kappa}}{dr} = \frac{\kappa - 1}{r} f_{\kappa} - (E - m_e - V) g_{\kappa} \quad (\text{A.32e})$$

where $E = \epsilon_1$ or ϵ_k and $V(r)$ are the appropriate static Coulomb potential generated by the nucleus. With (A.32), we have, for any tensor operator O_λ^μ

$$(\psi_f^{\sigma'} | O_\lambda^\mu | \psi_i^\sigma) = 8\pi^2 \sum_{\kappa \kappa' m m'} e^{i(\delta_\kappa + \delta_{\kappa'})} i^{\ell - \ell'} C_{\ell m \frac{1}{2} \sigma}^{j \nu} C_{\ell' m' \frac{1}{2} \sigma'}^{j' \ell'} Y_{\ell}^{m*}(\hat{p}_i) Y_{\ell'}^{m'}(\hat{p}_f) (\psi_{\kappa'}^{\nu'} | O_\lambda^\mu | \psi_{\kappa}^{\nu}). \quad (A.33)$$

From (A.32b)

$$\begin{aligned} (\psi_{\kappa'}^{\nu'} | (-i)^\lambda Y_\lambda^{*\mu} | \psi_{\kappa}^{\nu}) \\ = (-)^{\lambda+\mu} C_{j\nu, \lambda-\mu}^{j' \nu'} \{g_\kappa, g_\kappa(\phi_\kappa, \| i^\lambda Y_\lambda \| \phi_\kappa) + f_\kappa, f_\kappa(\phi_{-\kappa}, \| i^\lambda Y_\lambda \| \phi_{-\kappa})\} \end{aligned} \quad (A.34)$$

and

$$\begin{aligned} (\psi_{\kappa'}^{\nu'} | (-i)^\lambda \vec{\alpha} \cdot \vec{Y}_{\lambda \ell 1}^{*\mu} | \psi_{\kappa}^{\nu}) = -i(-)^{\lambda+\mu} C_{j\nu, \lambda-\mu}^{j' \nu'} \\ \times \{g_\kappa, f_\kappa(\phi_\kappa, \| i^{\lambda \rightarrow} \vec{Y}_{\lambda \ell 1} \| \phi_{-\kappa}) - f_\kappa, g_\kappa(\phi_{-\kappa}, \| i^{\lambda \rightarrow} \vec{Y}_{\lambda \ell 1} \| \phi_\kappa)\} \end{aligned} \quad (A.35)$$

where $j_\kappa = j_\kappa$, $j' = j_{\kappa'}$. From (14), and (A.19)

$$\begin{aligned} (\phi_{\pm k}, \| i^\lambda Y_\lambda \| \phi_{\pm \kappa}) = (-)^{j - \frac{1}{2} + \lambda} i^\lambda \frac{\hat{\lambda} \hat{j}}{\sqrt{4\pi}} \begin{pmatrix} j' & j & \lambda \\ \frac{1}{2} & -\frac{1}{2} & 0 \end{pmatrix}, \quad \ell_\kappa + \ell_{\kappa'} + \lambda \text{ or } \ell_{-\kappa} + \ell_{-\kappa'} + \lambda \text{ even} \\ = 0 \quad \text{otherwise.} \end{aligned} \quad (A.36)$$

$$\begin{aligned} (\phi_{\pm k}, \| i^\lambda Y_\lambda \| \phi_{\mp \kappa}) = (-)^{j - \frac{1}{2} + \lambda} i^\lambda \frac{\hat{\lambda} \hat{j}}{\sqrt{4\pi}} \begin{pmatrix} j' & j & \lambda \\ \frac{1}{2} & -\frac{1}{2} & 0 \end{pmatrix}, \quad \ell_\kappa + \ell_{\kappa'} + \lambda \text{ odd} \\ = 0 \quad \text{otherwise.} \end{aligned} \quad (A.37)$$

$$(\phi_{\pm\kappa}, \| i^{\lambda+1} \vec{\sigma} \cdot \vec{Y}_{\lambda+1} \| \phi_{\mp\kappa}) = i \frac{\pm(\kappa' - \kappa) + \lambda + 1}{\hat{\lambda} \sqrt{\lambda + 1}} (\phi_{\kappa}, \| i^{\lambda} Y_{\lambda} \| \phi_{\kappa}) \quad (\text{A.38})$$

$$(\phi_{\pm\kappa}, \| i^{\lambda-1} \vec{\sigma} \cdot \vec{Y}_{\lambda-1} \| \phi_{\mp\kappa}) = -i \frac{\pm(\kappa' - \kappa) - \lambda}{\hat{\lambda} \sqrt{\lambda}} (\phi_{\kappa}, \| i^{\lambda} Y_{\lambda} \| \phi_{\kappa}) \quad (\text{A.39})$$

$$(\phi_{\pm\kappa}, \| i^{\lambda} \vec{\sigma} \cdot \vec{Y}_{\lambda} \| \phi_{\mp\kappa}) = \mp \frac{\kappa' + \kappa}{\sqrt{\lambda(\lambda+1)}} (\phi_{\kappa}, \| i^{\lambda} Y_{\lambda} \| \phi_{-\kappa}) \quad (\text{A.40})$$

Putting (A.33-40) into (A.30a), we get

$$\begin{aligned} \langle H_{\text{int}} \rangle_{\text{DWBA}} = & -i 32\pi^2 e^2 k \sum_{\substack{\lambda\mu\kappa \\ \kappa'mm'}} C_{i,\lambda i}^f e^{i(\delta_{\kappa} + \delta_{\kappa'})} i^{\ell-\ell'+\lambda} (-)^{j-\frac{1}{2}+\mu} \frac{\hat{\lambda}\hat{j}}{\sqrt{4\pi}} \begin{pmatrix} j' & j & \lambda \\ \frac{1}{2} & -\frac{1}{2} & 0 \end{pmatrix} \\ & \times C_{\ell m \frac{1}{2} \sigma}^{j \nu} C_{\ell' m' \frac{1}{2} \sigma}^{j' \nu'} C_{j \nu, \lambda - \mu}^{j' \nu'} y_{\ell}^{m*}(\hat{p}_i) y_{\ell'}^{m'}(\hat{p}_f) \\ & \times \int \left\{ j_{\ell}(kr_{<}) h_{\ell}^{(1)}(kr_{>}) (T_{\kappa', \kappa \lambda}^S - i T_{\kappa', \kappa \lambda}^M) \right. \\ & \left. - \sum_{\ell=\lambda \pm 1} j_{\ell}(kr_{<}) h_{\ell}^{(1)}(kr_{>}) T_{\kappa', \kappa \lambda \ell}^{\text{EL}} \right\} r'^2 dr' r^2 dr \end{aligned} \quad (\text{A.41a})$$

where

$$T_{\kappa', \kappa \lambda}^S = G_{\kappa', \kappa}^+(r') \rho_{\lambda}(r) \quad (\text{A.41b})$$

$$T_{\kappa', \kappa \lambda, \lambda-1}^{\text{EL}} = \frac{1}{\hat{\lambda} \sqrt{\lambda}} (\lambda F_{\kappa', \kappa}^-(r') - (\kappa' - \kappa) F_{\kappa', \kappa}^+(r')) \rho_{\lambda, \lambda-1}(r) \quad (\text{A.41c})$$

$$T_{\kappa', \kappa \lambda, \lambda+1}^{\text{EL}} = \frac{1}{\hat{\lambda} \sqrt{\lambda+1}} ((\lambda+1) F_{\kappa', \kappa}^-(r') + (\kappa' - \kappa) F_{\kappa', \kappa}^+(r')) \rho_{\lambda, \lambda+1}(r) \quad (\text{A.41d})$$

$$T_{\kappa', \kappa \lambda}^M = \frac{\kappa' + \kappa}{\sqrt{\lambda(\lambda+1)}} F_{\kappa', \kappa}^+(r') \rho_{\lambda \lambda}(r) \quad (\text{A.41e})$$

$$G_{\kappa, \kappa}^+(r') = g_{\kappa, \kappa}(r')g_{\kappa, \kappa}(r') + f_{\kappa, \kappa}(r')f_{\kappa, \kappa}(r') \quad (\text{A.41f})$$

$$F_{\kappa, \kappa}^{\pm}(r') = g_{\kappa, \kappa}(r')f_{\kappa, \kappa}(r') \pm f_{\kappa, \kappa}(r')g_{\kappa, \kappa}(r') \quad (\text{A.41g})$$

where $\ell = \ell_{\kappa}$, $\ell' = \ell_{\kappa'}$, $j = j_{\kappa}$, $j' = j_{\kappa'}$. Note that for T^S and T^{EL} , $\ell + \ell' + \lambda$ must be even, whereas for T^M , $\ell + \ell' + \lambda$ must be odd. It follows that other than the phase shift factor $e^{i(\delta_{\kappa} + \delta_{\kappa'})}$, $\langle H_{int} \rangle$ is pure imaginary.

Equation (A.41) has been the basis for essentially all the computer programs written for (e, e') in DWBA. An updated and improved version of one of these, the code DUELS, originally written by the Yale-Duke-Ohio group^{19,20}, is available at CRNL.

It should be pointed out that in the development above, nowhere was the equation of continuity used. This results in T^S being due solely to the scalar multipole, and T^{EL} being due both to the longitudinal and transverse electric multipoles. Consequently T^S is not the equivalent of the Coulomb term F^C in PWBA. The latter includes the contribution from the longitudinal multipole as well. Similarly, T^{EL} is not the equivalent of F^E in PWBA. The latter is due only to the transverse electric multipole. It is desirable to combine the contributions to the scattering from the charge distribution and the longitudinal current. One does not do this for

esthetic reasons alone. For a many-body nucleus ($A > 2$), the approximation used in calculating the wavefunction often leads to a current which is not conserved, even when the Hamiltonian is formally gauge invariant. Therefore it is advantageous to work in a formalism that does not overly rely on the accuracy of the (irrotational component of the) current. In the following we shall derive an expression for the (e, e') amplitude in which explicit knowledge of the longitudinal current is not needed.

We start from equation 28, and use the shorthand notations for the electron current and density

$$\vec{j}^{\dagger} \equiv \psi_f^{\dagger}(\vec{r}') \vec{\alpha} \psi_i(\vec{r}'); \quad \rho_{(e)}^{\dagger} \equiv \psi_f^{\dagger}(\vec{r}') \psi_i(\vec{r}').$$

In making use of the continuity equation (A.2), we will have to do some integration by parts. Therefore it is necessary to separate the electron and nuclear coordinates. First we recall the integral representation

$$i4\pi k j_{\ell}(kr_{<}) h_{\ell}^{(1)}(kr_{>}) = 8 \int_0^{\infty} \frac{j_{\ell}(qr') j_{\ell}(qr) q^2 dq}{q^2 - k^2 - i\epsilon}. \quad (\text{A.42})$$

Substituting into (A.28), we get

$$\langle H_{int} \rangle = -8e \int d\vec{r}' d\vec{r} \int_0^\infty \frac{q^2 dq}{q^2 - k^2 - i\epsilon} \\ \times \sum_{\ell, m} \left\{ \rho_{(e)}^\dagger(\vec{r}') \rho(\vec{r}) - \vec{j}^\dagger(\vec{r}') \vec{j}(\vec{r})/c \right\} j_\ell(qr') Y_\ell^{m*}(\hat{r}') j_\ell(qr) Y_\ell^m(\hat{r}).$$

The electron and nuclear coordinates are now completely separated. We now use (A.29b) and get

$$\langle H_{int} \rangle = -8e \int d\vec{r}' d\vec{r} \int_0^\infty \frac{q^2 dq}{q^2 - k^2 - i\epsilon} \left\{ \sum_{\ell m} \rho_{(e)}^\dagger \rho j_\ell j_\ell Y_\ell^{m*} Y_\ell^m \right. \\ \left. - \sum_{\lambda, \mu} \frac{1}{4\pi(2\lambda+1)} \sum_{\mathcal{R}} (\vec{j}(\vec{r}') \cdot \vec{\mathcal{R}}_{\lambda\mu}(\vec{r}'))^\dagger (\vec{j}(\vec{r})/c \cdot \vec{\mathcal{R}}_{\lambda\mu}(\vec{r})) \right\}, \quad (A.43)$$

where $\sum_{\mathcal{R}}$ means summing over the three functionals $\vec{\mathcal{L}}$, $\vec{\mathcal{E}}$ and $\vec{\mathcal{M}}$ defined in (A.5). We are especially interested in the term when $\mathcal{R}$ is

$$\vec{\mathcal{L}}_{\lambda\mu}(\vec{r}) = \sqrt{4\pi} \hat{\lambda}(iq)^{-1} \vec{\nabla}(j_\lambda(qr) i^\lambda Y_\lambda^\mu(\hat{r})).$$

In this case

$$\frac{1}{4\pi(2\lambda+1)} \int d\vec{r}' (\vec{j}(\vec{r}') \cdot \vec{\mathcal{L}}_{\lambda\mu}(\vec{r}'))^\dagger \int d\vec{r} (j_\lambda(r)/c \cdot \vec{\mathcal{L}}(\vec{r})) \\ = \frac{k^2}{q^2} \int d\vec{r}' d\vec{r} \rho_\lambda(\vec{r}')^\dagger \rho_\lambda(\vec{r}) j_\lambda(qr') j_\lambda(qr) Y_\lambda^{\mu*}(\hat{r}') Y_\lambda^\mu(\hat{r}),$$

where we have integrated by parts, thrown away the vanishing surface terms at infinity, and used (A.2). We get a term which is k^2/q^2 times the scalar term, for each momentum transfer q . Combining this term with the scalar term in (A.43) and recalling that for $\vec{q} = \vec{k}$ the sum over λ is from $\lambda=0$ on, whereas for $\vec{q} = \vec{k}$ or $\vec{q}_i$ it is from $\lambda=1$ on, we have

$$\begin{aligned} \langle H_{\text{int}} \rangle = & -3e \int d\vec{r}' d\vec{r} \left\{ \int_0^\infty dq \rho_{(e)}^+(\vec{r}') \rho(\vec{r}) \sum_{\ell=0, m} j_\ell(qr') j_\ell(qr) Y_\ell^{m*}(\hat{r}') Y_\ell^m(\hat{r}) \right. \\ & \left. - \int_0^\infty \frac{q^2 dq}{q^2 - k^2 - i\epsilon} \sum_{\lambda=1, \mu} \frac{1}{4\pi(2\lambda+1)} \sum_{Q=\ell, m} (j(\vec{r}') \cdot \vec{Q}_{\lambda\mu}(\vec{r}'))^\dagger (j(\vec{r})/c \cdot \vec{Q}_{\lambda\mu}(\vec{r})) \right\} \end{aligned} \quad (\text{A.44})$$

This is a remarkable result. The expression for the Coulomb term (first term in the curly brackets) is similar to that for elastic scattering; that is, for $k=0$. (Naturally the densities $\rho_{(e)}$ and ρ still depend on the energy loss implicitly.) To proceed, we integrate over q and the angular variables, using (A.50e,f), (A.33-40) and (A.42). In using the last equation for the Coulomb term, we take the limit $k \rightarrow 0$,

$$8 \int_0^\infty j_\ell(qr') j_\ell(qr) dq = \frac{4\pi}{2\ell+1} \frac{r_{<}^\ell}{r_{>}^{\ell+1}} .$$

The final result is

$$\begin{aligned}
 \langle H_{int} \rangle_{DWBA} = & -32\pi^3 e^2 \sum_{\substack{\lambda\mu\kappa \\ k'mm'}} C_{i,\lambda\mu}^f e^{i(\delta_\kappa + \delta_{\kappa'})} i^{\ell-\ell'+\lambda} (-)^{j-\frac{1}{2}+\mu} \frac{\hat{\lambda}\hat{j}}{\sqrt{4\pi}} \begin{pmatrix} j' & j & \lambda \\ \frac{1}{2} & -\frac{1}{2} & 0 \end{pmatrix} \\
 & \times C_{\ell m \frac{1}{2} 0}^{j\nu} C_{\ell m \frac{1}{2} 0}^{j'\nu'} C_{j\nu\lambda-\mu}^{j'\nu'} Y_{\ell}^{m*}(\hat{p}_i) Y_{\ell}^m(\hat{p}_f) \int_0^\infty r'^2 dr' \int_0^\infty r^2 dr \\
 & \times \left\{ \frac{1}{2\lambda+1} G_{\kappa'\kappa}^+(r') \rho_{\lambda}(r) \frac{r_{<}^{\lambda}}{r_{>}^{\lambda+1}} P_{\ell\ell'\lambda}^{(+)} \right. \\
 & + k \frac{\kappa'+\kappa}{\sqrt{\lambda(\lambda+1)}} F_{\kappa'\kappa}^+(r') \rho_{\lambda\lambda}(r) j_{\lambda}(kr_{<}) h_{\lambda}^{(1)}(kr_{>}) (1-\delta_{\lambda 0}) P_{\ell\ell'\lambda}^{(-)} \\
 & - ik \frac{\sqrt{\lambda(\lambda+1)}}{\hat{\lambda}(2\lambda+1)} \left[(F_{\kappa'\kappa}^-(r') - \frac{\kappa'-\kappa}{\lambda} F_{\kappa'\kappa}^+(r')) h_{\lambda-1}(kr') \right. \\
 & \left. \left. + (F_{\kappa'\kappa}^-(r') + \frac{\kappa'-\kappa}{\lambda+1} F_{\kappa'\kappa}^+(r')) h_{\lambda+1}(kr') \right] \right. \\
 & \left. \times \left[\sqrt{\lambda+1} \rho_{\lambda,\lambda-1}(r) h_{\lambda-1}(kr) + \sqrt{\lambda} \rho_{\lambda,\lambda+1}(r) h_{\lambda+1}(kr) \right] (1-\delta_{\lambda 0}) P_{\ell\ell'\lambda}^{(+)} \right\}, \quad (A.45)
 \end{aligned}$$

where with $|\lambda'-\lambda| \leq 2$

$$h_{\lambda'}(kr') h_{\lambda}(kr) = \begin{cases} j_{\lambda'}(kr') h_{\lambda}^{(1)}(kr) + i(2\lambda-1) k^{-3} r'^{\lambda-2} / r^{\lambda+1} \delta_{\lambda'+2,\lambda}; & r' \leq r, \\ h_{\lambda'}^{(1)}(kr') j_{\lambda}(kr) + i(2\lambda+3) k^{-3} r^{\lambda} / r'^{\lambda+3} \delta_{\lambda'-2,\lambda}; & r' > r. \end{cases}$$

The function $P^{(\pm)}$ enforces the parity selection rule,

$$P_{\ell\ell'}^{(\pm)} = \frac{1}{2} (1 \pm (-)^{\ell+\ell'+\lambda}).$$

Although in principle they should lead to identical results, (A.45) should be chosen over (A.41) whenever possible. Some of the reasons for advocating this choice have been given earlier. In more practical terms, suppose we are interested in calculating the forward (e, e') amplitude for a natural parity transition. Equation (A.45) only requires that the charge density $\rho_\lambda(r)$ be known whereas (41) requires a knowledge of the current densities $\rho_{\lambda-1}(r)$ and $\rho_{\lambda+1}(r)$ as well. It can be shown that the error incurred in using (A.41) but setting $\rho_{\lambda\pm 1}(r) = 0$ is approximately k^2/q^2 , where $\vec{q}$ is the (asymptotic) momentum transfer. Therefore the error is small when $k^2/q^2 \ll 1$, and (A.45) offers no practical advantage over (A.41) in such cases. On the other hand, when moderate electron energy beams are used to highly excited nuclear states (such as the giant $E\lambda$ resonances), k^2/q^2 can be large, and (A.45) is definitely to be preferred.

Equation (A.45) has the correct limit for elastic scattering, but it is more convenient to use the relation

$$\lim_{k \rightarrow 0} k j_{\lambda}(kr_{<}) h_{\lambda}^{(1)}(kr_{>}) = \frac{-i}{2\lambda+1} \frac{r_{<}^{\lambda}}{r_{>}^{\lambda+1}}$$

for the magnetic term (second term in { }) at the outset. Also in this case the transverse electric (third) term vanishes since $\rho_{\lambda, \lambda \pm 1}(r)$ each vanishes individually, as was shown in section 2.8. We also expect the electron kernel to vanish for this term. We note that the electron kernel is an odd function under the permutation $\kappa' \nleftrightarrow \kappa$, therefore it must be zero when $\psi_f = \psi_i$.

A.4 (γ, N) REACTIONS

When the absorption of a γ -ray leads only to an internal excitation of the nucleus, the process can be described in terms similar to those describing γ -emission. A more complicated reaction occurs when the nucleus disintegrates. Here we describe the reactions (γ, p) and/or (γ, n).

We first recall the general (non-relativistic) expression for a scattering cross section

$$\frac{d\sigma}{d\Omega} = \frac{2\pi}{\hbar} |\langle H_{\text{int}} \rangle|^2 \frac{\rho(E_f)}{j_i}.$$

$\langle H_{\text{int}} \rangle$ is the interaction matrix, in this case

$$H_{\text{int}} = \frac{1}{\sqrt{2}} \sum_{\lambda} \sum_{\mu=\pm 1} (\vec{E}_{\lambda\mu} + i\vec{M}_{\lambda\mu}) \cdot \vec{J}/c \equiv \sum_{\lambda, \mu=\pm 1} O_{\lambda\mu}$$

j_i is the flux of the incoming (unpolarized) beam. For the EM fields given by (A.5), with $b_{+1} = 1/\sqrt{2}$,

$$j_i = k/2\pi\hbar.$$

$\rho_f(E)$ is the number of final states per unit volume per unit energy. For (γ, N) when the wavefunction of the outgoing particle is normalized as

$$\begin{aligned}\psi_{\mathbf{p}_f, \sigma}^{(+)} &= 4\pi \sum_{\ell m} u_{\ell}(r) Y_{\ell}^m(\hat{r}) Y_{\ell}^{m*}(\hat{\mathbf{p}}_f) \chi_{\sigma} \\ &\equiv 4\pi \sum_{\ell, m, a} C_{\ell m \frac{1}{2} \sigma}^{j a m a} \psi_{\alpha}^{(+)}(\hat{r}) Y_{\ell}^{*m}(\hat{k})\end{aligned}\quad (\text{A.46})$$

where $\lim_{r \rightarrow \infty} u_{\ell}(r) = \sin(kr - \frac{\ell\pi}{2} + \text{phase shifts})$, we have (ignoring recoil)

$$\rho_f(E) = \frac{M p_f}{(2\pi\hbar)^3}.$$

Therefore

$$\frac{d\sigma}{d\Omega} = \frac{M}{2\pi\hbar^2} \frac{p_f}{\hbar k} |\langle H_{\text{int}} \rangle|^2. \quad (\text{A.47})$$

Let $|\psi_r\rangle$ be the wavefunction of the residual nucleus.

The total final state is

$$\begin{aligned}|\Psi_f\rangle &= |\psi_{\mathbf{p}_f, \sigma}^{(+)}\rangle |\psi_r\rangle \\ &= 4\pi \sum_{\ell, m, a} C_{\ell m \frac{1}{2} \sigma}^{\alpha} Y_{\ell}^{*m}(\hat{\mathbf{p}}_f) |\psi_{\alpha}^{(+)}\rangle |\psi_r\rangle.\end{aligned}\quad (\text{A.48})$$

Therefore

$$\begin{aligned}
 \langle H_{int} \rangle &= \frac{1}{\sqrt{2}} \sum_{\lambda, \mu=\pm 1} 4\pi \sum_{\ell m a} C_{\ell \frac{1}{2}}^* Y_{\ell}^{*m}(\hat{p}_f) \langle \psi_{\alpha}^{(+)} \psi_r | O_{\lambda \mu} | \psi_1 \rangle \\
 &= \sum_{\lambda, \mu=\pm 1} \frac{4\pi}{\sqrt{2}} \sum_{\ell} Y_{\ell}^{*m}(\hat{p}_f) \sum_{ab} \rho_{ab, \lambda \mu}^{\hat{a}} \frac{\hat{a}}{\lambda} \langle a^{(+)} || O_{\lambda} || b \rangle
 \end{aligned} \tag{A.49}$$

where

$$\rho_{ab, \lambda \mu}^{i \rightarrow f} = (-)^{j_b - m_b} C_{\ell \frac{1}{2}}^{\alpha} C_{\alpha \beta}^{\lambda \mu} C_{J_r, \beta}^i \langle J_r; b || J_i \rangle \tag{A.50}$$

is the equivalent of the one-body transition density matrix.

$\langle J_r; b || J_i \rangle$ is the usual fractional parentage coefficient.

We can now use the formulas developed in section 2.

We find

$$\begin{aligned}
 \langle H_{int} \rangle &= \frac{4\pi e}{\sqrt{2}} \sum_{\lambda, \mu=\pm 1} \sum_{\ell} Y_{\ell}^{*m}(\hat{p}_f) \sqrt{4\pi} \\
 &\times \left\{ \int r^2 dr (\sqrt{\lambda+1} j_{\lambda-1}(kr) \rho_{\lambda \lambda-1}(r) + \sqrt{\lambda} j_{\lambda+1}(kr) \rho_{\lambda, \lambda+1}(r)) \right. \\
 &\left. + \mu \hat{\lambda} \int r^2 dr j_{\lambda}(kr) \rho_{\lambda \lambda}(r) \right\}.
 \end{aligned} \tag{A.57}$$

The similarity among the expressions in (A.51), (A.11), (A.27) and (A.45) that involve the nuclear current densities is quite transparent. In the curly bracket, the first two terms are due to $\vec{\epsilon}_{\lambda \mu}$ and the last term due

to $\vec{M}_{\lambda\mu}$. If the photon wave-length is long, we can again use Siegert's theorem to replace the first two terms by

$$\sqrt{\frac{(2\lambda+1)(\lambda+1)}{\lambda}} \int r^2 dr j_{\lambda}(kr) \rho_{\lambda}(r) \approx \sqrt{\frac{\lambda+1}{\lambda}} \frac{k\lambda}{(2\lambda-1)!!} \int r^{\lambda+2} \rho_{\lambda}(r) dr.$$

When magnetic states M_i and M_f are averaged and summed over, respectively, the parentage states $|J_f; b\rangle$ enter the sum incoherently. Furthermore, when σ is summed over and the cross section integrated over the angle of $\hat{P}_f$, the multipoles contribute incoherently. We get

$$\sigma = \frac{8\pi M c^2 e^2 P_f}{(\hbar c)^2 \hbar k} \sum_{a,b,\lambda} \frac{(2j_a+1)}{(2j_b+1)(2\lambda+1)} | \langle J_r; b | J_i \rangle |^2 | \langle a^{(+)} | 0_{\lambda} | b \rangle |^2.$$

The one-body matrix element can be evaluated straightforwardly using (A.5) and the formulas in Section 2, or alternatively, those given in (A.14-19). We only point out that since the radial wavefunction of the outgoing particle is dimensionless instead of having the normalization $\int |u|^2 r^2 dr = 1$ for bound states, the dimensionality of the matrix element is $L^{3/2}$.

In practice the (γ, N) spectrum (for $A \gg 1$) has a broad resonance structure. Thus at least in the giant resonance region ($E_{\gamma} \approx 20-30$ MeV) cross section is

dominated by the second order effect. That is, the photon is first absorbed by a resonance of the target, and the nucleon is then emitted through the residual nuclear interaction. In this case we have a "second order" term

$$\langle H_{\text{int}}^{(2)} \rangle = \frac{1}{\sqrt{2}} \sum_{n, \lambda, \mu = \pm 1} \langle \Psi | \hat{\mathcal{C}}_{\lambda\mu}^{\dagger} + \mu \hat{\mathcal{M}}_{\lambda\mu}^{\dagger} \cdot \frac{\hat{\mathbf{J}}_{\text{op}}}{c} | \Psi_n \rangle \frac{1}{E_f + E_i - E_n + i\Gamma_n/2} \langle \Psi_n | V_{\text{res}} | \Psi_f \rangle,$$

where E_n and Γ_n are respectively the energy and width of the intermediate (resonant) state $|\Psi_n\rangle$. The electronic matrix element can again be expressed in terms of the currents $\rho_{\lambda\ell}^{i \rightarrow n}$. The description of the nuclear vertex leading to the particle emission may be quite complex, and is outside the scope of this report.

An alternative to the perturbation treatment is to incorporate the resonance effect into the outgoing wave. This involves the solving of coupled-channel equations, as was done by Buck and Hill²¹⁾ and by Raynal et al.²⁰⁾. In this case (A.51) is formally retained, with the resonance (or second and higher order) effect included in the (complex) densities $\rho_{\lambda\ell}$.

FIGURE 1. $B(E\lambda)$ WEISSKOPF UNITS

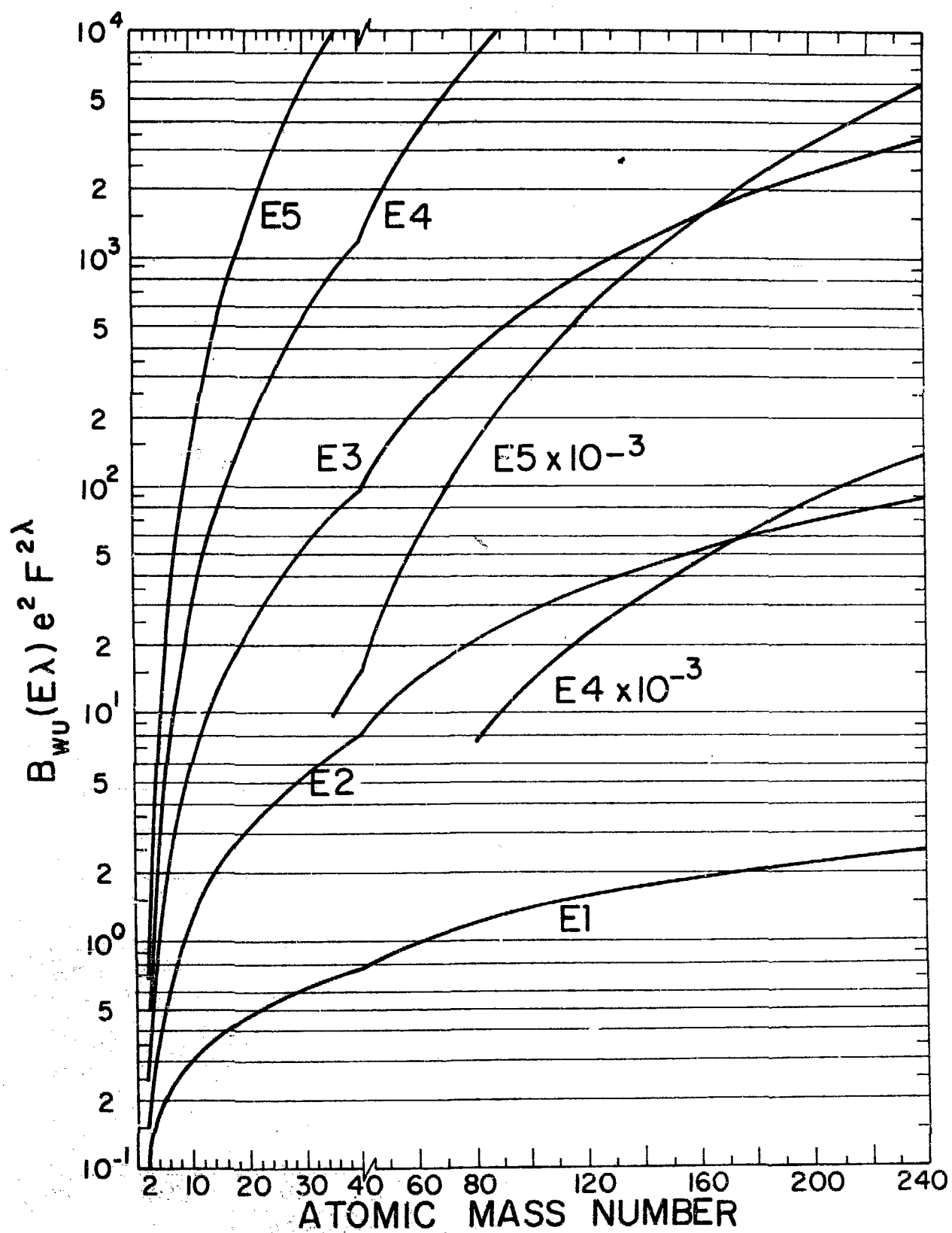
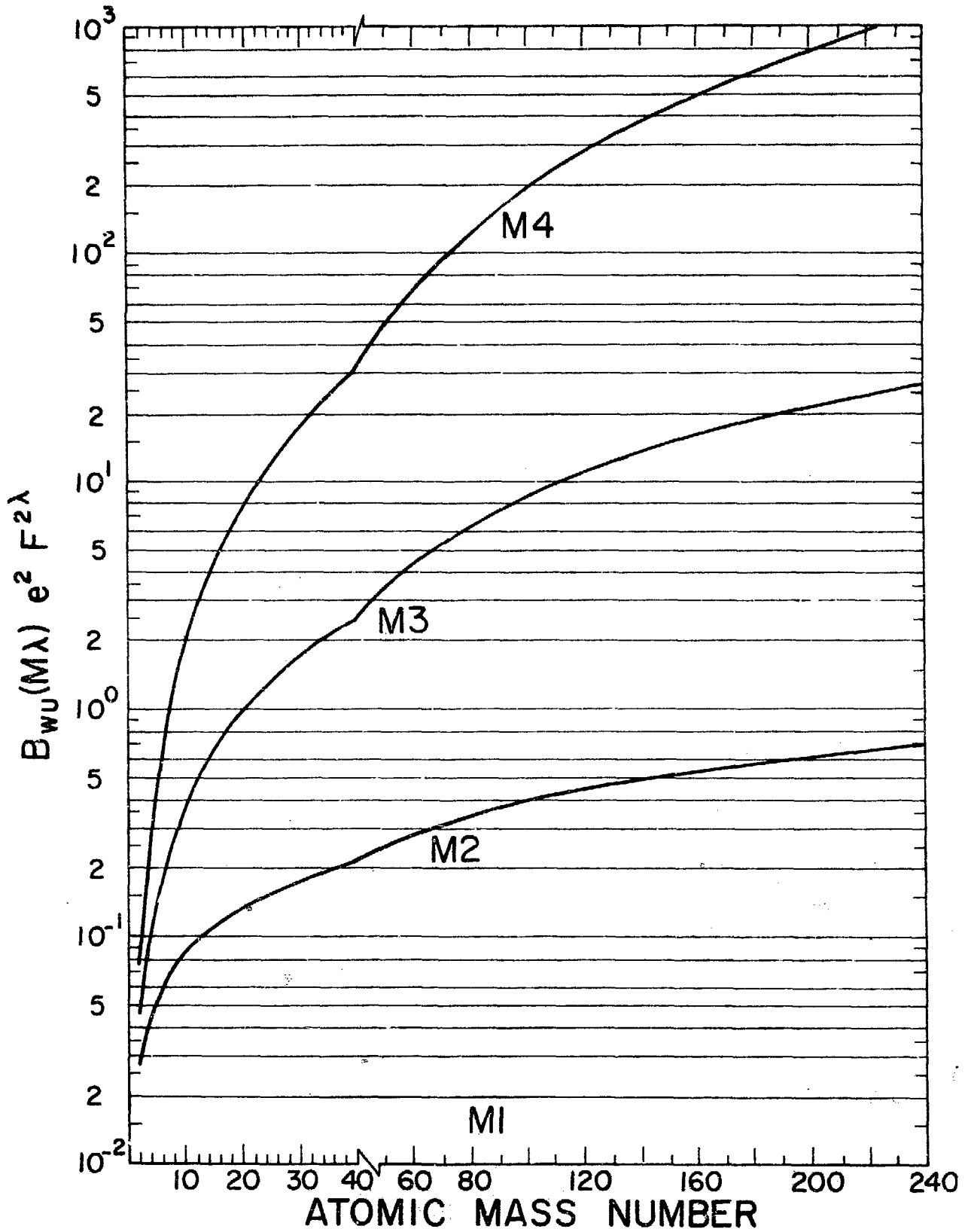


FIGURE 2. $B(M\lambda)$ WEISSKOPF UNITS



APPENDIX B

It was shown in Section 2 that the nuclear structural information leading to the one-body charge and current density is contained in the one-body transition density matrix, ρ . The computation of the density matrix depends on the specific microscopic model being used. In some models, such as the shell model and the Hartree-Fock model, the "complete" wavefunction of the nuclear state is obtained. In such cases the computation of ρ can be quite complex. In other models, such as the random-phase approximation and its variants, the transition amplitude instead of the wavefunction is obtained. In such cases the computation of ρ becomes a rather simple matter.

In this appendix we describe methods to calculate ρ in the three most commonly used nuclear models: The particle-hole model, the shell model and the Hartree-Fock model.

B.1 TRANSITION DENSITY MATRIX IN THE PARTICLE-HOLE MODEL

One of the simplest kinds of microscopic model describing nuclear excitations is the particle-hole model. In the Tamm-Dancoff approximation (TDA)

the particle-hole structure of an excited state of spin J, relative to the 0^+ ground state $|0\rangle$ is

$$|\Psi_{JM}\rangle = Q_{JM}^\dagger |0\rangle \equiv \sum_{ph} x_{ph}^J A_{phJM}^\dagger |0\rangle \quad (B.1)$$

where

$$A_{abJM}^\dagger = \sum_{mB} C_{\alpha\beta}^{JM} s_\beta C_\alpha^\dagger C_\beta; \quad (B.2)$$

p is an orbital not occupied (particle) in $|0\rangle$ and h is an orbital occupied (hole) in $|0\rangle$. The amplitudes x_{ph} are obtained by solving the linearized equation-of-motion²³⁾

$$\langle \Psi_{JM} | [H, Q_{JM}] | 0 \rangle = \omega_J \langle \Psi_{JM} | Q_{JM}^\dagger | 0 \rangle. \quad (B.3)$$

The normalization is

$$\sum_{ph} (x_{ph}^J)^2 = 1. \quad (B.4)$$

Since $A_{abJM}^\dagger$ is in fact just the transition density operator it is easy to see that in TDA

$$\rho_{ba\lambda}^{0 \rightarrow J} = \delta_{J\lambda} \delta_{ap} \delta_{bh} x_{ph}^\lambda \quad (TDA) \quad (B.5)$$

In the random-phase approximation ²³⁾ (RPA), one assumes that there are particle-hole components in the ground state as well and write

$$|\Psi_{JM}\rangle = Q_{JM}^\dagger |0\rangle \equiv \sum_{ph} (x_{ph}^J A_{phJM}^\dagger - (-)^{J-M} y_{ph}^J A_{phJ-M}) |0\rangle, \quad (B.6)$$

and again linearize (B.3). The normalization is now

$$\sum_{ph} ((x_{ph}^J)^2 - (y_{ph}^J)^2) = 1.$$

The non-vanishing density matrix elements are

$$\rho_{hp\lambda}^{0 \rightarrow J} = \delta_{\lambda J} x_{ph}^\lambda, \quad (B.7a)$$

$$\rho_{ph\lambda}^{0 \rightarrow J} = -(-)^{j_p + j_h + \lambda} \delta_{\lambda J} y_{ph}^\lambda. \quad (\text{RPA}) \quad (B.7b)$$

From (B.7) and the Hermitian properties of the density operators (equations (37) and (39)) we have

$$\rho_\lambda^{0 \rightarrow \lambda}(r) = \sum_{ph} \frac{\hat{j}_p}{\hat{\lambda}} (x_{ph}^\lambda + y_{ph}^\lambda) (p \| \rho_{\lambda, 0p} \| h) \quad (B.8a)$$

and

$$\rho_{\lambda\lambda}^{0 \rightarrow \lambda}(r) = \sum_{ph} \frac{\hat{j}_p}{\hat{\lambda}} (x_{ph}^\lambda - y_{ph}^\lambda) (p \| \rho_{\lambda\lambda, 0p} \| h). \quad (B.8b)$$

One should pay attention to the different signs that combine the x and y amplitudes in the charge and current densities.

The particle-hole model is generally applied only to close-shell nuclei. For some medium and heavy nuclei, which have superconducting ground states, the low excited states can be described in terms of two-quasi-particle excitations²⁴⁾. In this model the quasi-particle creation ($a_{\alpha}^{\dagger}$) and annihilation (a_{α}) operators are defined as

$$a_{\alpha}^{\dagger} = (a_{\alpha})^{\dagger} = u_a c_{\alpha}^{\dagger} - (-)^J a^{-m_a} v_a c_{\alpha}, \quad (B.9)$$

and the angular-momentum-coupled two-quasi-particle operator is defined as

$$A_{abJM}^{\dagger} = (A_{abJM})^{\dagger} = \sum_{m_a m_b} c_{\alpha\beta}^{JM} a_{\alpha}^{\dagger} a_{\beta}^{\dagger}. \quad (B.10)$$

v_a (u_a) is the amplitude that the orbital "a" is occupied (unoccupied) in the BCS ground state. In the quasi-particle random-phase-approximation²²⁾ (QRPA) the excitations are described as

$$|\Psi_{JM}\rangle = Q_{JM}^{\dagger} |0\rangle = \frac{1}{2} \sum_{ab} (1+\delta_{ab})^{\frac{1}{2}} (x_{ab}^J A_{abJM} - (-)^{J-M} y_{ab}^{JM} A_{abJ-M}) |0\rangle. \quad (B.11)$$

The x and y amplitudes are obtained as usual by solving (B.3), with the normalization

$$\sum_{a \geq b} ((x_{ab}^J)^2 - (y_{ab}^J)^2) = 1.$$

The density matrix in this case becomes

$$\rho_{ba\lambda}^{0 \rightarrow J} = \delta_{\lambda J} (1+\delta_{ab})^{\frac{1}{2}} (u_a v_b x_{ab}^{\lambda} + v_a u_b y_{ab}^{\lambda}). \quad (\text{QRPA})$$

Similar to (B.8) one can reduce the sum over the pairs (a,b) to that over the ordered pairs $(a \geq b)$ by using the symmetry properties of the amplitudes and matrix elements involved, and get

$$\rho_{\lambda}^{0+\lambda}(r) = \sum_{a \geq b} \frac{\hat{j}_a}{\hat{\lambda}} (1+\delta_{ab})^{-\frac{1}{2}} (u_a v_b + v_a u_b) (x_{ab}^{\lambda} + y_{ab}^{\lambda}) (a \parallel \rho_{\lambda,0p} \parallel b) \quad (B.13a)$$

$$\rho_{\lambda\ell}^{0+\lambda}(r) = \sum_{a \geq b} \frac{\hat{j}_a}{\hat{\lambda}} (1+\delta_{ab})^{-\frac{1}{2}} (u_a v_b - v_a u_b) (x_{ab}^{\lambda} - y_{ab}^{\lambda}) (a \parallel \rho_{\lambda\ell,0p} \parallel b). \quad (B.14b)$$

We note that the pair ($a=b$) does not contribute to the current.

The particle-hole RPA may be generalized so that it becomes applicable to open-shell nuclei with ground states not necessarily of 0^+ . For detail the reader is referred to the literature on the equation-of-motion method²⁵⁾.

B.2 TRANSITION DENSITY MATRIX IN THE SHELL MODEL

In the shell model²⁶⁾, the wavefunction of an n-particle state is described as a linear combination of orthonormal n-particle basis wavefunctions $|\alpha(n)J\rangle$

$$|\Psi_{JM}(n)\rangle = \sum_{\alpha} x_{\alpha}^J |\alpha(n)JM\rangle. \quad (B.15)$$

The (real) amplitudes x_{α}^J are obtained by diagonalizing the Hamiltonian sub-matrix

$$H_{\alpha\alpha'}^J = \langle \alpha(n)JM | H | \alpha'(n)JM \rangle.$$

In the j-j coupling scheme, the transition density matrix is

$$\begin{aligned} \rho_{ab\lambda}^{i \rightarrow f} = n \sum_{\alpha_i \alpha_f} x_{\alpha_i}^{J_i} x_{\alpha_f}^{J_f} \sum_{\alpha_p J_p} (-)^{J_p + \lambda - J_f - j_b} \hat{\lambda} \hat{J}_i W(j_a j_b J_f J_i; \lambda J_p) \\ \times \langle \alpha_f(n)J_f \{ | \alpha_p(n-1)J_p, a \rangle \langle \alpha_p(n-1)J_p, b | \} \alpha_i(n)J_i \rangle. \end{aligned} \quad (B.16)$$

The coefficients of fractional parentage (cfp) $\langle \alpha_p J_p; a | \} \alpha J \rangle$ are defined in the expansion⁶⁾

$$|\alpha(n)JM\rangle = \sum_{\alpha_p J_p M_p \gamma} C_{J_p M_p \gamma}^{JM} |\alpha_p(n-1)J_p M_p \gamma\rangle \langle \alpha_p(n-1)J_p; c | \} \alpha(n)J \rangle. \quad (B.17)$$

The values and phases of cfp's depend on the way the many-particle states are built up. In general they can be expressed in terms of single-shell cfp's, i.e. cfp's for which all particles occupied the same orbit a:

$$\langle \alpha(a^n)J || \alpha_p(a^{n-1})J_p, a \rangle.$$

Some single-shell cfp's have been tabulated^{27,28)}.

For the particularly simple case when $n=2$, the basis wavefunctions are

$$|ab; JM\rangle = \sqrt{\frac{1+\delta_{ab}}{2}} \sum_{m_b m_a} C_{\alpha\beta}^{JM} (|\alpha\rangle|\beta\rangle - (1-\delta_{ab})|\beta\rangle|\alpha\rangle) \quad (B.18)$$

and the cfp's are

$$\langle a, b || \alpha J \rangle = \sqrt{\frac{1+\delta_{ab}}{2}} = -(-)^{j_a+j_b+J} \langle b, a || \alpha J \rangle. \quad (B.19)$$

Writing a two-particle state as

$$|\Psi_{JM}(2)\rangle = \sum_{ab} a_{ab}^J |ab; JM\rangle \quad (B.20)$$

with the normalization

$$\sum_{ab} (a_{ab}^J)^2 = 1,$$

the density matrix becomes

$$\rho_{ba\lambda}^{i \rightarrow f} = \sum_c \sqrt{1+\delta_{ac}} \sqrt{1+\delta_{bc}} a_{ca}^{J_i} a_{cb}^{J_f} (-)^{J_c + \lambda + J_f - J_b} \hat{j}_1 W(j_a j_b J_f J_i; \lambda j_c) . \quad (B.21)$$

B.3 TRANSITION DENSITY MATRIX IN THE PROJECTED HARTREE-FOCK APPROXIMATION

In the Hartree-Fock (HF) approximation²¹⁾, a variational method is used to minimize the expectation value of the Hamiltonian for a single Slater determinant, $|\chi\rangle$.

$$|\chi\rangle = \prod_{i=1}^N c_{\mu_i}^\dagger |0\rangle$$

where N is the number of particles and $\{\mu_i\}$ are the set of single-particle HF orbitals. They are related to the set of basis orbital $\{\alpha\}$ by the linear transformation

$$c_{\mu}^\dagger = \sum_{\alpha} U_{\alpha\mu} c_{\alpha}^\dagger, \quad \sum_{\alpha} (U_{\alpha\mu})^2 = 1. \quad (B.22)$$

In general $|\chi\rangle$ does not have a specific angular momentum, but may be expanded in terms of normalized states with good angular momentum,

$$|\chi\rangle = \sum_{JK} n_{JK} |J(K\chi)K\rangle.$$

If the initial state is $|J_i(K_i\chi_i)M_i\rangle$, and the final state is $|J_f(K_f\chi_f)M_f\rangle$, then the transition density matrix is¹⁰⁾

$$\begin{aligned} \rho_{ba\lambda}^{i \rightarrow f} = & (n_{J_f K_f \chi_f} n_{J_i K_i \chi_i})^{-1} \frac{2J_i+1}{8\pi^2} \sum_{\mu=-\lambda}^{\lambda} C_{J_i K_f -\mu, \lambda \mu}^{J_f K_f} \int d^3\Omega D_{K_f -\mu 1 K_i}^{J*}(\Omega) \\ & \times \sum_{m_a m_b} C_{\alpha \beta}^{\lambda \mu} \langle \chi_f | C_{\alpha \beta}^{\dagger} R(\Omega) | \chi_i \rangle \end{aligned} \quad (B.23)$$

where $R(\Omega) = R(\alpha, \beta, \gamma)$ is the rotation operator specified by the Euler angles (α, β, γ) , $D(\Omega)$ is the rotation matrix and

$$\langle \chi_f | C_{\alpha \beta}^{\dagger} R(\Omega) | \chi_i \rangle = \sum_{\mu \nu} U^1(r)_{\beta \nu} (d^{-1})_{\nu \mu} U_{\alpha \mu}^{f*} \quad (B.24a)$$

$$d_{\mu \nu} = \sum_{\alpha} U_{\alpha \mu}^{f*} U^1(\Omega)_{\alpha \nu} \quad (B.24b)$$

and

$$U_{\alpha \mu}^f(\Omega) = \sum_{\beta} \delta_{J_a J_b} D_{m_a m_b}^J(\Omega) U_{\beta \mu}^f, \quad (B.24c)$$

where U is the transformation of (B.22). In (B.24b), the summations of μ and ν are over occupied orbits (in $|\chi^f\rangle$ and $|\chi^i\rangle$ respectively) only. In general, it is possible to replace the integral in (B.23) by a sum over a certain set of Euler angles, to be determined by the symmetry properties of $|\chi_i\rangle$ and $|\chi_f\rangle$ ¹⁰⁾.

If the initial and final states are described respectively in terms of linear combinations of projected HF states,

$$|\Psi_i\rangle = |\sigma J_i M_i\rangle = \sum_{\sigma} x_{\sigma}^i |J_i(K_{\sigma} \chi_{\sigma}) M_i\rangle \quad (\text{B.25a})$$

and

$$|\Psi_f\rangle = |\sigma' J_f M_f\rangle = \sum_{\sigma'} x_{\sigma'}^f |J_f(K_{\sigma'} \chi_{\sigma'}) M_f\rangle \quad (\text{B.25b})$$

then

$$\rho_{ba\lambda}^{i \rightarrow f} = \sum_{\sigma \sigma'} x_{\sigma'}^f x_{\sigma}^i \rho_{ba\lambda}^{\sigma \rightarrow \sigma'} \quad (\text{B.26})$$

where $\rho_{ba}^{\sigma \rightarrow \sigma'}$ is given by (B.23).

APPENDIX C

This appendix provides the complete listing of MICRØDENS and the outputs for three sample calculations.

The only comments that need be made on the listing is that the "\$" sign is a record separator, and that the subroutines DATE (called at MICRØD.77) and PLØT (called at LINPLØT.37 and SEMILØG.31) are system routines of the CDC-6600 computing system at CRNL. Users at other computer installations may have to use the respective equivalents available locally.

In all three test cases the initial state $|i\rangle$ is the vacuum (0^+) state, and the oscillator length is taken to be 2.0F. Other specifications are given preceding the output for each case.

In the form factor section of the output, quantities listed under the headings F(CHARGE), FC.SQ, F(SPIN), FS.SQ and FTOTAL.SQ are explained in table B.1, where the notations

$$f_{JL}^{C,m}(q) \equiv \int_0^\infty \rho_{JL}^{C,m}(r) j_L(qr) r^2 dr, \quad (C.1)$$

$$f_J(q) \equiv \int_0^\infty \rho_J(r) j_J(qr) r^2 dr, \quad (C.2)$$

and

$$N = \frac{4\pi(2J_f+1)}{(2J_i+1)Z^2} \quad (C.3)$$

are used. The form factors $F_J^{C,M,E}(q^2)$ are those defined in (A.27).

TABLE C.1

<u>TYPE OF SCAT.</u>	<u>Heading</u>				
	<u>F(CHARGE)</u>	<u>FC.SQ</u>	<u>F(SPIN)</u>	<u>FS.SQ</u>	<u>FTOTAL.SQ</u>
COULOMB	$f_j(q)$	$N f_j ^2$			$F_j^C(q^2)$
MAGNETIC	$f_{JJ}^C(q)$	$N f_{JT}^C ^2$	+	††	$F_j^M(q^2)$
ELECTRIC(L=J-1)	$f_{J,J-1}^C(q)$	$\frac{\sqrt{J+1}}{\hat{J}} f_{J,J-1}^C$	+	††	
ELECTRIC(L=J+1)	$f_{J,J+1}^C(q)$	$N\left \frac{\sqrt{J+1}}{\hat{J}} f_{J,J-1}^C + \frac{\sqrt{J}}{\hat{J}} f_{J,J+1}^C\right ^2$	+	††	$F_j^E(q^2)$

+ Same as under F(CHARGE), but replace superscript c by m.

†† Same as under FC.SQ, but replace superscript c by m.

C.1 LISTING OF MICRØDENS

MICROD

```

PROGRAM MICRØD(INPUT,OUTPUT,PUNCH,PLOT)
DIMENSION C(50),S(50),RO(50),NLJA(50),NAME(7),IRO(50),
1 SNAME(3),XX(50),TC(50),IS(50)
COMMON/CFAC/CLP,GLN,MUSP,MUSN,EFC(2),EFM(2),BOHR,TRI
EQUIVALENCE (C,NLJA), (S,NLJ)
REAL MUSP,MUSN
DATA CLP,GLN,MUSP,MUSN/1.,0.,2.79,-1.91/
DATA PI,BOHR,4AC/3.1415 92653, 0.105, 197.33/
DATA MODE/10LELECTRIC /
BOHR = 0.5*MBAR/H(MUCLEON)*C (FERMI)
X= WCOEF(1.,1.,1.,1.,1.,1.)
100 READ 1002, NAME,KIND
*
* IPL=0 NO PLOT. EXL IS = LARGER EXPONENT OF 10 FOR SEMILOG PLOT
* IPL = 1 DENSITY PLOT ONLY. = 2 FORMFACTORS ONLY =3 BOTH
* DENSITY AND FORMFACTORS. NC IS THE NUMBER OF CYCLES OR SEMILOG P
*
READ 1003, N,3,DX,NX,DQ,QMAX,VJI,VJF,Z,EFC,EFM,JS,JL,IPL,IPU
*
* N LENGTH OF RO ARRAY TO BE READ IN
* 3 OSCILLATOR LENGTH IN FERMI'S
* DX INCREMENT IN X. X= R/R
* NX NO. OF X VALUES FOR WHICH DENSITY BE CALCULATED
* FIRST POINT IS X=0. IF NX.LT.2 DENSITY NOT COMPUTED
* DQ INCREMENT IN MOMENTUM TRANSF Q IN MEV/C
* QMAX MAXIMUM VALUE FOR Q
* VJI,VJF SPIN OF INITIAL AND FINAL STATE. IF LEFT PLANK
* VJI IS SET TO 0 AND VJF= MULTIPOLARITY
* JS,JL SMALLEST AND LARGEST MULTIPOLE TO BE CONSIDERED
* IPL PLOT OPTION. =1, PLOT(LINEAR) DENSITY ONLY
* =2, PLOT(SEMI-LOG) FORM FACTORS ONLY
* =3, PLOT BOTH
*
* IPU PUNCH OPTION. =1, PUNCH DENSITY. =0, NO PUNCH
* EFC CHARGE ENHANCEMENT FOR PROTON(1) AND NEUTRON(2)
* EFM MAGNETIC MOMENT ENHANCEMENT
*
IF N=0 JOB TERMINATES. IF N= -1 USE RO-ARRAY OF LAST CASE
KIND MAY BE 10LCOULOMB , OR 10LELECTRIC , OR 10LMAGNETIC .
*
IF ( N.EQ. 0 ) GOTO 99 $ PRINT 1004
PRINT 1003, N,3,DX,NX,DQ,QMAX,VJI,VJF,Z,EFC,EFM,JS,JL,IPL,IPU
*
PRINT1012, NAME,KIND $ IF (N.EQ. -1) GOTO 15 $ NOLD= N
*
READ IN AND CONSTRUCT DENSITY ARRAY IRO
*
READ 1005,( (NLJA(I),NLJR(I),IRO(I),RO(I)), I= 1, N)
PRINT1005,( (NLJA(I),NLJR(I),IRO(I),RO(I)), I= 1, N)
*
NLJ = 32*N+ 2*L + (J+0.5- L)
CONSTRUCT THE WORD IPO AS FOLLOWS
* FIRST 22(0 - 21) BITS STORES ARS(RO)*1.E05
* NEXT 8 BITS(22 - 29) STORES NLJB
* NEXT 8 BITS(30 - 37) STORES NLJA
* THE 50TH BIT STORES ISO (=1 FOR PROTON, =0 FOR NEUTRON)
* THE SIGN OR 60TH BIT HAS THE SIGN OF RO.
*
DO 1 I= 1, N $ IR= ARS(RO(I))*1.E05 $ IS=1 $ IF(RO(I).LT.0) IS= 0
IS= SHIFT(IS,5) $ ISO= SHIFT(IRO(I),5) $ IP= IS .0. ISO .0. IR
1 IPO(I)= IP .0. SHIFT(NLJA(I),30) .0. SHIFT(NLJR(I),22)
*
PRINT 1006, (IPO(I), I= 1, N)
*
N= NOLD $ IF (KIND -400E) 11, 12, 13
15 COULOMB
*
11 ICD= -2 $ JML= 0 $ GOTO 14
* TRANSVERSE ELECTRIC
*
12 ICD= J $ JML= 1 $ GOTO 14
* MAGNETIC
*
13 ICD= 0 $ JML= 0
*
14 DO 2 J= JS,JL,2 $ LS= J- JML $ LL= J+ JML $ DO 2 L= LS,LL,2
C(1)= S(1)= PO(1)= 0 $ PRINT 1007, NAME,KIND,J,L
22= SORT( J/(J+J+1.) ) $ 71= SORT( (J+1.)/(J+J+1.) )
TRI= 1. $ IF(J.GT.L) TRI= -1.
IF (VJI.LT..1) VJF= J
LAPELS FOR PLOT ROUTINES
CALL DATE(SNAME(3)) $ SNAME(2)=10HLM+1000G
SNAME(1)=10000*9 + J

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MICROD (CONTINUED)

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*
* FIRST COMPUTE CHARGE AND SPIN DENSITIES
*
IF (NX .LT. 2) GOTO 3
X=XX(I)=5 ; DO 3 I=2,NX ; X=XX(I)=X+DX
CALL CURRENT(C(I),S(I),ICD,J,L,X,IR0,N,9) ; RO(I)= C(I)+ S(I)
3 PRINT 1001, X,C(I),S(I),RO(I)
IF (IPL.EQ.0) GO TO 5
PUNCH 1018, (NAME(I), I= 1,3),KIND,J,L,DX,NX,9
PUNCH 2000, ( (C(I),S(I)), I= 1, NX )
5 CONTINUE
IF (IPL.EQ.1.OR.IPL.EQ.3) CALL LINPLOT(XX,C,S,DX,NX,NAME,SNAME)

*
* NOW CALCULATE FORM FACTORS IN BORN APPROXIMATION
*
DO 7 I= 1, 150
7 C(I)= 0
PRINT 1009, NAME,KIND,J,L
XNORM= 4.*PI*(VJF+VJF+1.)/(VJI+VJI+1.)/777
NQ= QMAX/DQ ; Q=1.-Q ; DO 4 I= 1,NQ ; XX(I)=Q+Q+ DQ ; QH= Q/HQC
CALL FORMFAC(CF,SS,ICD,J,L,QH*9,IR0,N,9)
IF (L-J) 21,22,23
21 TC(I)= 21*CC ; TS(I)= 71*SS ; GO TO 4
23 C(I)= TC(I)+ 72*CC ; S(I)= TS(I)+ 72*SS ; Y= C(I)+ S(I)
C(I)= C(I)*C(I)*XNORM ; S(I)= S(I)*S(I)*XNORM ; RO(I)=Y*Y*XNORM
GOTO 4
22 C(I)= CC*CC*XNORM ; S(I)= SS*SS*XNORM
Y= CC+ SS ; Y= Y*Y ; RO(I)= Y*XNORM
4 PRINT 1010, Q,QH,CC,C(I),SS,S(I),RO(I)
Y= -1. ; DO 6 I= 1, 150
6 Y= AMAX1(Y, C(I)) ; EXL= ALOG10(Y+Y)
IF (IPL.EQ.2 .OR. IPL.EQ.3)
1 CALL SEMILOG(XX,C,S,RO,DQ,NQ,NAME,SNAME,EXL,-6)
2 CONTINUE ; GOTO 100
99 CONTINUE
1002 FORMAT(8A10)
1003 FORMAT(I4,1X,2F5.3,I5,9F5.1,4I2)
1004 FORMAT(1H1)
1005 FORMAT( 5(2I3,1I,F8.5) )
1006 FORMAT( 4(5X020) )
1007 FORMAT(////7A10/5X,A10* J=*I1* L=*I1*, X=R/R, DENSITY TIMES 9-CUB
1ED IS PRINTED*// 8X*X*13X*CHARGE*15X*SPIN*15X*TOTAL*//)
1008 FORMAT(3A10,X,A10* J=*I1* L=*I1* DX=*F4.2* NX=*I2* R=*F5.3*(F)*)
1001 FORMAT(5X,F6.2,3E20.5)
2000 FORMAT( 6E12.5 )
1009 FORMAT(////7A10,5X*3ORN APPROX *A10* J=*I1* L=*I1* FORM FACTORS*//
1 6X*Q(MEV/C) Q(1/FH)*5X*F(CHARGE)*8X*FC.SQ*9X*F(SPIN)*
2 9X*FS.SQ*8X*F(TOTAL.SQ*//)
1012 FORMAT(1X///8A10/)
1010 FORMAT(5X,2F10.3,6E15.5)
END

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CURRENT

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SUBROUTINE CURRENT(CHARGE,SPIN,ICD,J,L,X,IROA,N,9)
  ROUTINE CALCULATES CURRENT DENSITIES CORRESPONDING TO THE SINGLE
  PARTICLE DENSITY MATRIX, IPOA. CHARGE IS THE DENSITY OF THE
  GRADIENT OPERATOR. SPIN IS THE PURE IMAGINARY PART OF THE DENSITY
  OF THE GRADIENT-CROSS-SPIN (PAULI) OPERATOR.
  IRO = SIGN(RO)*(SHIFT(J,55) + SHIFT(NLJA,30) + SHIFT(NLJA,22)
        + 445(RO)*1.E95)
  RO = RO(IRO,J) . NLJ = SHIFT(N,5) + L*2+ JL, JL= J+ .5,
  X IS DIMENSIONLESS
  N IS THE LENGTH OF THE DENSITY ARRAY IROA
  ICD = -1 COMPUTE CHARGE ONLY
        = 0 COMPUTE BOTH CHARGE AND SPIN
        = +1 COMPUTE SPIN ONLY
  (FOR DETAIL SEE H. P. LEE, ...NUCLEAR CHARGE, CONVECTION CURRENT AND
  MAGNETIZATION CURRENT DENSITIES...AECL REPORT NO. 4839, NOV 19.4)

  CHANGED TO TIME-REVERSAL INVARIANT PHASE ( FEB 2 1974 ) ADDITIONAL
  PHASE (LB+ J- LA)/2 FOR ELECTRIC AND (LB+ J- LA+ 1)/2 FOR MAG.

  COMMON/CFACOR/GLP,GLN,MUSP,MUSN,EFC(2),EFM(2),ROHR,TRI
  REAL MUSP,MUSN,MUS $ LOGICAL SLNAB,SJL
  DIMENSION IROA(1) $ LOGICAL IPASS
  DATA IPASS,ESP/.F.,1.E-5/
  DATA PI/3.1415 92653/, I22,I9/177777778,3778/
  SJL= J.NF.L
  CHARGE= SPIN= 0 $ IF (IARS(J-L).GT. 1.OR. J.LT. 1) RETURN
  FIRST RECOVER QUANTUM NUMBERS AND DENSITY MATRIX ELEMENT FROM IRO
  DO 11 I= 1, N $ IRO= IROA(I) $ ISO= SHIFT(IRO,-59)
  ISO= ISO.A. 1 $ S= ISO $ ISO= IRO.A.I22 $ PO= (S+S- 1.)*ISO*ESP
  ISO= SHIFT(IRO,-58) $ ISO= ISO.A. 1
  ISO= 1 $ PROTON= 0 $ NEUTRON
  GL= ISO*(GLP+ EFC(1))+ (1- ISO)*(GLN+ EFC(2))
  MUS= ISO*(MUSP+ EFM(1))+ (1- ISO)*(MUSN+ EFM(2))
  IRO= SHIFT(IRO, -22) $ IJUMP= 1
  2 NLJ= IRO.A. I9 $ JA= NLJ.A. 1 $ NLJ= NLJ/2 $ LA= NLJ.A. 15
  JA= LA+JA $ NLJ= SHIFT(NLJ,-4) $ NA= NLJ.A. 7 $ GOTO (3,31) IJUMP
  3 LB= LA $ JB= JA $ NB= NA $ IRO= SHIFT(IRO,-8) $ IJUMP=23 GOTO 2
  31 HAVE Q,N AND PO, NOW COMPUTE DENSITY. FIRST CHARGE
  SLNAB= NA.EQ.N9.A. (LA.EQ.L1)
  IF (SLNAB.A. SJL.A. (JA.EQ.JR)) GOTO 11
  IFAS= LB+ J- LA $ IF (ICD.GT.-2) GOTO 4
  THIS SECTION COMPUTES COULOMB TRANSITION DENSITY
  IFAS= JB+ J+ IFAS/2 $ IFAS= IFAS.A. 1 $ FAS= 1- IFAS- IFAS
  FAS= FAS*SQRT( JA*JB/PI)
  CHARGE= CHARGE- FAS*RO*GL*CLERS(JA,JB,J)
  1 *ROFUNC(NA,LA,X)*ROFUNC(NB,LB,X)
  GOTO 11
  4 I6= LA+ LB $ IF (MOD(I6+L,2).EQ.0) GOTO 11 $ I5= JA+ JB- 1
  IF (I5.LT.J.OR. IARS(JA-JR).GT.J) GOTO 11
  IF (SLNAB.A. SJL) GOTO 12
  IF (I6.LT.J.OR. IARS(LA-LB).GT.J) GOTO 12 $ IF (ICD) 5,5,6
  5 I1= I5+ I6 $ I5= I5+ J $ I2= JA+ LB+ J
  IFA= I2+ ( IFAS+1-IARS(J-L) )/2 $ IFA= IFA.A. 1
  I3= LA+ JA $ I4= LB+ JB $ I6= I6+ J $ FAS= 1- IFA- IFA
  FAS= FAS*2.*SQRT(JA*JB/(J+J+1.))*GL
  RACAH IS W(JA,JB,LA,LB,J,1/2)
  TDQ= TDQE= SQRT(LA+LA+1.)*TLYDEL(NA,LA,NB,LB,J,L,X)
  IF (SLNAB) GOTO 13 $ IFA= (J+1).A. 1
  TDQE= SQRT(LB+LB+1.)*TLYDEL(NB,LB,NA,LA,J,L,X)*(1-IFA-IFA)
  13 CHARGE= CHARGE- RO*FAS*FWCOEF(I1,I2,I3,I4,I5,I6)*(TDQ+ TDQE)*0.5
  12 IF (ICD) 11,6,6
  NOW COMPUTE SPIN
  6 RA= ROFUNC(NA,LA,X) $ RB= ROFUNC(NB,LB,X)
  DRAB= 0 $ IF (LA.GT.0) GOTO 7
  RNO= 0 $ IF (NA.GT.0) RNO= SQRT(FLOAT(NA))*ROFUNC(NA-1,1,X)
  RNO=-SQRT(NA+1.5)*ROFUNC(NA,1,X)- RNO+ ROFUNC(NA,0,X)/X
  DRAB= DRAB+ RB*RNO $ GOTO 4
  7 DRAB= DRAB+ RA*( SQRT(LA+ NA+ 0.5)*ROFUNC(NA, LA-1, X)
        + SQRT(NA+ 1.)*ROFUNC(NA+1, LA-1, X) )
  8 IF (LB.GT.0) GOTO 9
  RNO= 0 $ IF (NB.GT.0) RNO= SQRT(FLOAT(NB))*ROFUNC(NB-1,1,X)
  RNO=-SQRT(NB+1.5)*ROFUNC(NB,1,X)- RNO+ ROFUNC(NB,0,X)/X
  DRAB= DRAB+ RA*RNO $ GOTO 10
  9 DRAB= DRAB+ RA*( SQRT(LB+ NB+ 0.5)*ROFUNC(NB, LB-1, X)
        + SQRT(NB+ 1.)*ROFUNC(NB+1, LB-1, X) )
  10 RAB= RA*RB/X $ DRAB= DRAB- (LA+ LB)*RAB

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CURRENT (CONTINUED)

* * * 20 * * 21 22 23 11	<pre> RA= (LA- JA+ 0.5)*(JA+ JA) \$ RA= (LR- JR+ 0.5)*(JB+ JB) RAR= RA*RA/Y , DRAR= D/DX(RA*RA*X*X)/Y**2 , RA= XA , RB= XB S= JA*JA/PI \$ IF (J-L) 20, 21, 22 FOR RO(J,J+1) S= S/(J+1.)/(J+ J+ 1.) S= SORT(S) FOR RO(J,J) S= SORT(S/(J*J+ J))*((RA+RA)*(RA- DRAR)- (J*J+ J)*RAR) \$ GOTO 23 FOR RO(J,J-1) S= S/(J+ J+ 1.) S= SORT(S) S= (RA- RA)*(DRAR+ (J- 1.)*RAR) IFA= J+ JR+ (IFA+1- IABS(J-L))/2 \$ IFA= IFA , A. 1 SPIN= SPIN+ (1- IFA- IFA)*CLES(JA,JB,J)*S*RO*MUS CONTINUE \$ IF (ICQ.EQ. -2) RETURN CHARGE= 2.*BOHR*TRI*CHARGE/B \$ SPIN= BOHR*TRI*SPIN/B \$ RETURN END </pre>	CURRENT 79 CUPPENT 80 CUPPENT 81 CUPPENT 82 CUPPENT 83 CUPPENT 84 CUPPENT 85 CUPPENT 86 CUPPENT 87 CUPPENT 88 CUPPENT 89 CUPPENT 90 CUPPENT 91 CUPPENT 92 CUPPENT 93 CUPPENT 94
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TLYDEL

* * * * * * * * * * 11 * 1 * 2 13 * 3 14 * 4 * 5 6 15 * 5	<pre> FUNCTION TLYDEL(NA,LA,NB,LR,LAM,L,X) ROUTINE CALCULATES REDUCED MATRIX ELEMENT OF TENSORIAL OPERATOR OF ORDER LAM FORMED BY THE VECTOR COUPLING OF THE SPHERICAL HARMONIC OF ORDER L TO THE GRADIENT OPERATOR. ONLY THE ANGLES ARE INTEGRATED OVER. THE RADIAL FUNCTIONS ARE ASSUMED TO HARMONIC OSCILLATOR FUNCTIONS. X = R/P IS DIMENSIONLESS. Q = OSC. LENGTH. THE BRINK AND SACHLER, NOT THE TIME-REVERSAL-INVARIANT PHASE IS USED. LOGICAL RHO DATA PI/3.1415 92653/ RHO= .T. \$ GOTO 11 \$ ENTRY TLYDELO \$ RHO = .F. I4= LA+ LR+ L \$ TLYDEL= 0 \$ GOTO 1 FOLLOWING ENTRY POINT DO TESTING OF TRIANGULAR RELATIONS. ENTRY TLYDEL \$ TLYDEL= 0 \$ IF (MOD(I4,2).NE.1) RETURN IF (IABS(LAM-L) .GT. 1) RETURN IF (LAM .GT. LA+LR .OR. LAM .LT. IABS(LA-LR)) RETURN I5= LAM+ LR \$ FAS= 1- 2*MOD(I5,2) Z= (LAM+LAM+1)*(L+L+1)/4./PI \$ I2= I5+ L+ LR- 1 \$ I3= LR+ LR I5= I5+LA \$ I6= L+LAM+ 1 \$ I1= I4+ 1 \$ I4= I4- 1 FAS= FAS*SORT(7) \$ IF (RHO) FAS= FAS*ROFUNC(NA,LA,X) IF (LR.GT.0) GOTO 13 \$ IF (RHO) GOTO 2 RHO= 0 \$ IF (NR.GT.0) RHO= SORT(FLOAT(NR))*RADQ(NA+1,LA,NB,1,L,X) RHO= -SORT(NR+1.5)*RADQ(NA+1,LA,NB+1,1,L,X)- RHO GOTO 15 RHO= 0 \$ IF (NR.GT.0) RHO= SORT(FLOAT(NR))*ROFUNC(NR-1,1,X) RHO= -SORT(NR+1.5)*ROFUNC(NR,1,X)- RHO GOTO 15 IF (RHO) GOTO 3 RHO= SORT((LR+NR+ .5) *RADQ(NA+1,LA,NR+1,LR-1,L,X) 1 + SORT(NR+ 1.) *RADQ(NA+1,LA,NR+2,LR-1,L,X)) \$ GOTO 14 RHO= SORT(LR+NR+.5)*ROFUNC(NR,LR-1,X) 1 + SORT(NB+1.) *ROFUNC(NB+1,LR-1,X) CONTINUE LPA= LA+L \$ LMA= IABS(LA-L) IF (LPA+1 .LT. LB .OR. LMA+1 .GT. LB) GOTO 4 TLYDEL= -SORT(LA*(I3-1.))*CG000(LA,L,LR-1) RACAH IS W(LB,1,LA,L,LR-1,LAM) TLYDEL= TLYDEL*FWCOEF(I1,I2,I3,I4,I5,I6)*RMO IF (LPA-1 .LT. LB .OR. LMA-1 .GT. LR) GOTO 5 IF (RHO) GOTO 6 RMO= RMO- (I3+1)*X/(L+L+1.) 1 * (RADQ(NA+1,LA,NR+1,LR,L-1,X) + RADQ(NA+1,LA,NB+1,LR,L+1,X)) GOTO 15 RMO= RMO- (I3+1)*ROFUNC(NR,LR,X)/X CONTINUE RMO= RMO*SORT((LB+1.)*(I3+3)) *CG000(LA,L,LR+1) RACAH IS W(LB,1,LA,L,LR+1,LAM) TLYDEL= TLYDEL+ RMO*FWCOEF(I1,I2+2,I3+2,I1,I5,I6) TLYDEL= TLYDEL*FAS \$ RETURN END </pre>	TLYDEL 2 TLYDEL 3 TLYDEL 4 TLYDEL 5 TLYDEL 6 TLYDEL 7 TLYDEL 8 TLYDEL 9 TLYDEL 10 TLYDEL 11 TLYDEL 12 TLYDEL 13 TLYDEL 14 TLYDEL 15 TLYDEL 16 TLYDEL 17 TLYDEL 18 TLYDEL 19 TLYDEL 20 TLYDEL 21 TLYDEL 22 TLYDEL 23 TLYDEL 24 TLYDEL 25 TLYDEL 26 TLYDEL 27 TLYDEL 28 TLYDEL 29 TLYDEL 30 TLYDEL 31 TLYDEL 32 TLYDEL 33 TLYDEL 34 TLYDEL 35 TLYDEL 36 TLYDEL 37 TLYDEL 38 TLYDEL 39 TLYDEL 40 TLYDEL 41 TLYDEL 42 TLYDEL 43 TLYDEL 44 TLYDEL 45 TLYDEL 46 TLYDEL 47 TLYDEL 48 TLYDEL 49 TLYDEL 50 TLYDEL 51 TLYDEL 52
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FORMFAC

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SUBROUTINE FORMFAC(CHARGE,SPIN,ICD,J,L,Q,IROA,N,B)
  COMPUTE FORM FACTORS
  ROUTINE SIMILAR TO CURRENT, EXCEPT THAT THE DELTA FUNCTION IN
  X IS REPLACED BY THE SPHERICAL BESSEL FUNCTION OF ORDER J.
  FOR MAGNETIC FORM FACTOR, USE THE RELATION
  MAG*(P+ CURL(MU)) = -K*( -LONG*L/SQRT(J*J+1) + ELEC*MU )
  O = K*9 , X = R/R , K IS THE MOMENTUM-TRANSFER.

  COMMON/CFACOR/GLP,GLN,MUSP,MUSN,EFC(2),EFM(2),BOHR,TRI
  REAL MUSP,MUSN,MUS , LOGICAL SLNAB,SJL
  DIMENSION IQQ(1) , LOGICAL IPASS
  DATA IPASS,ESP/.F.,1.E-5/
  DATA PI/3.1415 92653/, I22,I8/17777777B,377B/
  SJL = J.NE.L
  CHARGE = SPIN= 0 & IF (IARS(J-L).GT. 1 .OR. J.LT. 1 ) RETURN
  Z1= J*(J+J+1) & Z1= 1./SQRT(Z1)
  Z2= (J+1)*(J+J+1) & Z2= 1./SQRT(Z2)
  FIRST RECOVER QUANTUM NUMBERS AND DENSITY MATRIX ELEMENT FROM IRO
  DO 11 I= 1, N & IRO= IQQ(I) & ISO= SHIFT(IRO,-59)
  ISO= ISO.A.1 & S= ISO & ISO= IRO.A.I22 & RO= (S+S- 1.)*ISO*ESP
  ISO= SHIFT(IRO,-59) & ISO= ISO.A.1
  ISO= 1, PROTON, = 0, NEUTRON
  GL= ISO*(GLP+EFC(1))+ (1- ISO)*(GLN+EFC(2))
  MUS= ISO*(MUSP+EFM(1))+ (1- ISO)*(MUSN+EFM(2))
  IRO= SHIFT(IRO,-22) & IJUMP= 1
  NLJ= IRO.A.I8 & JA= NLJ.A.1 & NLJ= NLJ/2 & LA= NLJ.A.15
  JA= LA+JA & NLJ= SHIFT(NLJ,-4) & NA= NLJ.A.7 & GOTO (3,31) IJUMP
  LB= LA & JA= JA & NR= NA & ISO= SHIFT(IRO,-3) & IJUMP=2 & GOTO 2
  HAVE Q.N. AND RO, NOW COMPUTE DENSITY. FIRST CHARGE
  SLNAB= NA.EQ.NR.A. (LA.EQ.LB)
  IF (SLNAB.A.SJL.A. (JA.EQ.JR) ) GOTO 11
  IFAS= LB+J-LA & IF (ICD.GT.-2) GOTO 4

  THIS SECTION COMPUTES COULOMB FORM FACTOR

  IFAS= JR+J+ IFAS/2 & IFAS= IFAS.A.1 & FAS= 1- IFAS- IFAS
  FAS= FAS*SQRT(JB+JB/PI)
  CHARGE= CHARGE- FAS*RO*GL*CLERS(JA,JR,J)
  *RADD(NA+1,LA,NR+1,LR,J,Q)
  GOTO 11
  I6= LA+LR & IF (MOD(I6+L,2).EQ.0) GOTO 11 & I5= JB+JB- 1
  IF (I6.LT.J .OR. IARS(JA-JR).GT.1) GOTO 11
  IF (SLNAB.A.SJL) GOTO 7
  IF (I6.LT.J .OR. IARS(LA-LR).GT.J) GOTO 7 & IF (ICD) 5,5,6
  I1= I5+ I6 & I5= I5+ J & I2= JA+LR+J
  IFA= I2+ ( IFAS+1-IARS(J-L) )/2 & IFA= IFA.A.1
  I3= LA+JA & I4= LR+JR & I6= I4+J & FAS= 1- IFA- IFA
  FAS= FAS*2.*SQRT( (LA+LA+1.)*JA*JB/(J+J+1.))
  RAGAH IS W(LA,LR,JA,JR,J,1/2)
  FAS= FAS*PO*FWCOEF(I1,I2,I3,I4,I5,I6)*GL & IF (J.NE.L) GOTO 10
  IF (LA+LR.GE.J+1.A. IARS(LA-LR).LE.J+1)
  1CHARGE= CHARGE- FAS* Z1*TJYLL(NA,LA,NR,LR,J,J+1,Q)
  IF (LA+LR.GE.J-1.A. IARS(LA-LR).LE.J-1)
  1CHARGE= CHARGE- FAS* Z2*TJYLL(NA,LA,NR,LR,J,J-1,Q)
  GOTO 7
  IFA= (J+1).A.1 & TDD= TLYDELO(NA,LA,NR,LR,J,L,Q)
  TDD= TDD+ (1-IFA-IFA)*SQRT( (LB+LB+1.)/(LA+LA+1.))
  *TLYDELO(NR,LR,NA,LA,J,L,Q)
  1CHARGE= CHARGE- 0.5*FAS*TDD/2
  IF (ICD) 11, 6, 6
  NOW COMPUTE SPIN
  IF (J.EQ.L) GOTO 8
  ELCTRIC MATRIX ELEMENT. L = J+1 OR J-1
  S= JA*JR*2./PI/(J+1.)/(J+J+1.) & S= (L- J)*SQRT(S)
  XAR= (LA-JA+ 0.5)*(JA+JA) - (LR-JR+ 0.5)*(JB+JB)
  S= S*XAR*RADD(NA+1,LA,NR+1,LR,J,Q) & GOTO 9
  MAGNETIC MATRIX ELEMENT. L = J
  S= JA*JR/J/PI/(J+1.) & S= SQRT(S)/(J+J+1.)
  XAP= (LA-JA+ 0.5)*(JA+JA) + (LR-JR+ 0.5)*(JB+JB)
  S= S*(XAP+J+1.)*RADD(NA+1,LA,NR+1,LR,J+1,Q)
  1 - (J+1.)/J*(XAR-J)*RADD(NA+1,LA,NR+1,LR,J-1,Q)
  IFA= J+JR+ ( IFAS+1-IARS(J-L) )/2 & IFA= IFA.A.1
  SPIN= SPIN- (1- IFA- IFA) *CLERS(JA,JR,J)*S*RO*Q*MUS
  CONTINUE & IF (ICD.EQ.-2) RETURN
  CHARGE= 2.*BOHR*TRI*CHARGE/9 & SPIN= BOHR*TRI*SPIN/9 & RETURN
END

```

RADQ

FUNCTION RADQ(NP,LP,NF,LF,L,Q)

PROGRAM CALCULATES THE RADIAL INTEGRAL
(NP,LP,LP) J_L(Q) I_{NF}(LF)
WHERE J_L(P) IS THE SPHERICAL BESSEL FUNCTION OF ORDER L
AND Q IS A DIMENSIONLESS QUANTITY.
NFACT(N) IS EQUAL TO FACTORIAL OF (N-1),
NDFACT(N) IS EQUAL TO DOUBLE FACTORIAL OF N,
GAMMA(N) IS EQUAL TO GAMMA(N-.5)

DIMENSION GAMMA(100),NFACT(50),NDFACT(100)

LOGICAL IPASS

DATA IPASS/.FALSE./

ENTRY RADQ

IF((-1)**(LP+LF+L))10,10,21

21 IF(NF.LE.0)GOTO4; IF(NF.LE.0)GOTO4; IF(IPASS)GOTO2; IPASS=.TRUE.

PI=3.14159265359; B=SQR(PI); X=-.5; GAMMA(1)=PI DO3 I=2,100;X=X+1.

3 GAMMA(I)=X*GAMMA(I-1); NFACT(1)=1; DO5 I=2,50

5 NFACT(I)=(I-1)*NFACT(I-1); NDFACT(1)=1; NDFACT(2)=2; DO7 I=3,100

7 NDFACT(I)=I*NDFACT(I-2)

2 RNF=NFACT(NF)

RNP=NFACT(NP)

RLF=NDFACT(2*LF+1)

RLP=NDFACT(2*LP+1)

RF=NDFACT(2*LF+2*NF-1)

RP=NDFACT(2*LP+2*NP-1)

IF(7.LT.1.E-06) Q=1.E-06

A=((2.**((LF-NF+3))/B)*(RF/RNF)*(1./RLF**2)

C=((2.**((LP-NP+3))/B)*(RP/RNP)*(1./RLP**2)

AC=SQR(A*C)

SUM1=0.0

DO 1 K=1,NF

KT=K-1

RK=NFACT(KT+1)

RNK=NFACT(NF-KT)

RFK=NDFACT(2*LF+2*KT+1)

S=((-1.**KT*2.**KT*RNF*RLF)/(RK*RNK*RFK)

SUM2=0.0

DO 5 KS=1,NP

KS=KP-1

RKP=NFACT(KS+1)

RNKP=NFACT(NP-KS)

RFKP=NDFACT(2*LP+2*KS+1)

M=2*(KT+KS)+LF+LP+2

XS=FLOAT(M-L)*0.5-1.

9 NNS=XS

I=FLOAT(L+M)*0.5+1.

XA=PI*Q**L*GAMMA(I)*EXP(-I**2/4.)/(2.**((L+2)*GAMMA(L+2)))

* FOLLOWING TO ST. NO.11 COMPUTE CONFL. HYPER-G IF1(-NNS,L+1.5,OSQ/4)

* OSQ=Q*Q/4. \$ SUMM= SUMN= 1. \$ NSTP= NNS \$ NSB= 1

IF (NNS) 12, 11, 13

12 NSTP= OSQ+ 5

13 XD= -NNS \$ XN= L+ 1.5

14 DO 8 NS= NSB, NSTP \$ SUMN= SUMN*XD*OSQ/XN/NS

SUMM= SUMM+ SUMN \$ XN= XD+ 1. \$ XN= XN+ 1.

CONTINUE \$ IF ((NNS.GT.0) .OR. (ABS(SUMN/SUMM).LT.1.E-6))GOTO11

8 NSB= NSTP+ 1 \$ NSTP= NSTP+ 5 \$ GOTO 14

11 SUM2=SUM2+XA*SUMM*(-1.**KS*2.**KS*RNP*RLP/(RKP*RNKP*RFKP)

6 CONTINUE

SUM1=SUM1+S*SUM2

1 CONTINUE

XLP=LP

XL=LF

XL=L

RAD1= RADQ= AC*SUM1

RETURN

10 PRINT20

20 FORMAT(10X*MESSAGE FROM RADQ----L+LP+LAMBDA IS ODD, RADQ FAILS*)

RAD1= RADQ= 0.

RETURN

4 PRINT100

100 FORMAT(10X*MESSAGE FROM RADQ ---*/

1 10X*ERROR IN THE PRINCIPLE QUANTUM NUMBER OF THE OSCILLA-

2 TOP WAVE FUNCTION. THE W.F. WITH ZERO NODE MUST HAVE N=1. *)

RETURN

END

RADQ 2
RADQ 3
RADQ 4
RADQ 5
RADQ 6
RADQ 7
RADQ 8
RADQ 9
RADQ 10
RADQ 11
RADQ 12
RADQ 13
RADQ 14
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RADQ 77
RADQ 78
RADQ 79

TJYLL

<pre> * * FUNCTION TJYLL (NA,LA,NA,LR,LAM,L,Q) * * REDUCED MATRIX ELEMENT OF T(LAM) (Y(L), L) * J(L,Q) * T(LAM) IS SPHERICAL TENSOR OF RANK LAM * Y(L) IS SPHERICAL HARMONICS OF ORDER L * L IS THE ORBITAL ANGULAR MOMENTUM OPERATOR * J(L,Q) IS SPHERICAL BESSEL FUNCTION OF ORDER L, ARGUMENT Q * Q = MOMENTUM-TRANSFER*LENGTH-PARAMETER, DIMENSIONLESS * ONLY ENTRY POINT TJYLT TEST ALL TRIANGULAR AND PARITY SELECTION * RULES. * * DATA PI/3,1415 92653/ * TJYLL=0 \$ GOTO 1 \$ ENTRY TJYLT \$ TJYLL= 0 * IF (IABS(J-L) .GT. 1) RETURN * IF (LA+LR.LT.LAM .OR. IABS(LA-LR).GT.LAM) RETURN * IF (LA+LR.LT.L .OR. IABS(LA-LR).GT.L) RETURN * IF (MOD(LA+LR+L,2).EQ.1) RETURN * FIRST COMPUTE ARGUMENTS FOR FAST RACAH * I2= LR+ LR+ L+ LAM \$ I3= LR+ LR+ 1 \$ I4= LA+ LR+ L * I5= LA+ LR+ LAM \$ I6= L+ LAM+ 1 * COMPUTE GEOMETRIC FACTORS * TJYLL= (LR+ LR+LR)*(L+ L+ 1)*(LAM+ LAM+ 1) * TJYLL= 0.5*SQRT(TJYLL/PI)*I3 * COMPUTE PHASE * TJYLL= TJYLL*(1- 2*MOD(LAM+ LR, 2)) * NOW COMPUTE THE REST. RACAH IS W(LR,1,LA,L,LR,LAM) * TJYLL= TJYLL*FWCOEF(I4+1,I2,I3,I4,I5,I6) * TJYLL= TJYLL*RADDQ(NA+1,LA,NA+1,LR,L,Q)*CG000(LA,L,LR) * RETURN * END </pre>	<pre> TJYLL 2 TJYLL 3 TJYLL 4 TJYLL 5 TJYLL 6 TJYLL 7 TJYLL 8 TJYLL 9 TJYLL 10 TJYLL 11 TJYLL 12 TJYLL 13 TJYLL 14 TJYLL 15 TJYLL 16 TJYLL 17 TJYLL 18 TJYLL 19 TJYLL 20 TJYLL 21 TJYLL 22 TJYLL 23 TJYLL 24 TJYLL 25 TJYLL 26 TJYLL 27 TJYLL 28 TJYLL 29 TJYLL 30 TJYLL 31 </pre>
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RDFUNC

<pre> C C FUNCTION RDFUNC(N,L,X) C NORMALIZED H.O. RADIAL-FUNCTION UPTO NOSC=20 (NOSC=2*N+L) C DIMENSION FL(11),GL(21) C LOGICAL IPASS; DATA IPASS/.F./ C DATA N1MAX,L1MAX/11,21/ C N1= N+1; L1= L+1; RDFUNC=1; IF(N.LT.0.OR.L.LT.0) RETURN C NL= N+L+1; IF(N1.GT.N1MAX.OR.NL.GT.L1MAX) GO TO 5 C IF(IPASS) GO TO 2; IPASS= .T.; FL(1)=0.; GL(1)= ALOG(0.5) C DO 1 M = 2, L1MAX; A= 4-1; IF(M.GT.N1MAX) GO TO 1 C FL(M)= FL(M-1)+ALOG(A) C GL(M)= GL(M-1)+ALOG(A+0.5) C A= 2./SQRT(3.141593) C XSO= X*X C EVALUATE LAGUERRE POLYNOMIAL POLAG(N,L+1/2,XSO) C POLAG= EXP(GL(NL)-GL(L1)-FL(N1)) \$ IF(N.EQ.0) GO TO 4; B= 1. C DO 3 M=1,N; B= -B*XSO C POLAG= POLAG+B*EXP(GL(NL)-GL(L1+M)-FL(N1-M)-FL(M+1)) C B= EXP(FL(N1)-GL(NL)) \$ B= SQRT(A*B) C RDFUNC= B*(X*L)*EXP(-0.5*XSO)*POLAG; RETURN C PRINT6,N,L; RETURN C FORMAT(1X,'XXXXX',N=*,I2,2X,'L=*,I2,2X,'N OR L TOO LARGE', 11* CANNOT CALCULATE RDFUNC.CHANGE DIMENSION AND DATA 'XXXXX') C END </pre>	<pre> RDFUNC 2 RDFUNC 3 RDFUNC 4 RDFUNC 5 RDFUNC 6 RDFUNC 7 RDFUNC 8 RDFUNC 9 RDFUNC 10 RDFUNC 11 RDFUNC 12 RDFUNC 13 RDFUNC 14 RDFUNC 15 RDFUNC 16 RDFUNC 17 RDFUNC 18 RDFUNC 19 RDFUNC 20 RDFUNC 21 RDFUNC 22 RDFUNC 23 RDFUNC 24 </pre>
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WCOEF

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FUNCTION WCOEF(A,B,C,D,E,F)
* THIS PROGRAM, WRITTEN BY R.V. CUSSON AND J.H. SCHMIDT IN SEPT. 1971,
* USES REGGE'S FORMULA (NUOVO CIMENTO, VOL11(1959), P116) TO EVALUATE THE
* RACAH COEFFICIENT  $W(A,B,C,D,E,F) = (-1)^{A+B+C+D} \cdot SIXJ(A,B,E/D,C,F)$  .
*
* SINCE MANY PROGRAMS WHICH USE RACAH COEFFICIENTS ALSO USE CLEBSCH -
* GORDANS, COMMON BLOCK CGRACAH HAS BEEN SET UP SO TABLES OF LOGS AND
* SQUARE ROOTS DO NOT HAVE TO BE CALCULATED EACH TIME THE FUNCTIONS
* ARE CALLED. NOTE THAT ALOGF(0) IS SOMETIMES REQUIRED IN THE
* CALCULATIONS AND SO ALOGF0=ALOGF(0)=0.
*
* COMMON BLOCK RACASYS IS INCLUDED TO ENSURE THAT ALL ITS ELEMENTS ARE
* STORED SEQUENTIALLY IN CORE. THIS IS IMPORTANT FOR EQUIVALENCING.
*
* COMMON BLOCK XCRACAH IS USED IN CALCULATING X-COEFFICIENTS.

COMMON/CGRACAH/INITIAL,ALOGF0,ALOGF(80),SQRF(80)
COMMON/RACASYS/X1,X2,X3,Y1,Y2,Y3,Y4,IX1,IX2,IX3,IY1,IY2,IY3,IY4
COMMON/XCRACAH/JOUT,JXSWTCH
DIMENSION RV(7),IV(7)
EQUIVALENCE (RV,X1),(IV,IX1,VA),(IX2,VB),(IY1,VC),(IY2,VD),
1 (IY3,VE),(IY4,VF)
DATA INITIAL/0/,ALOGF0/0./,JXSWTCH/0/
400 FORMAT(' ILLEGAL PARAMETERS IN WCOEF(*,6F5.1,*)')
* ENSURES THAT ALOGF AND SQRF ARE SET UP ONLY ONCE IN EACH PROGRAM
IF (INITIAL .NE. 0) GO TO 20
ALOGF(1)=0.
SQRF(1)=1. $ SM=1.
DO 25 J=2,80
SM=SM*1.
SQRF(J)=SQRT(SM)
25 ALOGF(J)=ALOGF(J-1)+ALOG(SM)
INITIAL = 1
20 SIXJ=1. $ IFAS=0 $ WCOEF=0.
* SET UP THE REGGE SYMBOL
Y4= A+B $ Y3=C+D $ Y2=E+F
X1=Y4+Y3 $ X2=A+D+Y2 $ X3=B+C+Y2
Y1=Y4+E $ Y2=Y3+E $ Y3=A+C+F $ Y4=B+D+F
* CHECK THAT THE REGGE SYMBOL CONTAINS ONLY NON-NEGATIVE INTEGERS
DO 1 K=1,7
J=A35(RV(K))
IF (FLOAT(J) .NE. RV(K)) GO TO 998
1 IV(K)= J
* CHECK FOR TRIANGLE AND IN RANGE CONDITION .
GO TO 201
ENTRY F4COEF
* CALLING SEQUENCE IS F=F4WCOEF(IX1,IX2,IY1,IY2,IY3,IY4) WHERE
* IX1 = A+B+C+D IY2=C+D+E
* IX2 = A+D+E+F IY3=A+C+F
* IY1 = A+B+E IY4=B+D+F
* IF PARAMETERS ARE LEGAL THE VALUE OF WCOEF(A,B,C,D,E,F) IS RETURNED
* WCOEF MUST BE CALLED FIRST TO INITIALISE /CGRACAH/
* NOTE - IF F4WCOEF IS CALLED, NO ARGUMENT CHECKING IS DONE
VA=A $ VB=B $ VC=C $ VD=D $ VE=E $ VF=F
IX3= IY1+IY2+IY3+IY4-IX1-IX2
201 IY0= MAX0(IY1,IY2,IY3,IY4)
IX0= MIN0(IX1,IX2,IX3)

```

WCOEF (CONTINUED)

IF (IY0 .GT. 78) GO TO 938	WCOEF	59
X=IX0-IY0+1	WCOEF	60
IF (X .LT. 1.) GO TO 998	WCOEF	61
IFAS=IX1	WCOEF	62
IF (IX0 .LT. 0) GO TO 939	WCOEF	63
* THE EXPRESSION (IY0 .AND. 1) SIMPLY CHECKS IF IY0 IS ODD OR EVEN	WCOEF	64
K=IY0 .AND. 1	WCOEF	65
* NOTE DELC IS EVALUATED AS FOLLOWS	WCOEF	66
DELC=(1)-(2)+((3)-(4))	WCOEF	67
* WHERE (1)=SUM OF SMALLER LOGS	WCOEF	68
(2)=SUM OF SMALLER LOGS	WCOEF	69
(3)=MOST LIKELY TO BE THE LARGEST LOG	WCOEF	70
(4)=SUM OF LARGE LOGS	WCOEF	71
DELC=((ALOGF(IX1-IY1)+ALOGF(IX1-IY2)+ALOGF(IX1-IY3)+ALOGF(IX1-IY4)	WCOEF	72
1 +ALOGF(IX2-IY1)+ALOGF(IX2-IY2)+ALOGF(IX2-IY3)+ALOGF(IX2-IY4)	WCOEF	73
2 +ALOGF(IX3-IY1)+ALOGF(IX3-IY2)+ALOGF(IX3-IY3)+ALOGF(IX3-IY4))*5)	WCOEF	74
3 - (ALOGF(IY0-IY1)+ALOGF(IY0-IY2)+ALOGF(IY0-IY3)+ALOGF(IY0-IY4)	WCOEF	75
4 +ALOGF(IX1-IY0)+ALOGF(IX2-IY0)+ALOGF(IX3-IY0))+(ALOGF(IY0+1)	WCOEF	76
5 -.5*(ALOGF(IY1+1)+ALOGF(IY2+1)+ALOGF(IY3+1)+ALOGF(IY4+1)))	WCOEF	77
DELC=(1-K-K)*EXP(DELC)	WCOEF	78
* INITIALISE THE SUM	WCOEF	79
SM=1.	WCOEF	80
IF (X .LE. 1.) GO TO 13	WCOEF	81
Z0=IX0+3	WCOEF	82
K=IX0+1 \$ X1=IX1-K \$ X2=IX2-K \$ X3=IX3-K	WCOEF	83
K=<+1 \$ Y1=K-IY1 \$ Y2=K-IY2 \$ Y3=K-IY3 \$ Y4=K-IY4	WCOEF	84
* THE SUM IS DONE RECURSIVELY WITHOUT ALOGF	WCOEF	85
* NOTE BECAUSE OF THE DEFINITION OF X, THE SUM FROM 2 TO X IS THE	WCOEF	86
* SAME AS THE SUM FROM IY0 TO IX0	WCOEF	87
XX=2.	WCOEF	88
15 SM= 1.-SM*(Z0-XX)*(X1+XX)*(X2+XX)*(X3+XX)/	WCOEF	89
1 ((Y1-XX)*(Y2-XX)*(Y3-XX)*(Y4-XX))	WCOEF	90
XX=XX+1.	WCOEF	91
IF (XX .LE. X) GO TO 15	WCOEF	92
* SIXJ IS THE USUAL 6-J SYMBOL OF WIGNER .	WCOEF	93
13 SIXJ= DELC*SM	WCOEF	94
999 IFAS=IFAS .AND. 1	WCOEF	95
WCOEF=SIXJ*(1-IFAS-IFAS)	WCOEF	96
997 RETJRN	WCOEF	97
* JXSWTCH DETERMINES WHETHER OR NOT WCOEF WAS CALLED BY XCOEF. IF IT	WCOEF	98
* WAS, NO ERROR MESSAGE IS PRINTED BY WCOEF	WCOEF	99
998 IF (JXSWTCH .EQ. 0) GO TO 996	WCOEF	100
JJUT=1 \$ JXSWTCH=0	WCOEF	101
GO TO 937	WCOEF	102
996 PRINT 400,A,B,C,D,E,F	WCOEF	103
GO TO 937	WCOEF	104
END	WCOEF	105

LINPLOT

```

SUBROUTINE LINPLOT(X,Y,Z,NX,N,NAME,SNAME)
DIMENSION X(50),Y(50),Z(50),NAME(4),SNAME(3)
DATA FORM1/6H(F7.4)/,FORM2/10H(F10.2A10)/,FORM3/6H(F5.2)/
DATA FORM4/6H(4A10)/
FIND MAX ABSOLUTE VALUE FOR ARRAY Y AND Z
AM=0.0
DO 1 I=1,N
1 AM=AMAX1(AM,ABS(Y(I)),ABS(Z(I)))
IF(AM)6,6,7
6 PRINT *,NAME
GO TO 3
8 FORMAT(*C*,5A10,* ARRAYS ARE ZERO*)
SELECT VEPTICAL SCALE FACTORAPPLICABLE WHEN AM IS LESS THAN 1
7 AM=2*AM $ NP=0
2 IF(AM.GT.1) GO TO 3 $ NP = NP +1 $ AM = AM*10 $ GO TO 2
3 AM=AIN*(AM+1.0) $ AM=AM*10.0*(-NP)
UPPER AND LOWER LIMITS FOR AXIS
YMAX= AM/2.0 $ YMIN = - YMAX $ XMIN = 0.0 $ XMAX = INT(X(N) + 1.0)
PLOTING THE FRAME
TICX = 10* NX $ TICZ = AM*0.1 $ XL = XMAX/(10*NX)
CALL PLOT(1,XL,10.0,XMIN,XMAX,YMIN,YMAX,TICX,TICZ)
PLOTING THE ARRAYS BOTH ON THE SAME FRAME
CALL PLOT(2,1,3,X,Y,N) $ CALL PLOT(2,1,0,X,Z,-N)
LABELS ON Y AXIS
DO 4 I=1,11 $ XLOC= XMIN -0.8*TICX
APP = YLOC = YMIN + (I-1)*TICZ
4 CALL PLOT(3,FORM1,1.0,0,XLOC,YLOC,ARR,1)
IDENTIFICATION ON UPPER AXIS
XLOC=XMIN $ YLOC= YMAX + 0.2 + TICZ
CALL PLOT(3,FORM4,1.0,0,XLOC,YLOC,NAME,4)
XLOC= XMIN+3.5*TICX $ CALL PLOT(3,FORM2,1.0,0,XLOC,YLOC,SNAME,3)
LABELS ON X AXIS
YLOC= YMIN - 0.2*TICZ $ IXL= INT(XL)+1
DO 5 I= 1,IXL $ XLOC = XMIN + (I-1)*TICX - 0.3*TICX
APP = X*IN + (I-1)*TICX
5 CALL PLOT(3,FORM3,1.0,0,XLOC,YLOC,APP,1)
9 RETURN
END

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LINPLOT 2
LINPLOT 3
LINPLOT 4
LINPLOT 5
LINPLOT 6
LINPLOT 7
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LINPLOT 9
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LINPLOT 39

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SEMILOG

```

SUBROUTINE SEMILOG(X,Y,Z,NX,N,NAME,SNAME,EXL,NC)
DIMENSION X(50),Y(50),Z(50),NAME(4),SNAME(3)
DATA FORM/10H(F10.2A10)/
DATA FORM4/6H(4A10)/
DATA FORMX/8H(MEV) C/,FORMY/10HFORMFACTOR /,FORMSL/6H(F5.1)/
SET LIMITS FOR PLOT YMAX = 10EXP EXL CYCLES = NC
XMIN=0,XMAX=X(N),TIC=5.0,XSL=XMAX/TIC,SYL=10
CALL S4LOGY(XL,YL,XMIN,XMAX,TIC,EXL,NC,FORMX,8,FORMY,12,FORMSL,5)
F(CHARGE) SYMBOL 7 FOR NEGATIVE VALUES AND 8 FOR POSITIVE
DO 1 I=1,N
IF(Y(I)-0.)2,3,4
2 SYM1=7SY(I)=ALOG10(ABS(Y(I)))GO TO 5
3 SYM1=8 $ Y(I)=ALOG10(Y(I)) $ GO TO 5
4 Y(I) = -10 $ SYM1= 8
5 CALL PLOT(2,1,SYM1,X(I),Y(I),1)
1 CONTINUE
F(SPIN) SYMBOL 1 FOR NEG AND 2 FOR POSITIVE
DO 11 I=1,N
IF(Z(I)-0.) 6,7,8
6 SYM2=1 $ Z(I) = ALOG10(ABS(Z(I))) $ GO TO 9
7 SYM2=2 $ Z(I) = -10 $ GO TO 9
8 SYM2=2 $ Z(I) = ALOG10(Z(I))
9 CALL PLOT(2,1,SYM2,X(I),Z(I),1)
11 CONTINUE
SYM3=63 DO 12 I=1,N
12 YZ(I)=ALOG10(YZ(I))
CALL PLOT(2,1,SYM3,X,YZ,N)
WRITE LABEL
YLOC=EXL+0.05 $ XLOC=0.0 $ CALL PLOT(3,FORM4,1.0,XLOC,YLOC,NAME,4)
XLOC=4.0 $ CALL PLOT(3,FORM,1.0,0,XLOC,YLOC,SNAME,3)
RETURN
END

```

```

SEMILOG 2
SEMILOG 3
SEMILOG 4
SEMILOG 5
SEMILOG 6
SEMILOG 7
SEMILOG 8
SEMILOG 9
SEMILOG 10
SEMILOG 11
SEMILOG 12
SEMILOG 13
SEMILOG 14
SEMILOG 15
SEMILOG 16
SEMILOG 17
SEMILOG 18
SEMILOG 19
SEMILOG 20
SEMILOG 21
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SEMILOG 23
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SEMILOG 25
SEMILOG 26
SEMILOG 27
SEMILOG 28
SEMILOG 29
SEMILOG 30
SEMILOG 31
SEMILOG 32
SEMILOG 33

```

C.2 SAMPLE OUTPUT OF MICRODENS

Case No. 1: $|f\rangle = |(\pi 1 S_{1/2})(\pi 0 d_{5/2})^{-1}; 2^+ \rangle$; Coulomb transition

1 2.000 .100 48 20.0300.0 -0.0 -0.0 20.0 -0.0 -0.0 -0.0 -0.0 2 2-0-0

COULOMB

TEST CASE NO.1, 2+, P(005/2)*P(1S1/2), B=2.0(F)

33 51 1.00000
60000000410120303240

TEST CASE NO.1, 2+, P(005/2)*P(1S1/2), B=2.0(F)
COULOMB, J=2 L=2, X=R/B, DENSITY TIMES B-CUBED IS PRINTED

X	CHARGE	SPIN	TOTAL
.10	.43376E-02	0.	.43376E-02
.20	.16499E-01	0.	.16499E-01
.30	.34102E-01	0.	.34102E-01
.40	.53723E-01	0.	.53723E-01
.50	.71563E-01	0.	.71563E-01
.60	.84119E-01	0.	.84119E-01
.70	.89150E-01	0.	.89150E-01
.80	.85337E-01	0.	.85337E-01
.90	.73108E-01	0.	.73108E-01
1.00	.54086E-01	0.	.54086E-01
1.10	.30762E-01	0.	.30762E-01
1.20	.60192E-02	0.	.60192E-02
1.30	.17423E-01	0.	.17423E-01
1.40	.33229E-01	0.	.33229E-01
1.50	.56160E-01	0.	.56160E-01
1.60	.65646E-01	0.	.65646E-01
1.70	.64922E-01	0.	.64922E-01
1.80	.60589E-01	0.	.60589E-01
1.90	.53867E-01	0.	.53867E-01
2.00	.45865E-01	0.	.45865E-01
2.10	.37583E-01	0.	.37583E-01
2.20	.29723E-01	0.	.29723E-01
2.30	.22852E-01	0.	.22852E-01
2.40	.16855E-01	0.	.16855E-01
2.50	.12168E-01	0.	.12168E-01
2.60	.84753E-02	0.	.84753E-02
2.70	.57537E-02	0.	.57537E-02
2.80	.38042E-02	0.	.38042E-02
2.90	.24449E-02	0.	.24449E-02

Case No. 2: $|f\rangle = |(\pi S_{1/2})(\pi O d_{5/2})^{-1}; 2^+\rangle$; Electric transition

1 2.000 .100 48 20.0300.0 -0.0 -0.0 20.0 -0.0 -0.0 -0.0 2 2-0-0

ELECTRIC

TEST CASE NO.2, 2+, P(0D5/2) + P(1S1/2), B=2.0(F)

33 51 1.00000
60000000410120303240

TEST CASE NO.2, 2+, P(0D5/2) + P(1S1/2), B=2.0(F)
ELECTRIC J=2 L=1, X=R/B, DENSITY TIMES B-CUBED IS PRINTED

X	CHARGE	SPIN	TOTAL
100	72207E-02	19957E-01	12737E-01
120	13384E-01	3718E-01	12335E-01
140	19335E-01	4945E-01	30060E-01
160	22565E-01	5423E-01	31899E-01
180	22524E-01	5474E-01	20417E-01
200	22981E-01	4398E-01	10284E-01
220	16163E-01	3205E-01	19205E-01
240	12372E-01	7767E-02	12117E-01
260	8891E-01	5019E-02	12117E-01
280	5157E-01	4433E-02	12117E-01
300	11573E-01	2437E-02	12117E-01
320	11793E-01	2337E-02	12117E-01
340	11573E-01	2337E-02	12117E-01
360	11793E-01	2337E-02	12117E-01
380	11573E-01	2337E-02	12117E-01
400	11793E-01	2337E-02	12117E-01
420	11573E-01	2337E-02	12117E-01
440	11793E-01	2337E-02	12117E-01
460	11573E-01	2337E-02	12117E-01
480	11793E-01	2337E-02	12117E-01
500	11573E-01	2337E-02	12117E-01
520	11793E-01	2337E-02	12117E-01
540	11573E-01	2337E-02	12117E-01
560	11793E-01	2337E-02	12117E-01
580	11573E-01	2337E-02	12117E-01
600	11793E-01	2337E-02	12117E-01
620	11573E-01	2337E-02	12117E-01
640	11793E-01	2337E-02	12117E-01
660	11573E-01	2337E-02	12117E-01
680	11793E-01	2337E-02	12117E-01
700	11573E-01	2337E-02	12117E-01
720	11793E-01	2337E-02	12117E-01
740	11573E-01	2337E-02	12117E-01
760	11793E-01	2337E-02	12117E-01
780	11573E-01	2337E-02	12117E-01
800	11793E-01	2337E-02	12117E-01
820	11573E-01	2337E-02	12117E-01
840	11793E-01	2337E-02	12117E-01
860	11573E-01	2337E-02	12117E-01
880	11793E-01	2337E-02	12117E-01
900	11573E-01	2337E-02	12117E-01
920	11793E-01	2337E-02	12117E-01
940	11573E-01	2337E-02	12117E-01
960	11793E-01	2337E-02	12117E-01
980	11573E-01	2337E-02	12117E-01
1000	11793E-01	2337E-02	12117E-01

[illegible]

Case No. 3: $|f\rangle = |(OP_{1/2})(OP_{3/2})^{-1}; 1^+, T=1\rangle$; Magnetic transition

2 2.000 .100 48 20.0300.0 -0.0 -0.0 20.0 -0.0 -0.0 -0.0 -0.0 1 1-0-0

MAGNETIC

TEST CASE NO.3, 1+,T=1, (OP3/2) $\times$ (OP1/2), B=2.0(F)

2 31 .70710 2 30 -.70710
6000000020060212066 00000000020060212066

TEST CASE NO.3, 1+,T=1, (OP3/2) $\times$ (OP1/2), B=2.0(F)
MAGNETIC J=1 L=1, X=R/B, DENSITY TIMES B-CUBED IS PRINTED

X	CHARGE	SPIN	TOTAL
.10	25472E-02	5663E-02	84135E-02
.20	494339E-02	6889E-01	15633E-01
.30	76542E-02	593E-01	20648E-01
.40	87697E-02	414E-01	22791E-01
.50	10019E-01	772E-01	21785E-01
.60	10770E-01	867E-01	11552E-01
.70	11085E-01	57E-02	37077E-01
.80	10301E-01	414E-01	4778E-01
.90	10455E-01	24E-01	12723E-01
1.0	94394E-02	162E-01	25006E-01
1.1	73149E-02	311E-01	28346E-01
1.2	61716E-02	518E-01	29742E-01
1.3	50676E-02	816E-01	29389E-01
1.4	40623E-02	15E-01	27627E-01
1.5	31823E-02	307E-01	24874E-01
1.6	24330E-02	57E-01	21507E-01
1.7	18124E-02	93E-01	18001E-01
1.8	13247E-02	150E-01	14561E-01
1.9	94567E-02	209E-01	11441E-01
2.0	64756E-02	37E-01	86818E-01
2.1	49458E-02	57E-01	64157E-01
2.2	39458E-02	7E-01	46157E-01
2.3	29417E-02	107E-01	32049E-01
2.4	19415E-02	155E-01	22609E-01
2.5	12415E-02	21E-01	14509E-01
2.6	77545E-02	36E-01	9609E-01
2.7	47399E-02	55E-01	6071E-01
2.8	28361E-02	83E-01	371E-01
2.9	1652E-02	155E-01	109E-01
3.0	952E-02	35E-01	33E-01
3.1	52E-02	54E-01	10E-01
3.2	45E-02	5E-01	3E-01
3.3	45E-02	5E-01	3E-01
3.4	45E-02	5E-01	3E-01
3.5	45E-02	5E-01	3E-01
3.6	45E-02	5E-01	3E-01
3.7	45E-02	5E-01	3E-01
3.8	45E-02	5E-01	3E-01
3.9	45E-02	5E-01	3E-01
4.0	45E-02	5E-01	3E-01
4.1	45E-02	5E-01	3E-01
4.2	45E-02	5E-01	3E-01
4.3	45E-02	5E-01	3E-01
4.4	45E-02	5E-01	3E-01
4.5	45E-02	5E-01	3E-01

C.3 SPECIAL RELATIONS FOR CHECKING THE RESULTS

There are some special properties of the densities and form factors which are useful to know.

Using the "form factors" given in equations (C.1,2)

$$\langle \vec{L}_\lambda \cdot \vec{J} \rangle = \sqrt{4\pi} (\sqrt{\pi} f_{\lambda, \lambda-1}(q) - \sqrt{\lambda+1} f_{\lambda, \lambda+1}(q)). \quad (C.4)$$

Here $\langle \rangle$ means the reduced matrix element between the initial and final states.

For the magnetization current

$$\langle \vec{L}_\lambda \cdot (\vec{\nabla} \times \vec{\mu}) \rangle = \langle \phi_\lambda \vec{\nabla} \cdot (\vec{\nabla} \times \vec{\mu}) \rangle \equiv 0.$$

Therefore we have the first general relation

$$(i) \quad \sqrt{\lambda} f_{\lambda, \lambda-1}^m - \sqrt{\lambda+1} f_{\lambda, \lambda+1}^m \equiv 0; \quad \text{general.} \quad (C.5)$$

If the convection current is conserved we have

$$\langle \vec{L}_\lambda \cdot \vec{J}^c \rangle = \frac{k}{q} \langle \phi_\lambda \rho \rangle,$$

where k is the wave-number of the energy transfer. This leads to the second relation

$$(ii) \quad \sqrt{\lambda} f_{\lambda, \lambda-1}^c - \sqrt{\lambda+1} f_{\lambda, \lambda+1}^c = \frac{k}{q} \hat{\lambda} f_\lambda; \quad \text{if } \vec{J}^c \text{ conserved.} \quad (C.6)$$

This relation must be used with discretion since in a microscopic, many-body calculation $\vec{J}^c$ is almost never conserved. However in the single-particle model, where the one-body potential is momentum-independent, (ii) is satisfied. In the special case of harmonic oscillator, the energy difference in the transition $|b\rangle \rightarrow |a\rangle$ is

$$\Delta N \hbar \omega = (2n_a + l_a - 2n_b - l_b) \hbar \omega$$

and

$$k = \Delta N \cdot \hbar c / (M d^2)$$

where M is the nucleon mass and d is the oscillator length parameter. For such cases

$$(iii) \quad \sqrt{\lambda} f_{\lambda, \lambda-1}^c - \sqrt{\lambda+1} f_{\lambda, \lambda+1}^c = \frac{\Delta N \cdot \hbar c}{q M d^2} \hat{\lambda} f_{\lambda};$$

$$\underline{\text{harm. osc. s.p. transition}} \quad (C.7)$$

A special case for (iii) is when $|a\rangle$ and $|b\rangle$ are orbitals in the same shell, in which case $\Delta N = 0$; therefore the LHS of (C.7) must vanish.

We recall (see section 2.8) that $\vec{J}^c$ vanishes in an elastic transition. This can be extended to transitions between two s.p. states split by a simple spin-orbit interaction (constant $\times \vec{L} \cdot \vec{\sigma}$), since this

interaction does not affect the spatial wavefunctions. The harmonic oscillator model provides such an example. Therefore we have

$$(iv) \quad c_{\lambda, \lambda \pm 1}^C = f_{\lambda, \lambda \pm 1}^C \equiv 0; \quad \text{s.p. transition, } n_a = n_b, l_a = l_b. \quad (C.8)$$

Equations (C.4,6,7) can be used to check the results of MICRØDENS. Equation (C.8) has been incorporated into the coding of the program.

As was mentioned earlier, the form factors in MICRØDENS are calculated analytically rather than by Bessel transforming the computed densities. In particular the magnetic form factor is calculated using (the first line of) (A.27c') which involves formulas significantly different from those used in computing the magnetic density. Therefore an additional consistency check can be made by performing the integrations in (C.1,2).



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